Measurements of the Hubble Constant: Tensions in Perspective

Freedman (2021)



Wei Zhu (祝伟)

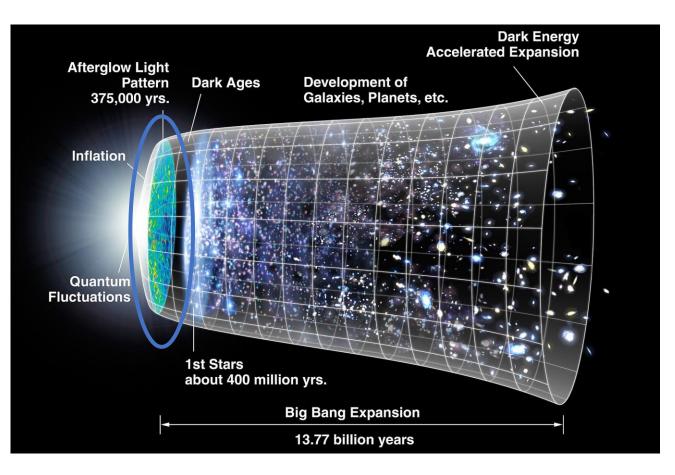
Astro Student Seminar, 2021-09-24

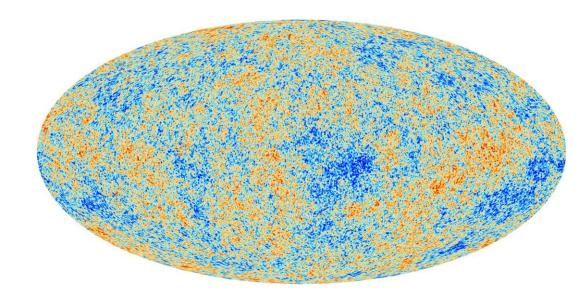


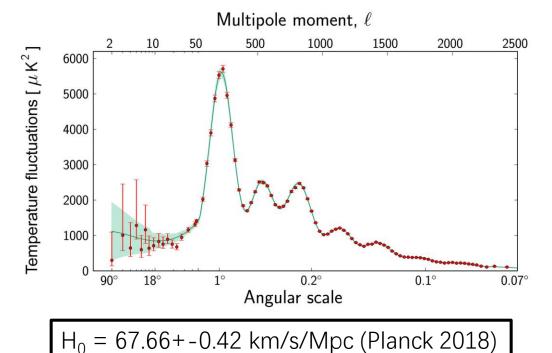
Outline

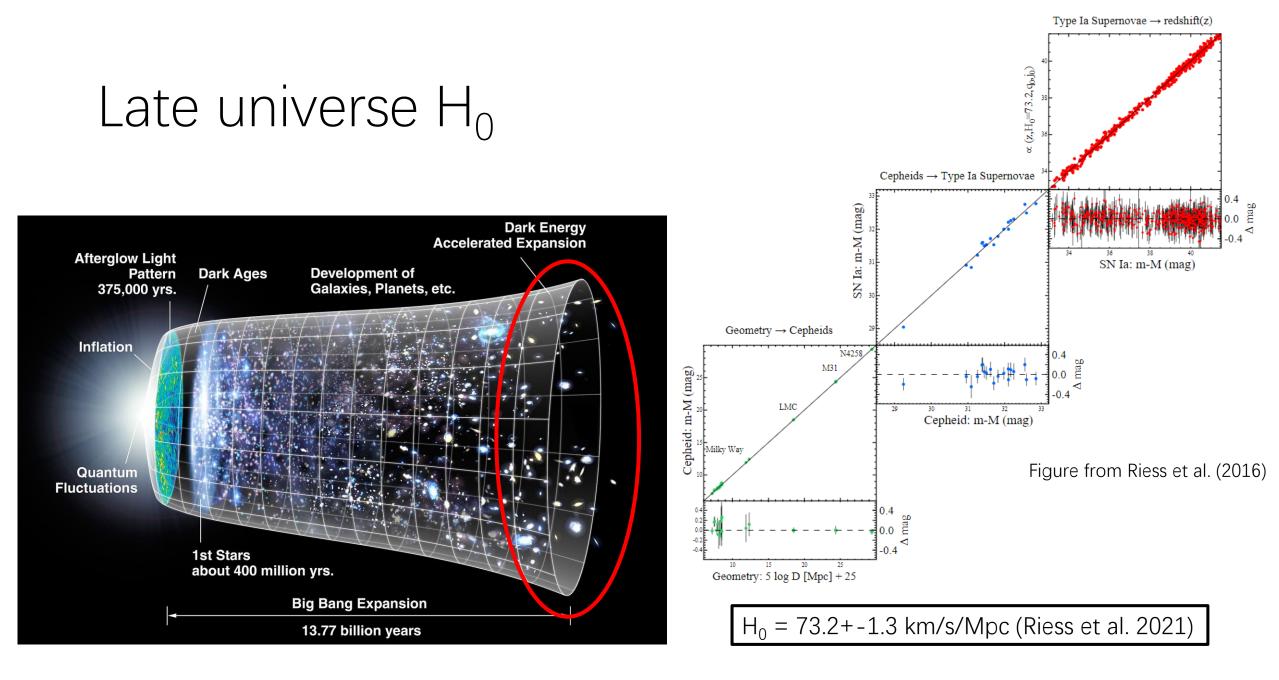
- Hubble tension & TRGB method
- Freedman (2021)
- My comments
- Summary

Early universe H₀

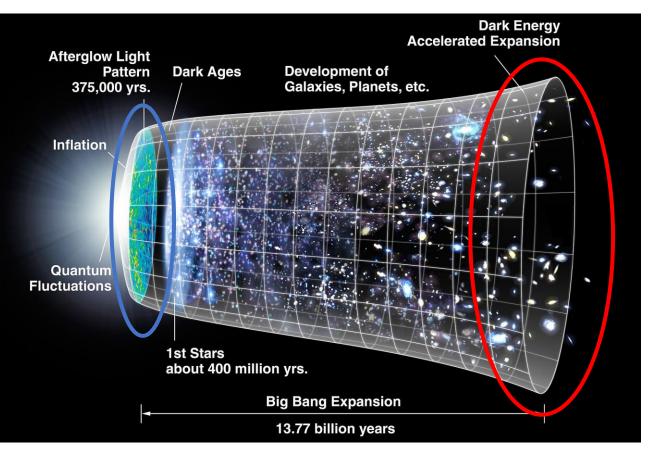


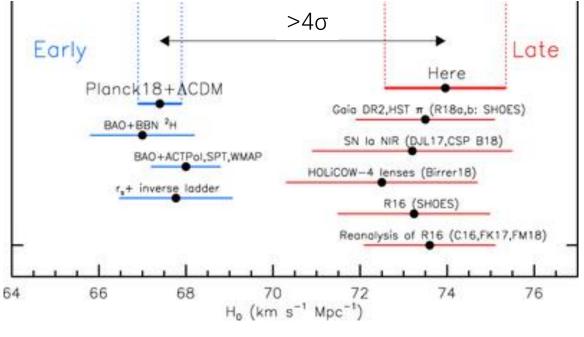






Hubble tension



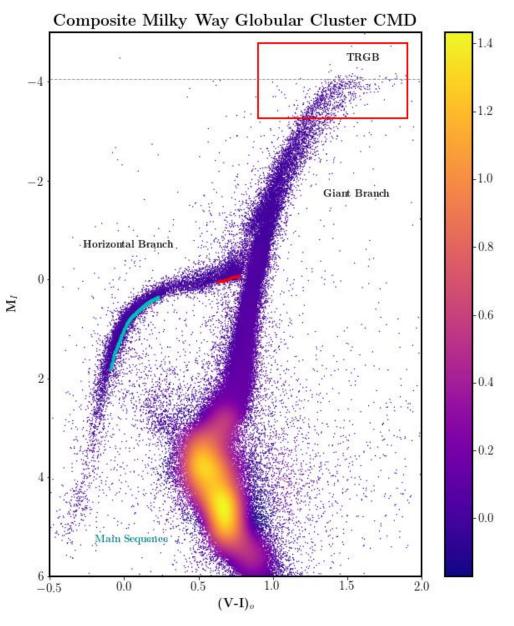


Riess et al. (2016, 2019, 2021)

New physics? Or Systematics?

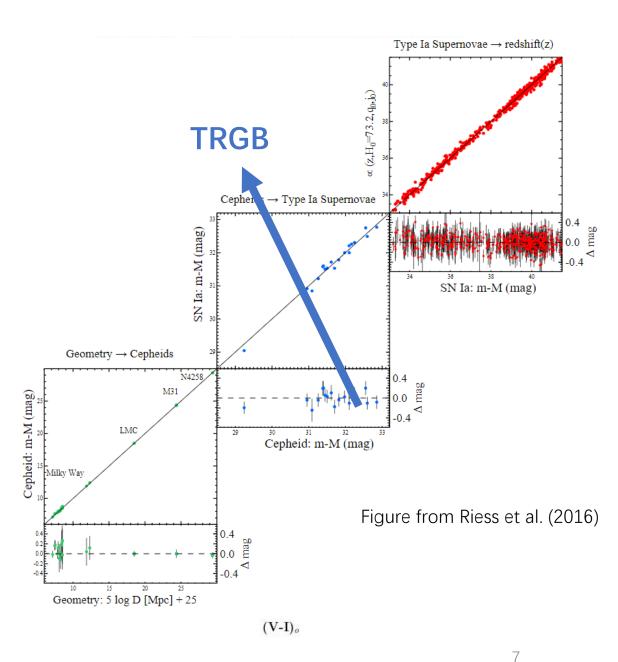
TRGB method

 Helium flash in low-mass stars
 → Tip of the Red Giant Branch (TRGB) as a standard candle (especially in I band).



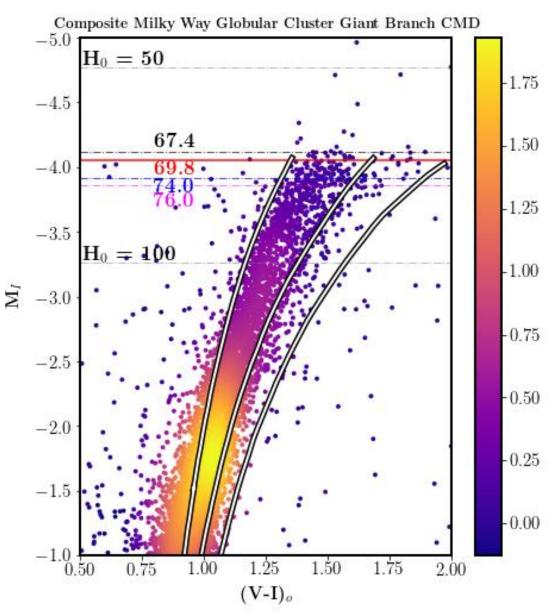
TRGB method

- Helium flash in low-mass stars
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- Replacing Cepheids with TRGB in the distance ladder \rightarrow H₀



TRGB method

- Helium flash in low-mass stars → Tip of the Red Giant Branch (TRGB) as a standard candle (especially in I band).
- Replacing Cepheids with TRGB in the distance ladder \rightarrow H₀
- Different calibrations of the TRGB method \rightarrow different TRGB zero point M_I \rightarrow different H₀

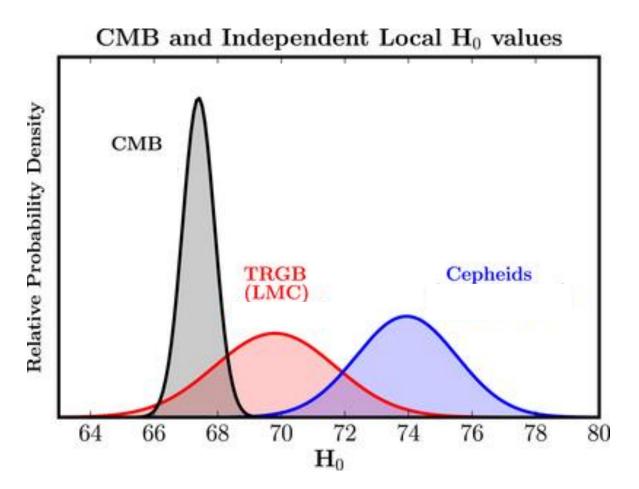


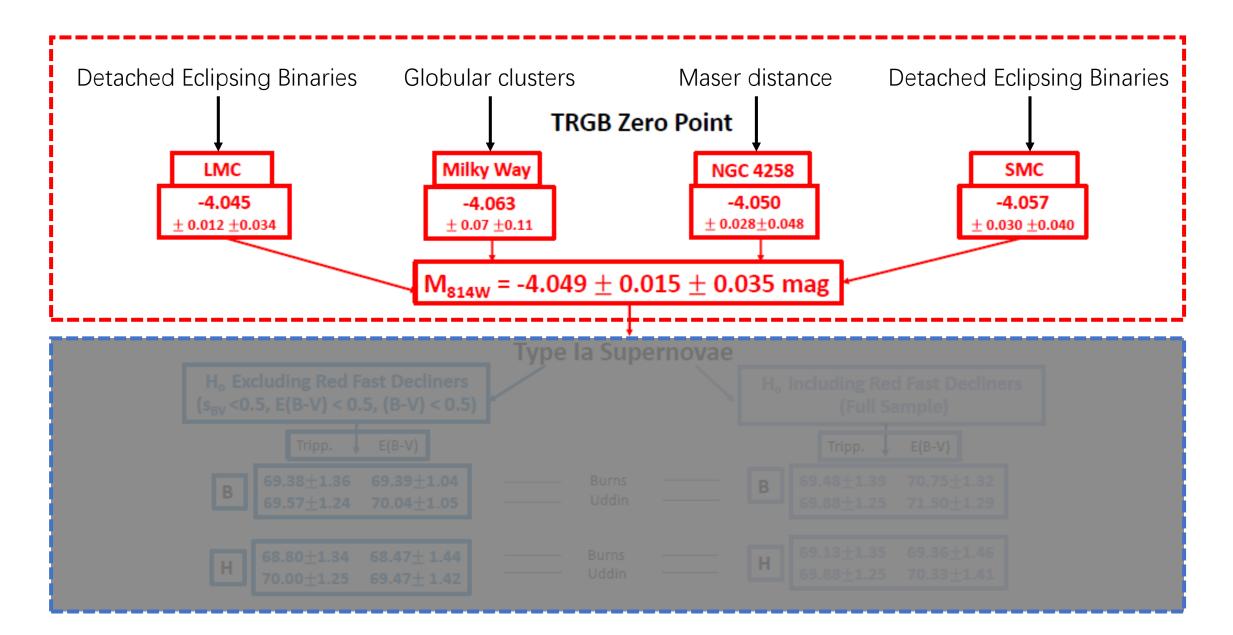
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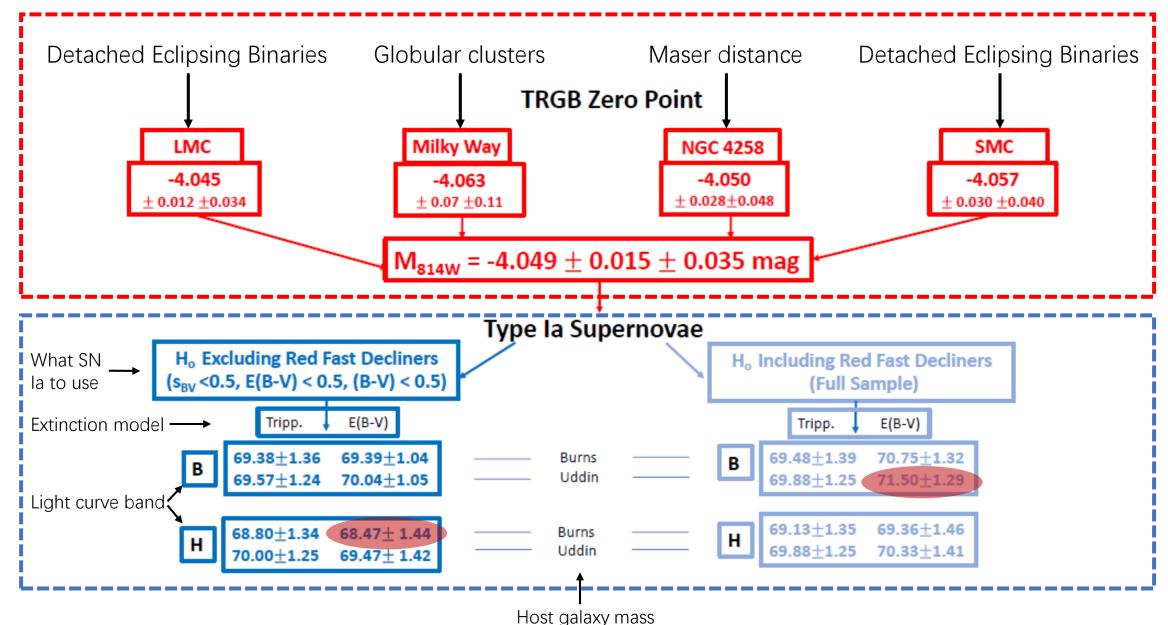
Take-home message

Freedman (2021): Local measurement of H_0 from TRGB is more consistant with the early universe result.



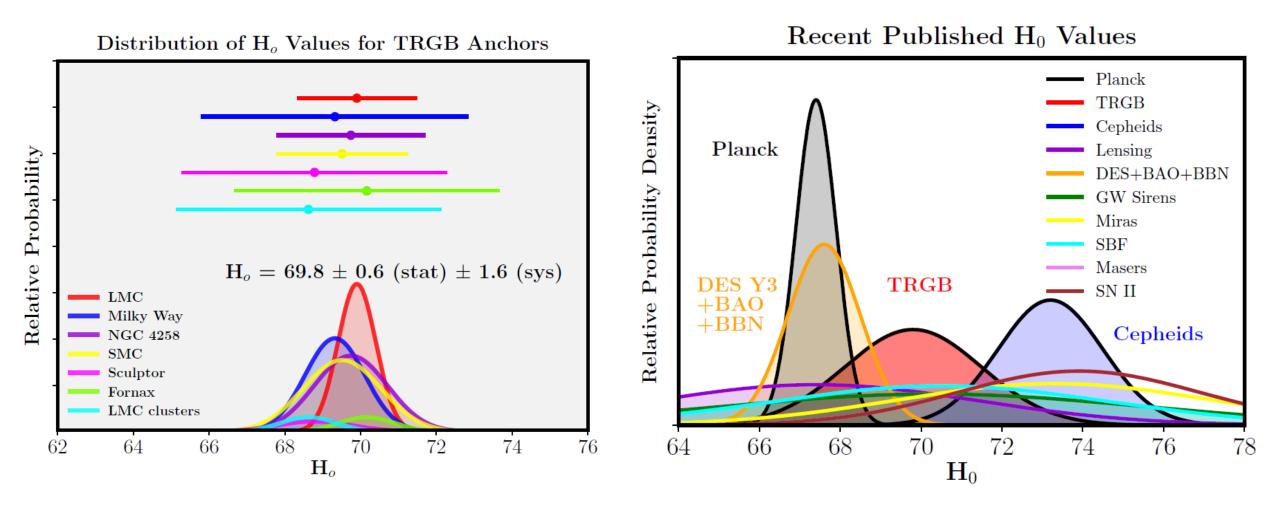


Freedman (2021)



Freedman (2021)

H₀ from TRGB method



Freedman (2021)

3. My comments

- Robustness of the empirical determination of H_0 .
- Cepheid vs. TRGB calibrators

Calibrator	TRGB	Cepheids
LMC	DEB^{a}	DEB^{a}
NGC 4258	masers^b	masers^{b}
Milky Way	$\omega~{\rm Cen}~{\rm DEB}^c$	EDR3 parallaxes ^{d}
SMC	DEB^{e}	
\overline{a} Pietrzyński (2019)		
b Reid et al. (2019)		
c Thompson et al. (2001)		
d Riess et al. (2021)		
e Graczyk et al. (2020)		



- Hubble tension: Early and late universe probes result in >4-sigma difference in the Hubble constant (67 vs. 73 Km/s/Mpc).
- Using four anchors to calibrate the TRGB method, Freedman (2021) derived

 $H_0 = 69.8 \pm 0.6 \text{ (stat)} \pm 1.6 \text{ (sys)} \text{ km s}^{-1} \text{ Mpc}^{-1}$

This is more consistent with the early universe measurements.

• The discrepancy in the late universe probes (TRGB and cepheid) needs to be resolved in order to understand the nature of Hubble tension.

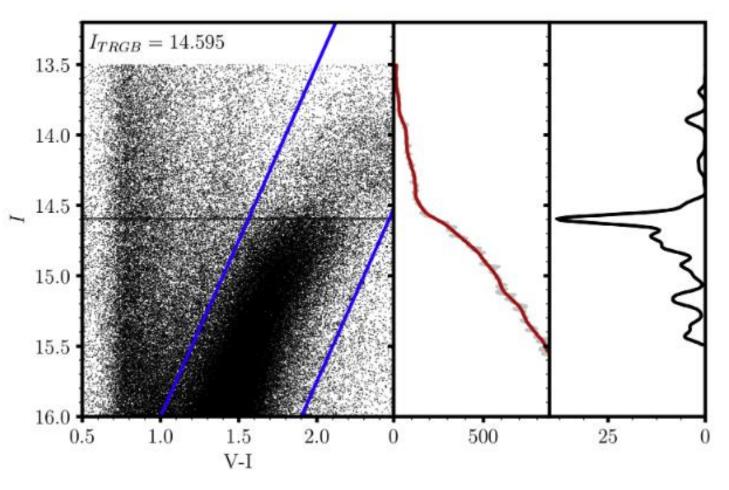
Questions I would ask as an audience

- 1. What types of galaxies can TRGB method apply to? How is it compared to the cepheid method?
- 2. What is the advantage of TRGB method over cepheid method?
- 3. Why is TRGB independent of stellar metallicity?
- 4. How will future (larger) telescopes help in resolving the discrepancy?

Back-up slides

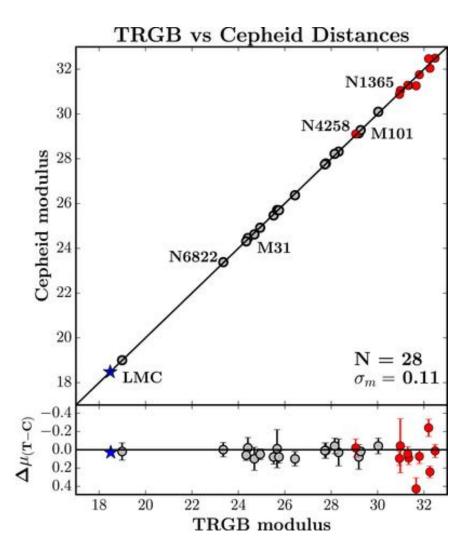
How to determine the TRGB

• Sobel edge-detection filter.



Common hosts of cepheid and TRGB

 Comparison of published TRGB and Cepheid distance moduli for 28 nearby galaxies. The distances span a range from 50 kpc to 30 Mpc.



Hubble constant measurements over time

