# A bimodal burst energy distribution of a repeating fast radio burst source

D Li et al. 2021

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#### Overview towards FRB

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Signal Milky Way Host Galaxy IGM  $DM = \int n_e dl$  $t_2 - t_1 = DM \frac{2\pi^2 e^2 c}{m_e} (\nu_1^{-2} - \nu_2^{-2})$  $DM = DM_{MW} + DM_{IGM} + DM_{host}$ 

 $DM_{host}$  is unknown, so it's quite hard to measure how far the source is.

Lorimer et al. 2007



# Observational profiles



Keane et al. 2018

#### FRB121102



Repeated Bursts Arecibo Detection

Spitler et al. 2016 Chatterjee 2017



z = 0.193 $D_L = 949 \text{ kpc}$ 

# FAST Observation





- 59.5 hours observation towards FRB121102 in two months.
- Two peaks in energy distribution
- LogNormal and Cauchy model to fit burst rate curve

$$N(E) = \frac{N_0}{\sqrt{2\pi}\sigma_E E} \exp\left[\frac{-(\log E - \log E_0)^2}{2\sigma_E^2}\right] + \frac{\epsilon_E}{1 + (E/E_0)^{\alpha_E}},$$
$$E_0 = 7.2 \times 10^{37} \text{ erg} \qquad \alpha_E = 1.85$$
$$E_E = 3 \times 10^{38} \text{ erg}$$

### FAST Observation



Extremely high rate disfavors the models which need expensive trigger to each burst.

#### Mechanism



#### 'Pulsar Like'

'GRB Like'

Lu et al.2020 Metzger et al. 2019

# FRB from Magnetar



### Take Home Message

1) FAST detected FRB121102 for 59.5 hours and 1652 events were derived.

- Energy distribution can be fitted by Lognormal and Cauchy model, which implies complicated mechanism.
- 3) Periodic search got no result.
- 4) Coherent emission model of neutron star is preferred.





#### **Open Questions**

1) Are all FRBs repeaters? If not, what're their mechanism?

2) Are there engines other than magnetars that can power FRBs? If so, what are they?

3) How is FRB generated from magnetar? Magnetosphere or relativistic shock?