ARE THEY ULTRA-SPECIAL?

ULTRA-DIFFUSE GALAXIES

SO...IT IS POPULAR...

At least 201 papers since 2015...

ultra diffuse galaxies edit

description edit

My ADS library

Number of Papers:

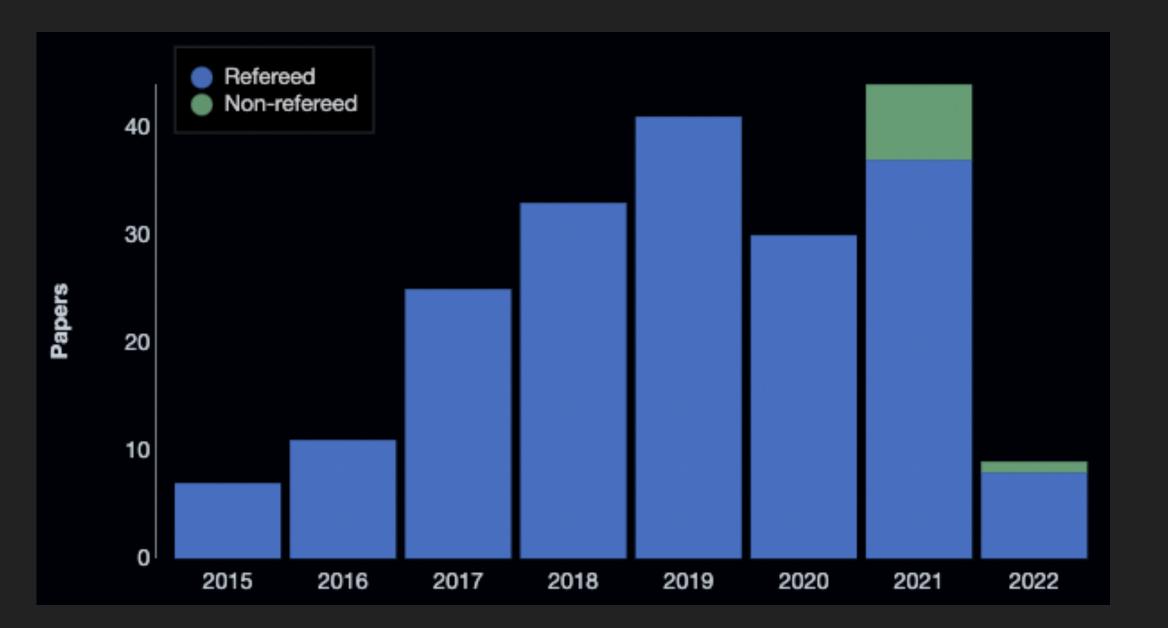
Date Created:

Mar 3 2022, 10:54pm

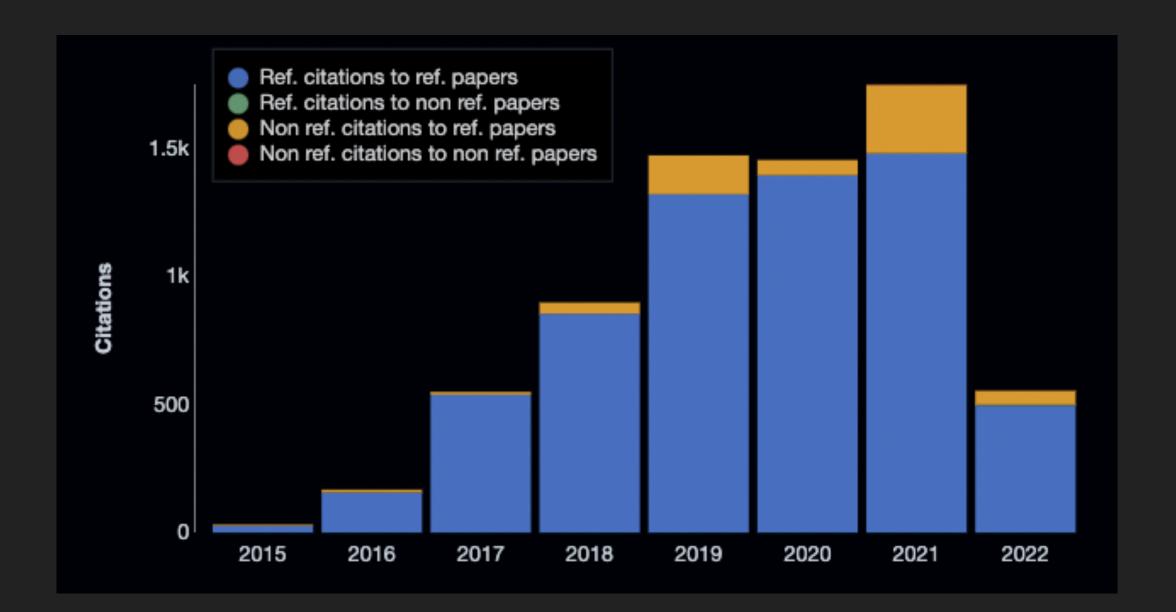


The library is public if you are interested...

Publication



Citation

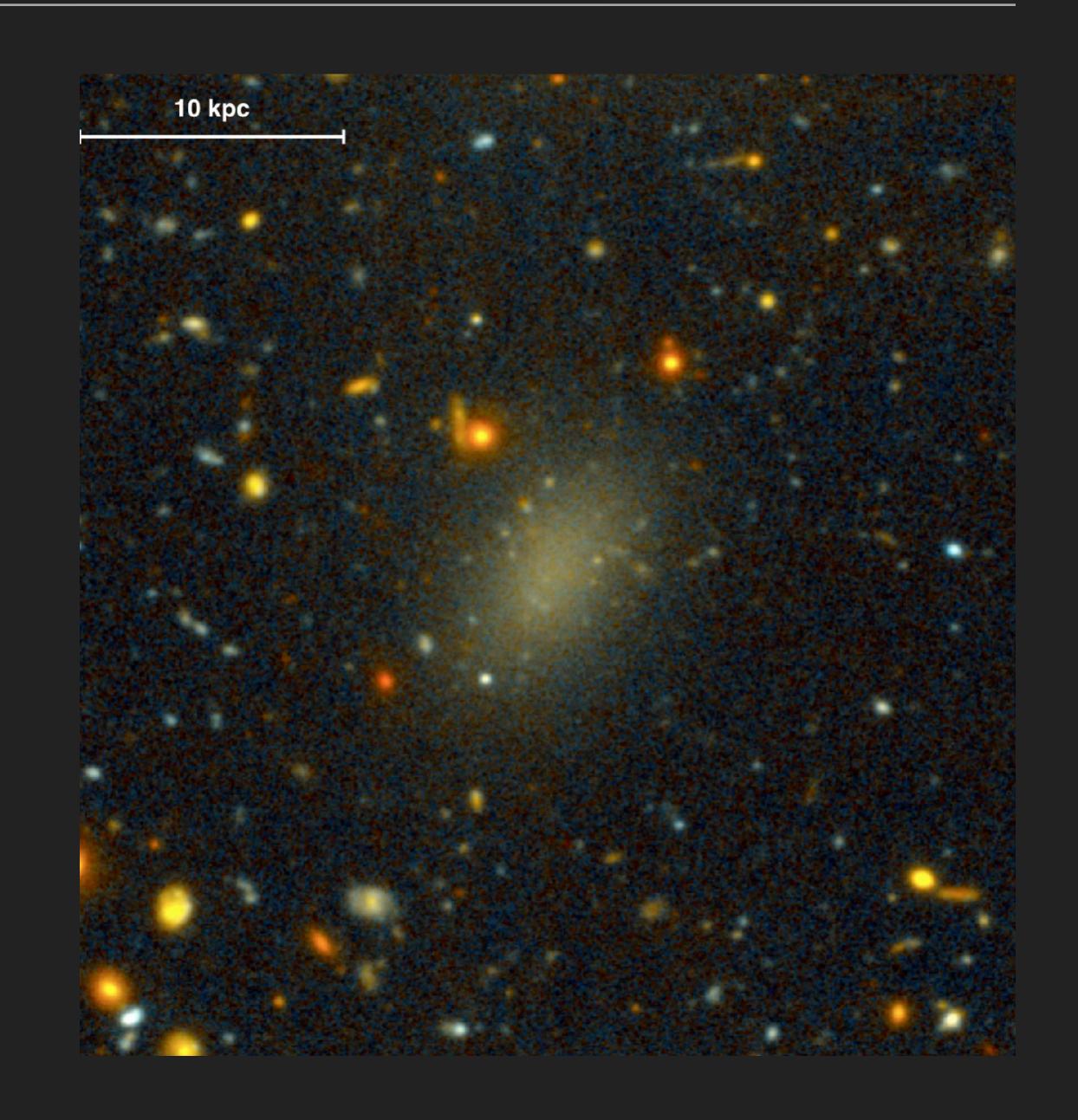


A PERSONAL VIEW

- From an observational point-of-view
- I am very biased
- It is highly incomplete...

WHAT I WILL COVER TODAY

- > 7 basic facts (given what we know for now)
- ≥ 2 controversial (and open) questions
- $\triangleright 2$ positive and 2 negative aspects of this topic

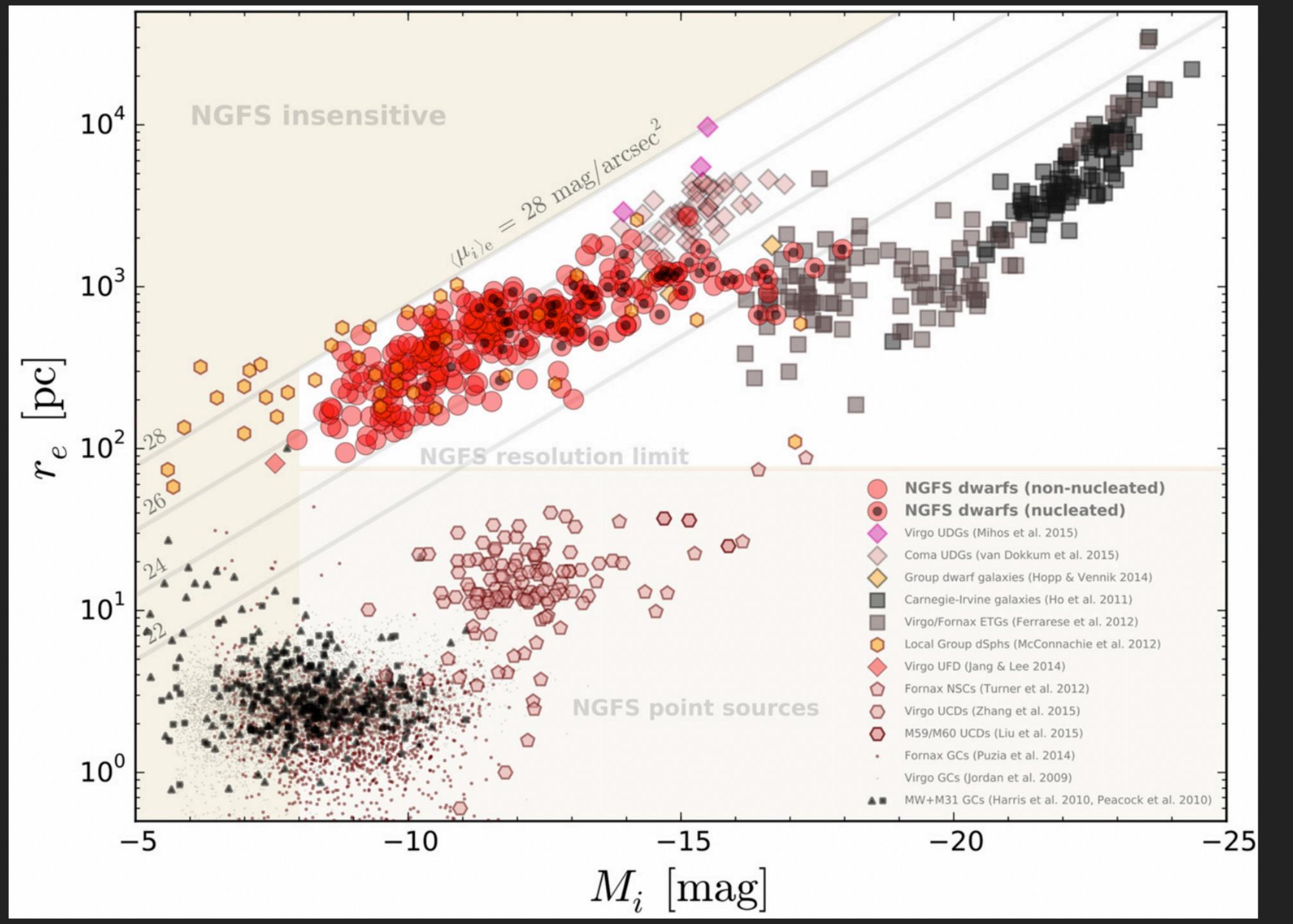


The Spitzer Infrared Nearby Galaxies Survey (SINGS) Hubble Tuning-Fork The Spitzer Space Telescope observed 75 galaxies as part of its SINGS (Spitzer Infrared Nearby Galaxies Survey) Legacy Program. The galaxies are presented here in a Hubble Tuning-Fork diagram, which groups galaxies according to the morphology of their nuclei and spiral arms. The designation of these galaxies and their placement in the diagram is based on their visible-light appearance. The main goal of the SINGS program is to characterize the infrared properties of a wide range of galaxy types. The images of the galaxies are composites created from data taken by IRAC (the Infrared Array Camera) at 3.6 and 8.0 µm, and MIPS (the Multiband Imaging Photometer for Spitzer) at 24 µm. The infrared range probed by these and other observations taken for the SINGS project allows for the detailed study of star formation, dust emission, and the distribution of stars in each galaxy. Light from old stars appears as blue in the images, while the lumpy knots of green and red light are produced by dust clouds surrounding newly born stars. The elliptical galaxies on the left are almost entirely made of old stars, while spiral galaxies like our own Milky Way are rich in young stars and the raw materials for future star formation. More information can be found at: http://sings.stsci.edu/ Intermediate Spirals Irregulars SINGS Team Poster and composite images created from Robert Kennicutt, Jr. (Principle Investigator), Daniela Calzetti (Deputy Principle Investigator), Charles Engelbracht (Technical Contact), Lee Armus, George Bendo, Caroline Bot, Brent Buckalew, John Cannon, Daniel Dale, Bruce Draine, Karl Gordon, Albert Grauer, David Hollenbach, Tom Jarrett, Lisa SINGS observations by Karl D. Gordon (Oct 2007) Blue=IRAC 3.6µm (stars) Green=IRAC 8µm (aromatic features from dust grains/molecules) Kewley, Claus Leitherer, Aigen Li, Sangeeta Malhotra, Martin Meyer, John Moustakas, Eric Murphy, Michael Regan, George Rieke, Marcia Rieke, Helene Roussel, Kartik Sheth, J.D. Smith, Michael Thornley, Fabian Walter & George Helou

So, why do we think UDGs are special?

Or...What is a UDG?

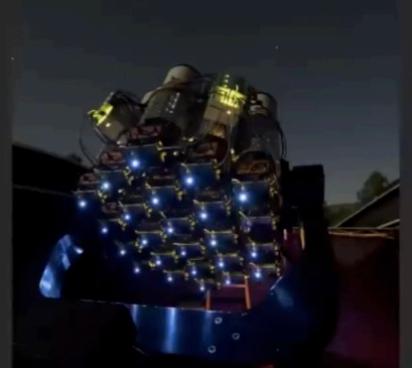
LET'S GET BACK TO THE DRAWING BOARD



Munoz et al. 2015

THE BEGINNING...















Shany Danieli

Dhruba Dutta Chowdhury

Galaxies lacking dark matter

Pieter van Dokkum Shany Danieli Zili Shen Asher Wasserman Dhruba Dutta Chowdhury Allison Merritt Jielai Zhang Deborah Lokhorst Colleen Gilhuly Roberto Abraham Aaron Romanowsky Charlie Conroy Ava Polzin Tim Miller Jiaxuan Li Diederik Kruijssen Michael Keim



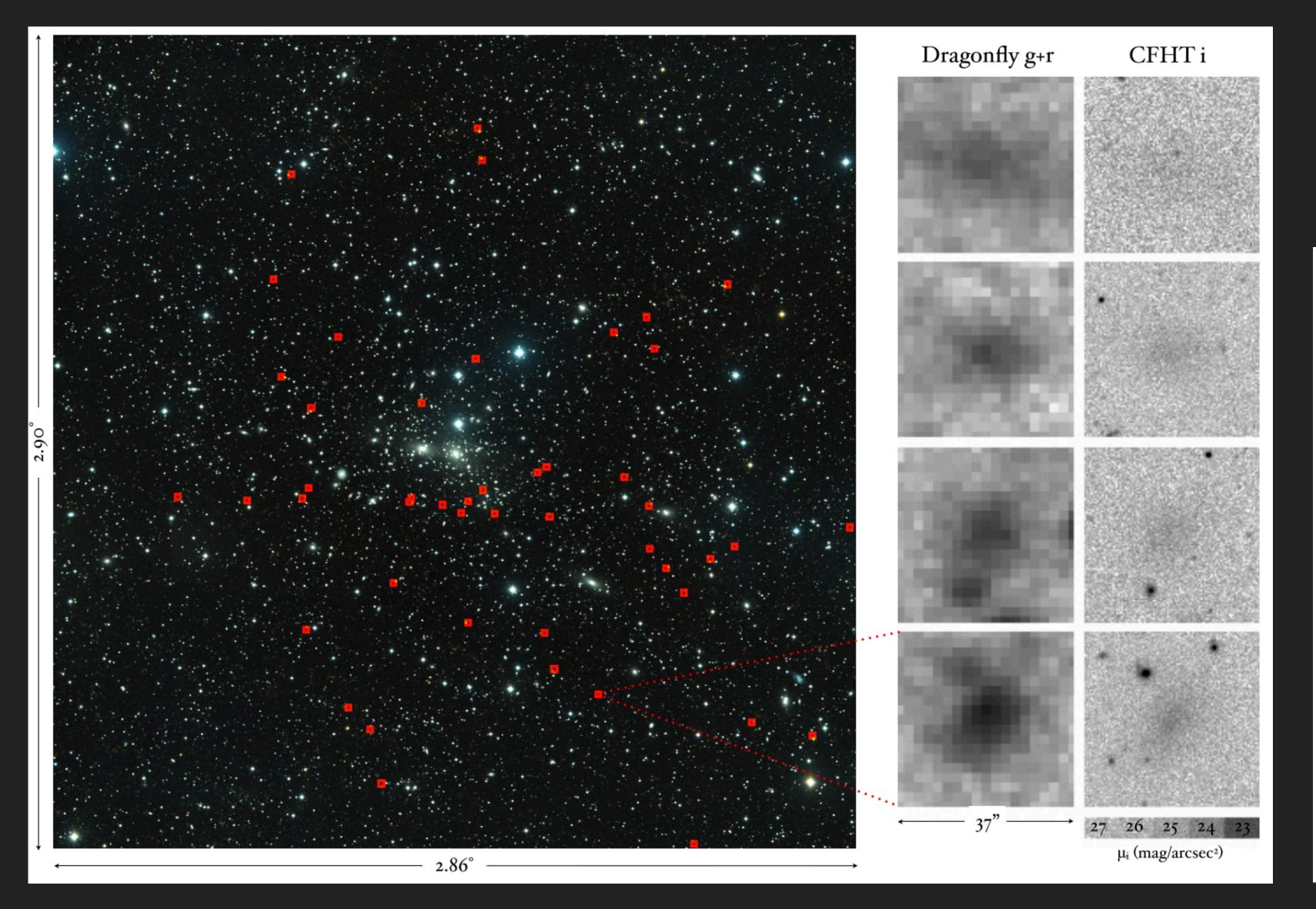




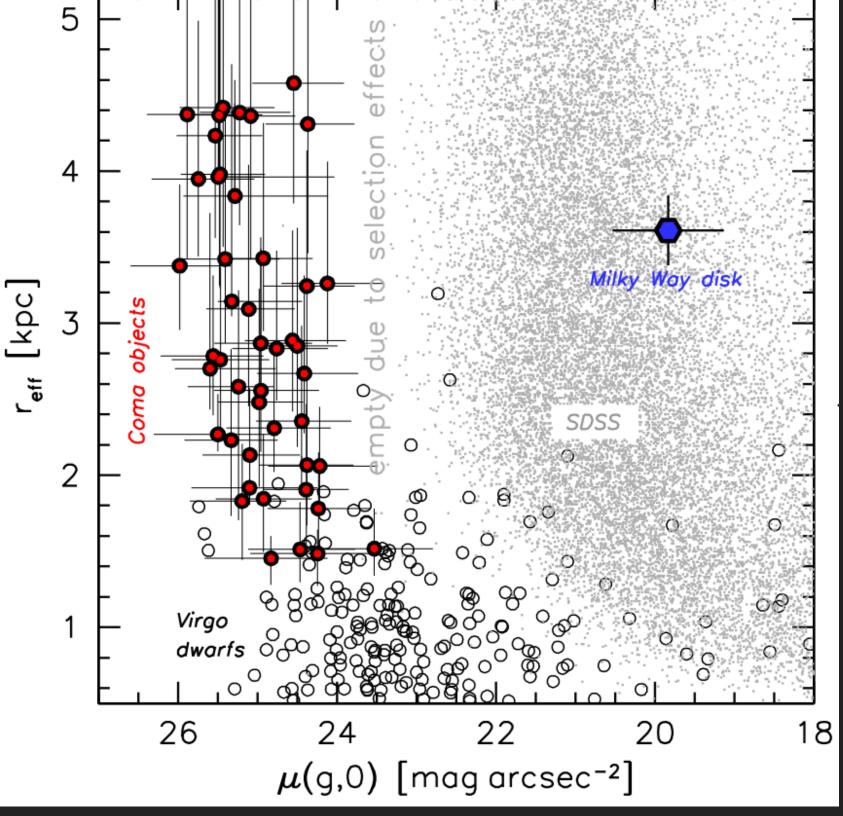




47 "SMUDGES" IN THE COMA CLUSTER



van Dokkum et al. 2015

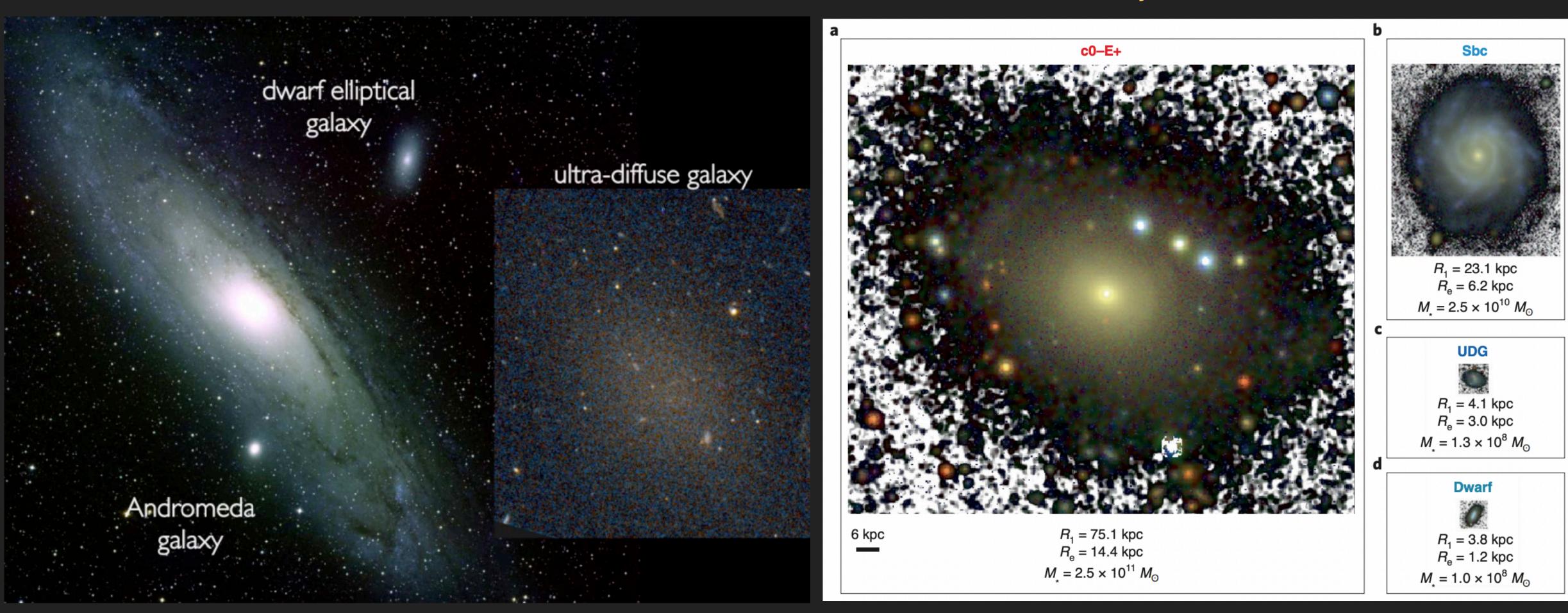


"MILKY WAY SIZE" GALAXY?

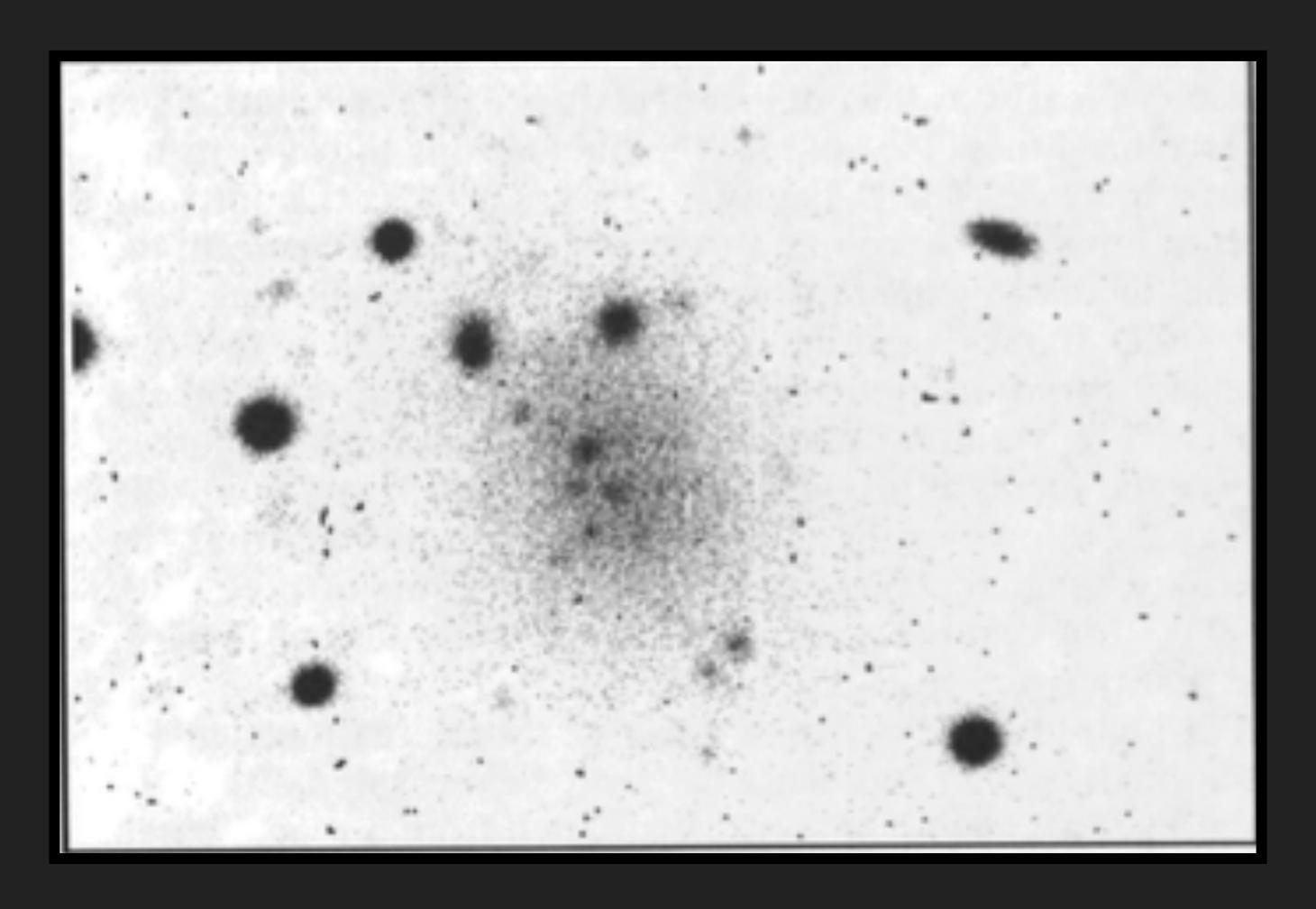
Who's Right?

van Dokkum et al. 2015

Trujillo 2021



Low-Surface Brightness (LSB) Dwarf Galaxies



Bothun+1991

in Fornax

Also see:

Impey+1988;

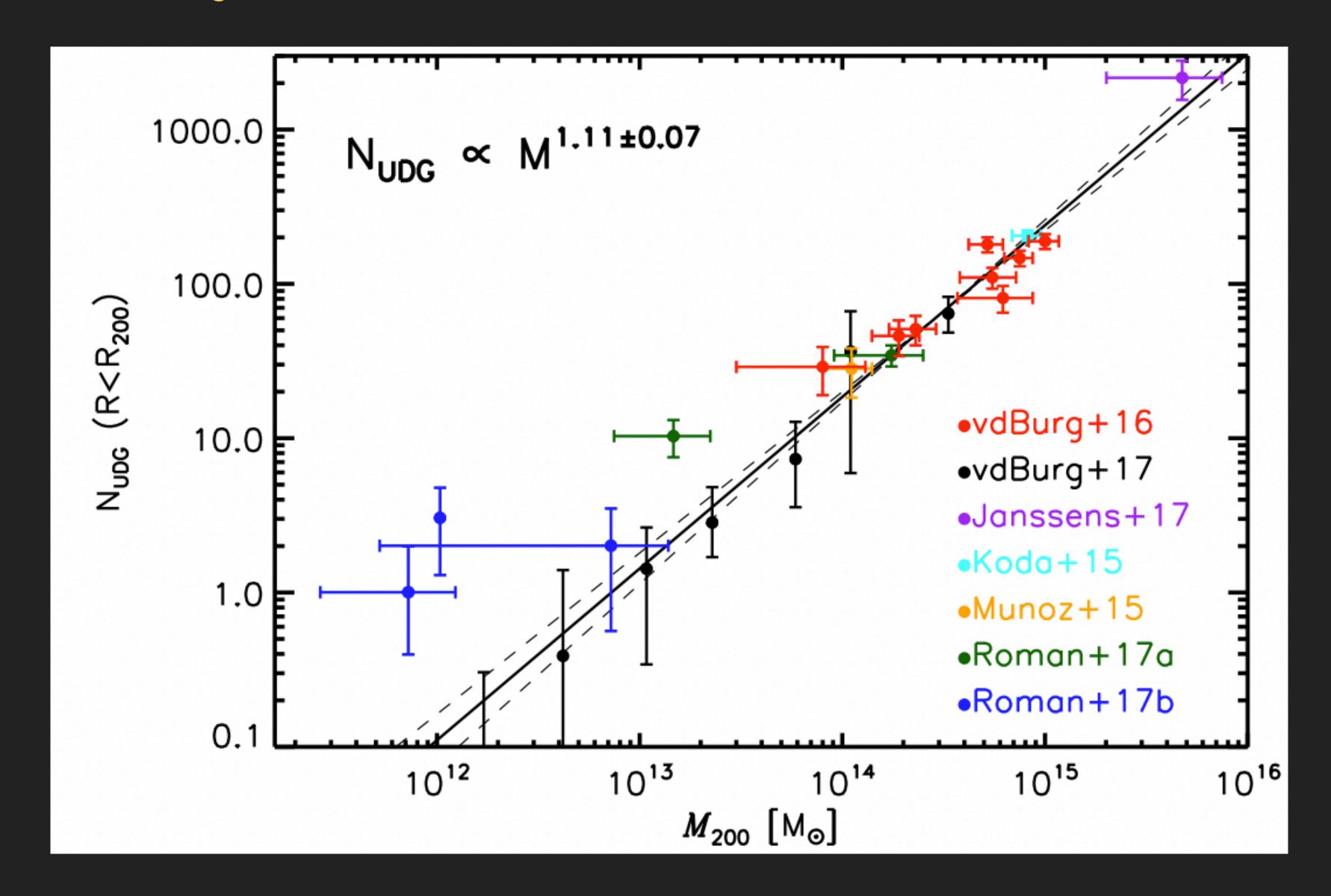
Turner+1993

Dalacanton+1997

Caldwell 2006

Why haven't people notice they are "special" before?

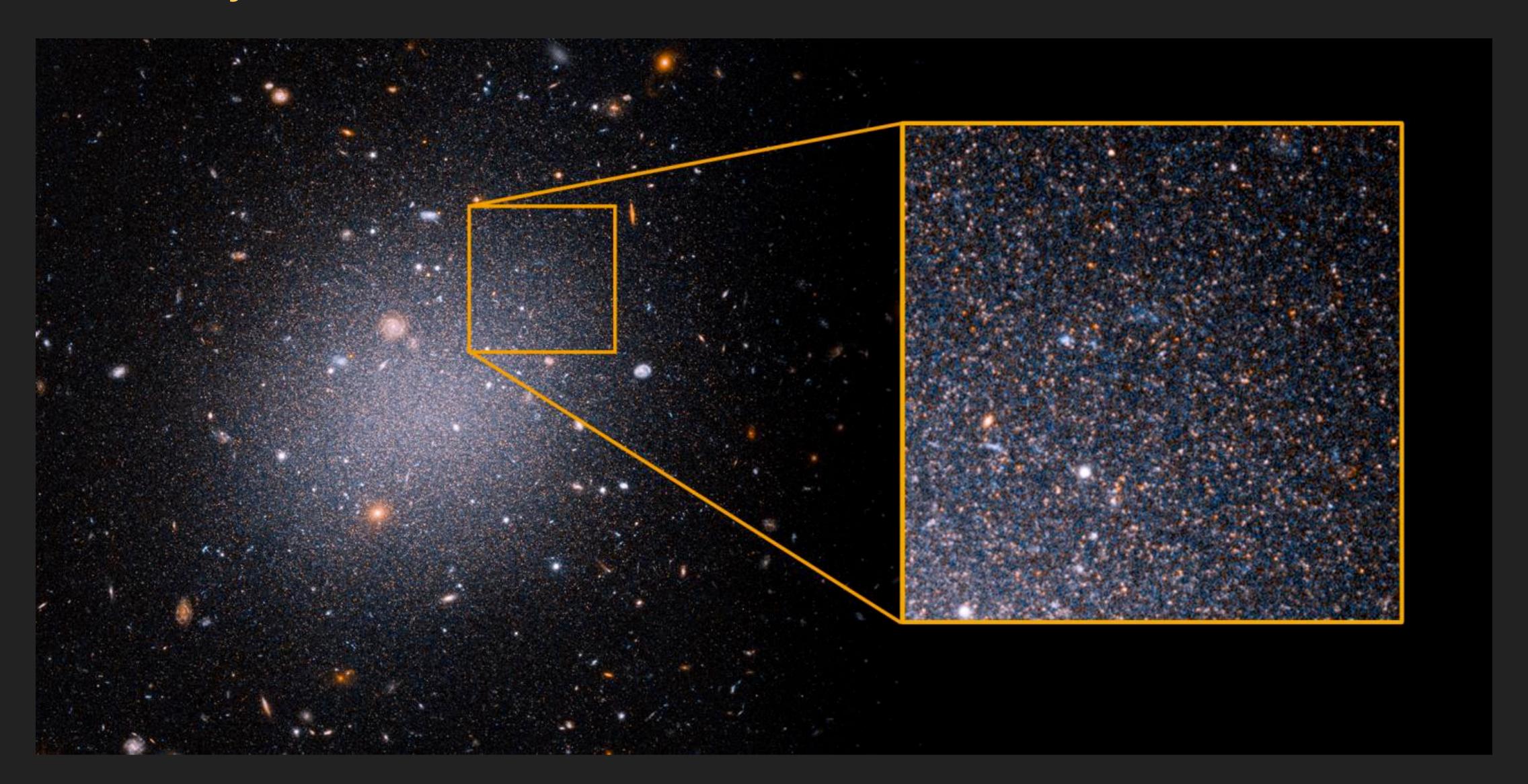
They are everywhere



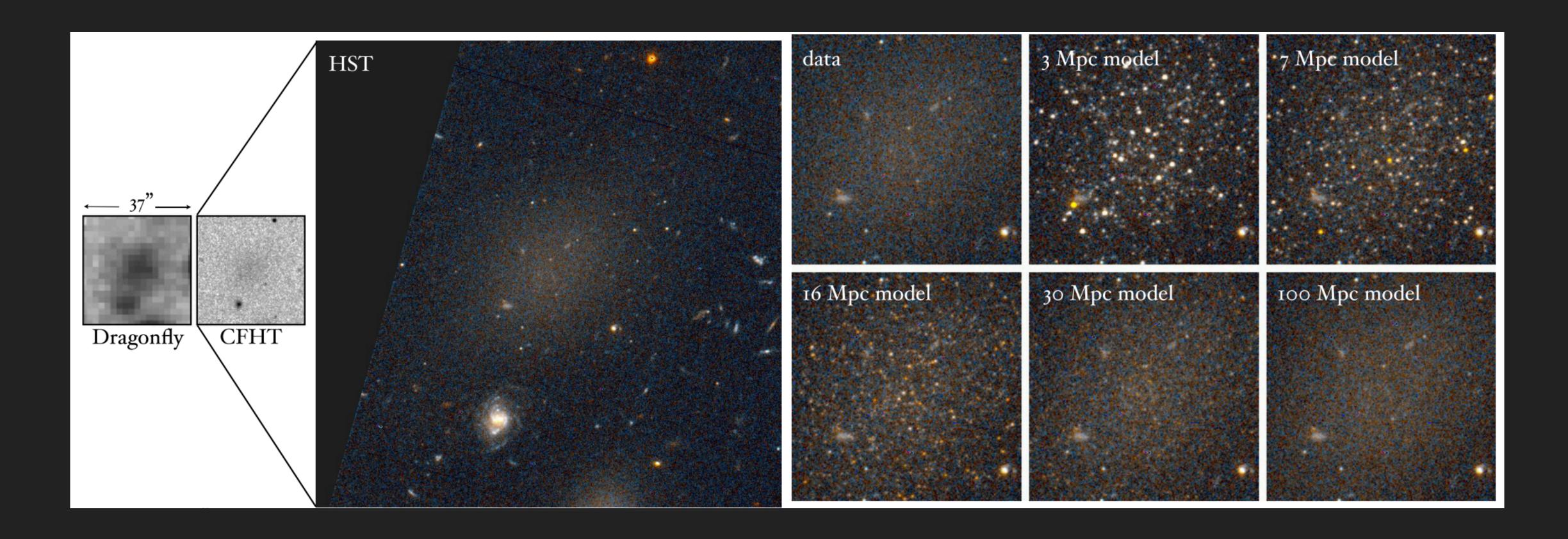


Ah...Distance...What a cursed word in astronomy

Why redshift or line-of-sight velocity is no longer good? How do you measure the distance of them?



Surface Brightness Fluctuation

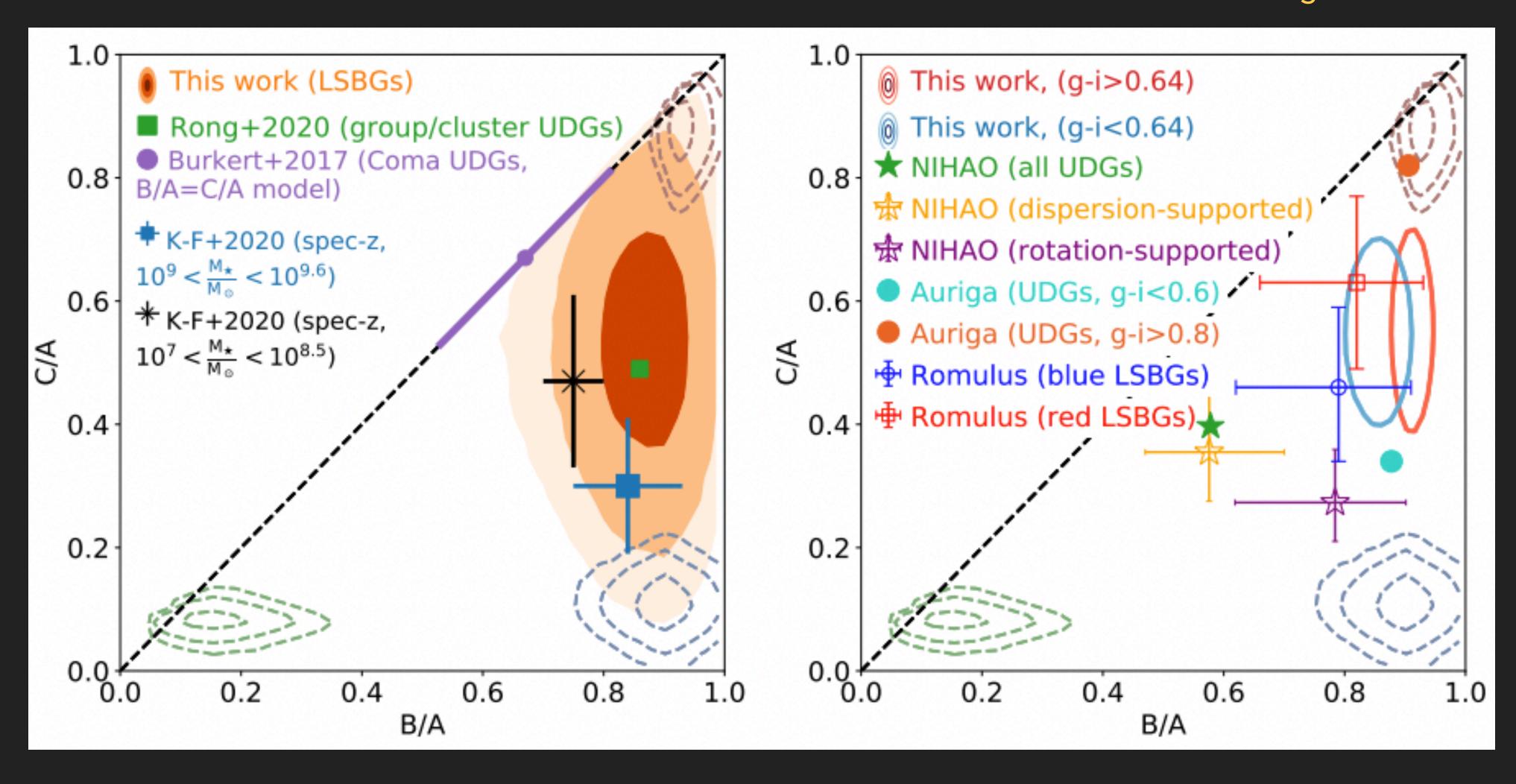


There is a range of looks...

Blue LSBGs			Red LSBGs		
LSBG-305	LSBG-750	LSBG-497	LSBG-435	LSBG-505	LSBG-709
LSBG-157	LSBG-274	LSBG-540	LSBG-17	LSBG-507	LSBG-711
LSBG-621	LSBG-605	LSBG-603	LSBG-490	LSBG-506	LSBG-238
LSBG-522	LSBG-407	LSBG-333	LSBG-159	LSBG-431	LSBG-622

Intrinsic Shape: between a football and a frisby

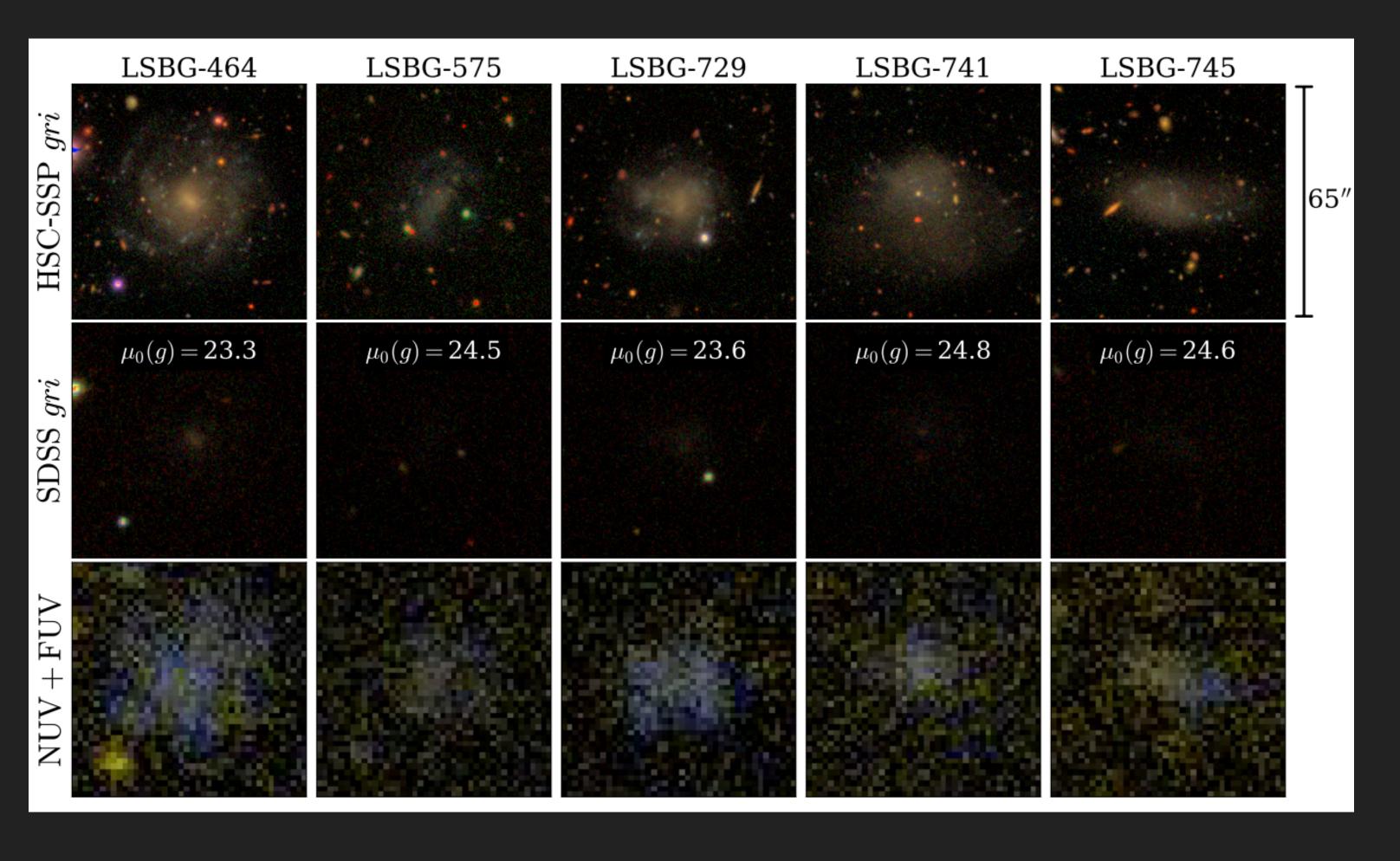
Kado-Fong et al. 2021



Yes, they can form stars; Yes, they can be gas rich

Mostly not in cluster environment; Also has emission line detections

Greco et al. 2018

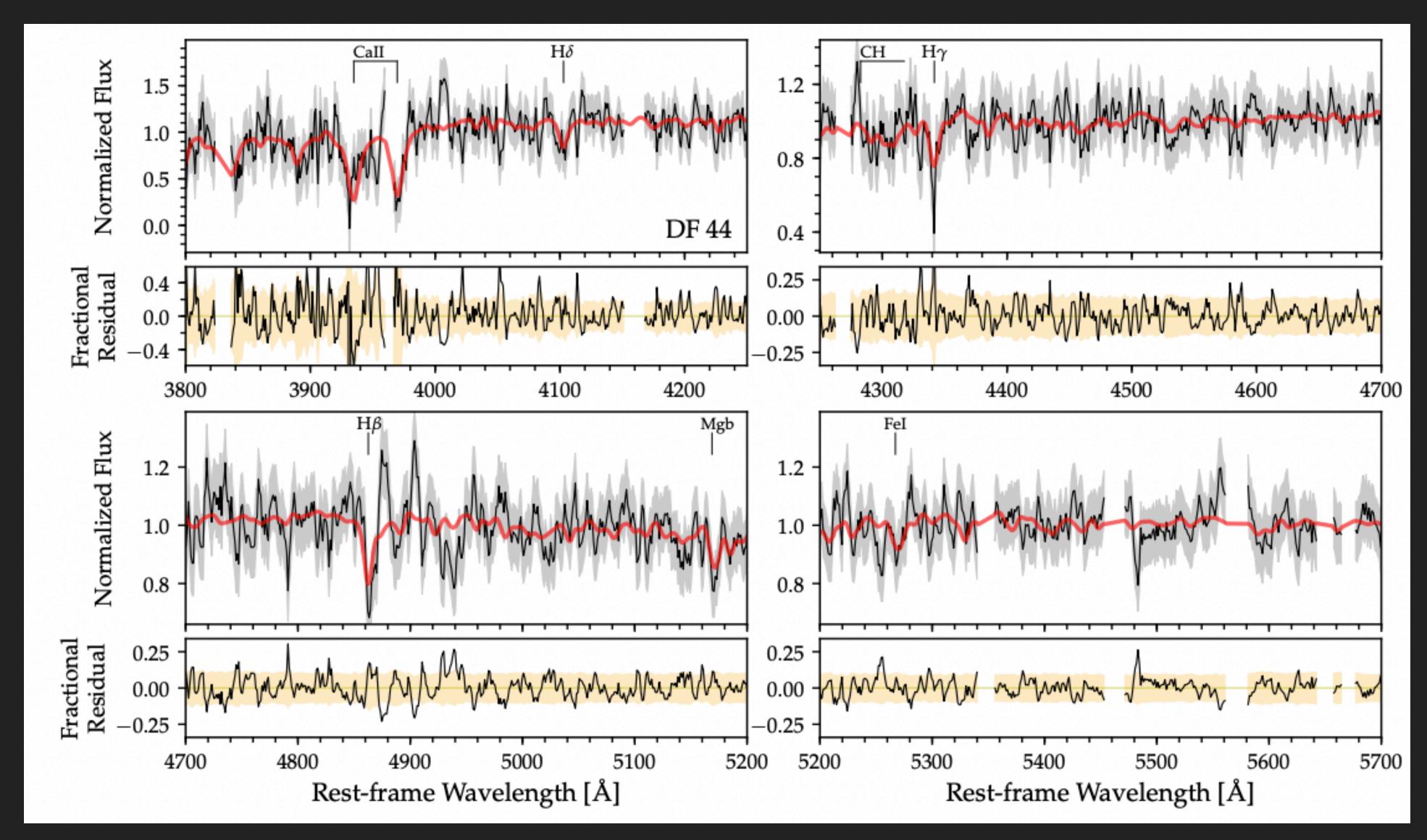


Also see "HI-rich LSB/UDGs"

Stellar population and Star
Formation history are diverse as well

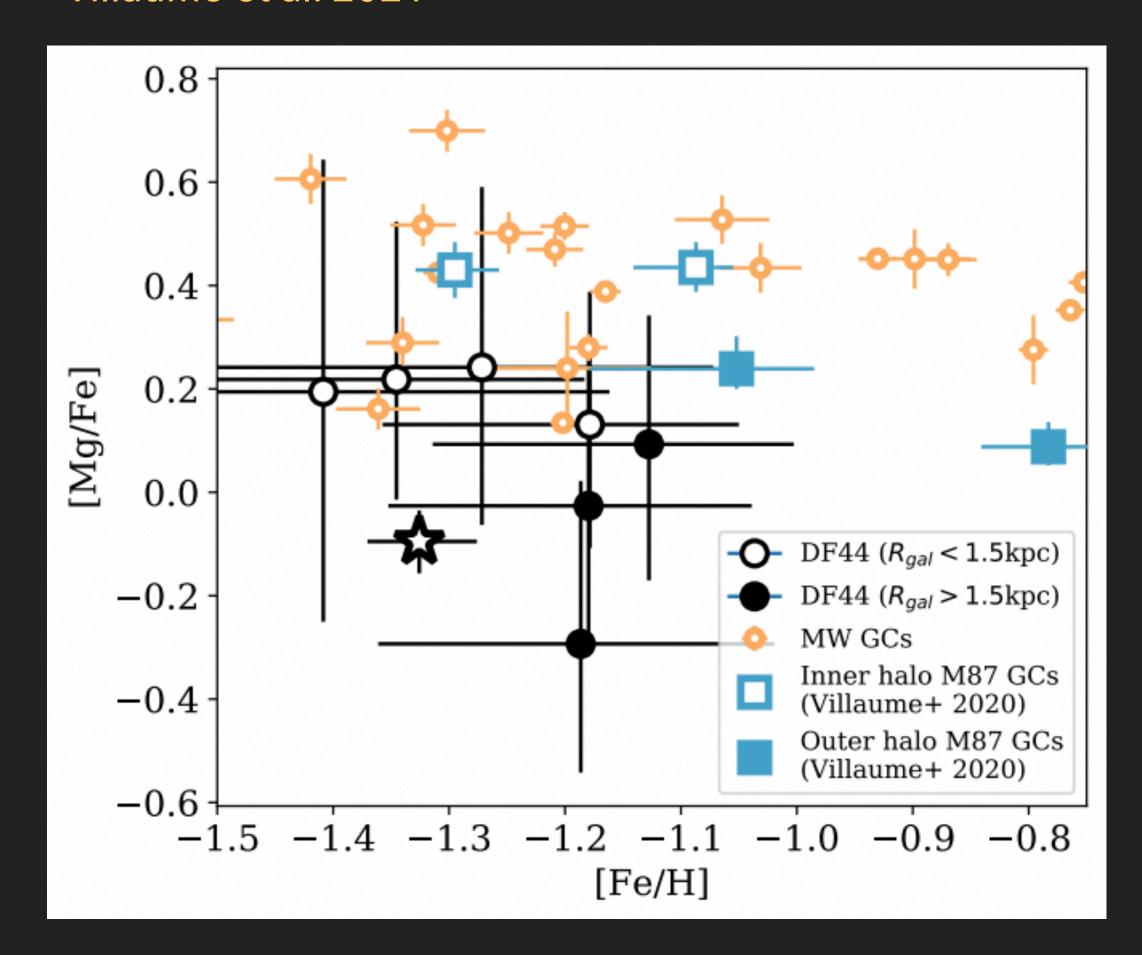
Spectroscopic observation is extremely difficult...

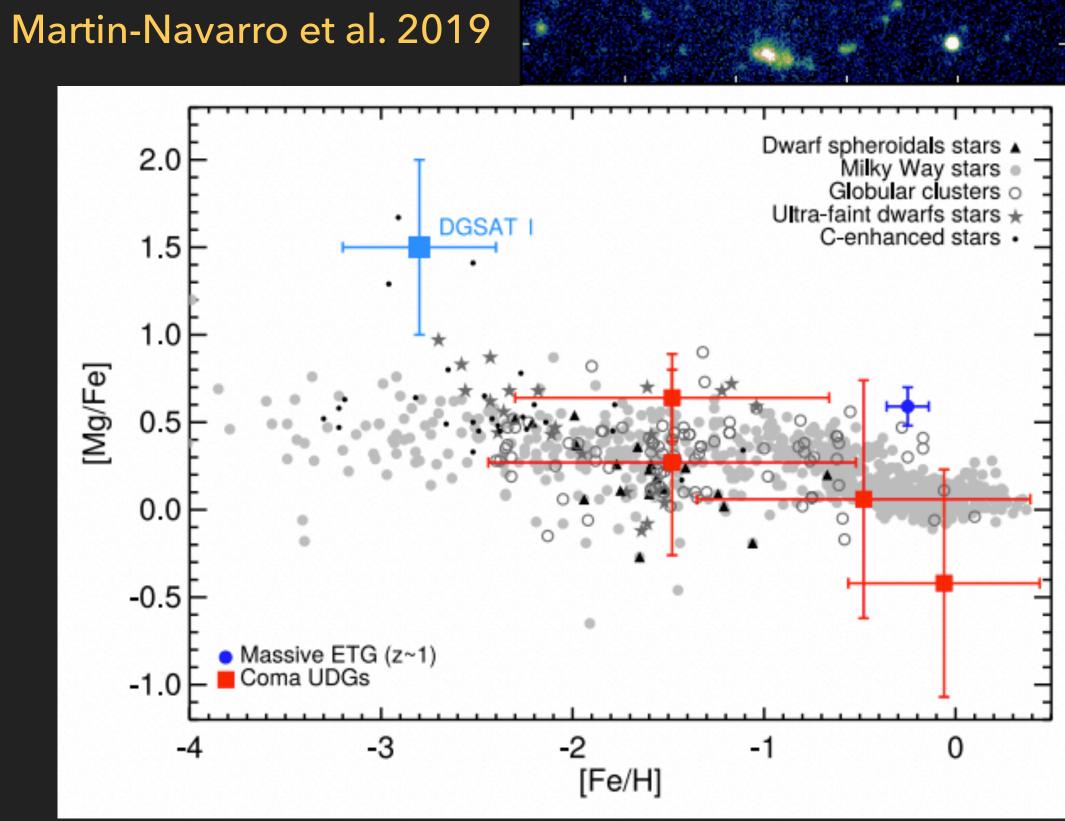
Gu et al. 2017



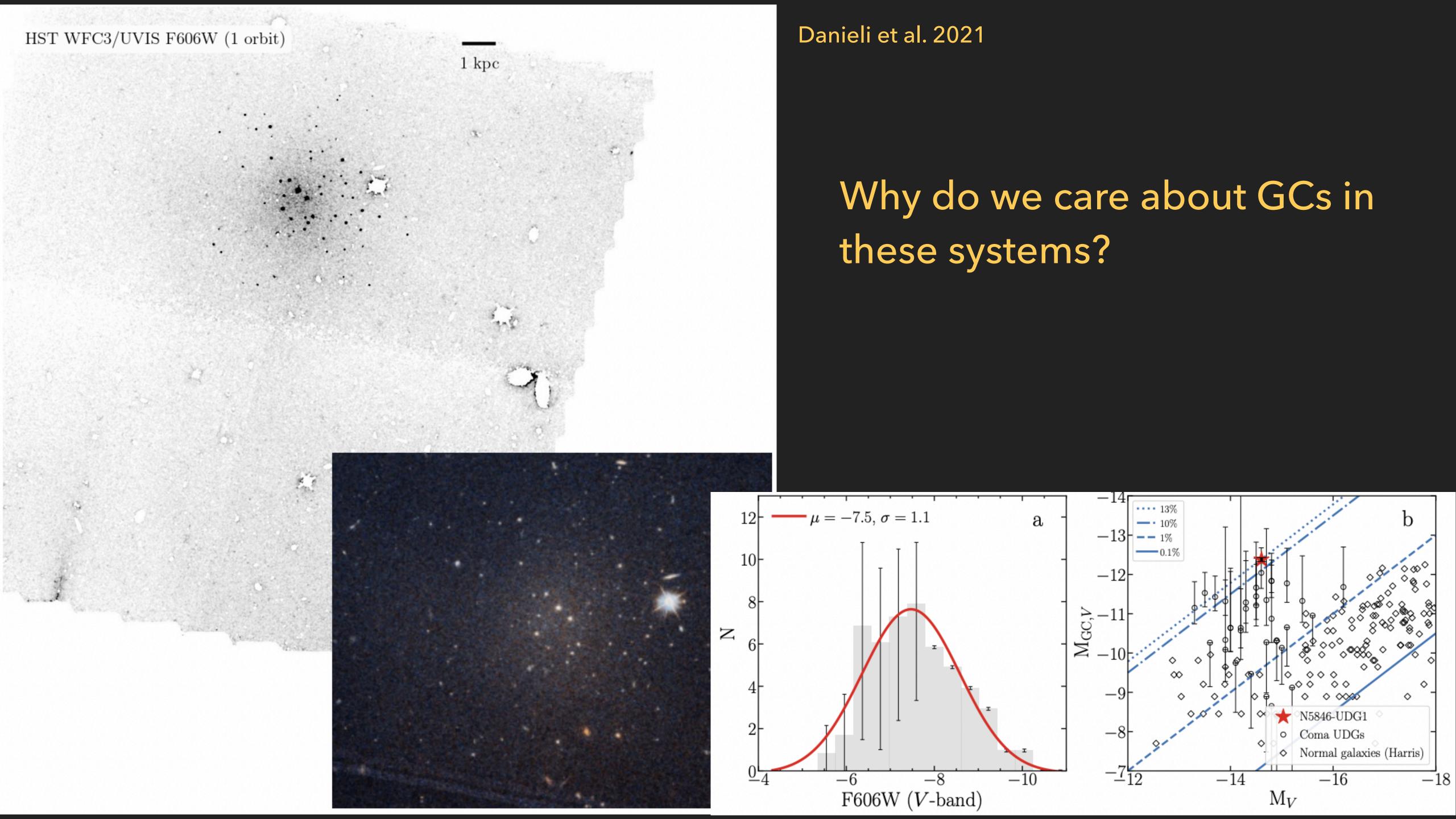
UDGs can show very broad range of star formation efficiency

Villaume et al. 2021



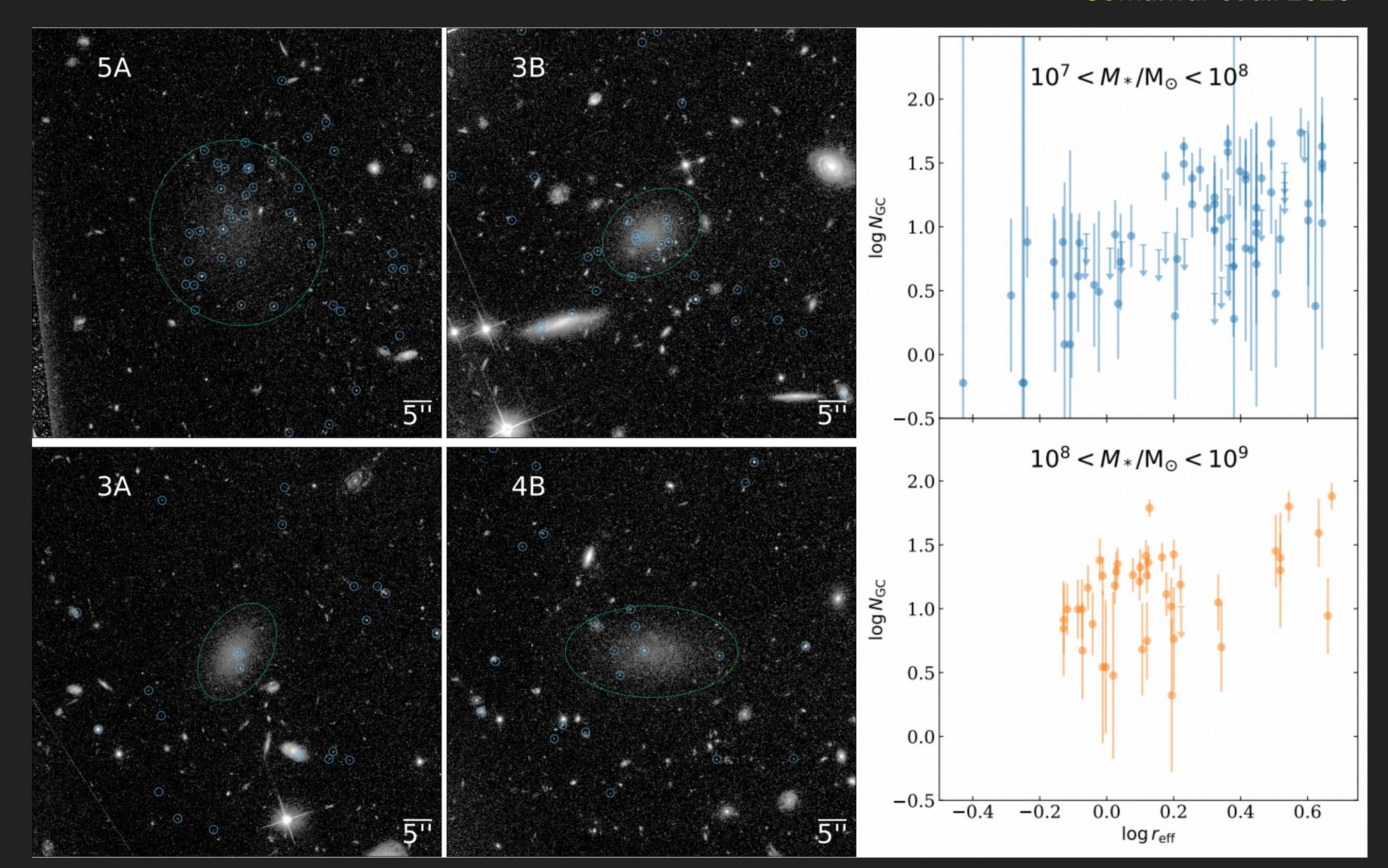


They are mysterious little jewelry boxes...



There is a wide range of GC abundance

Somalwar et al. 2020

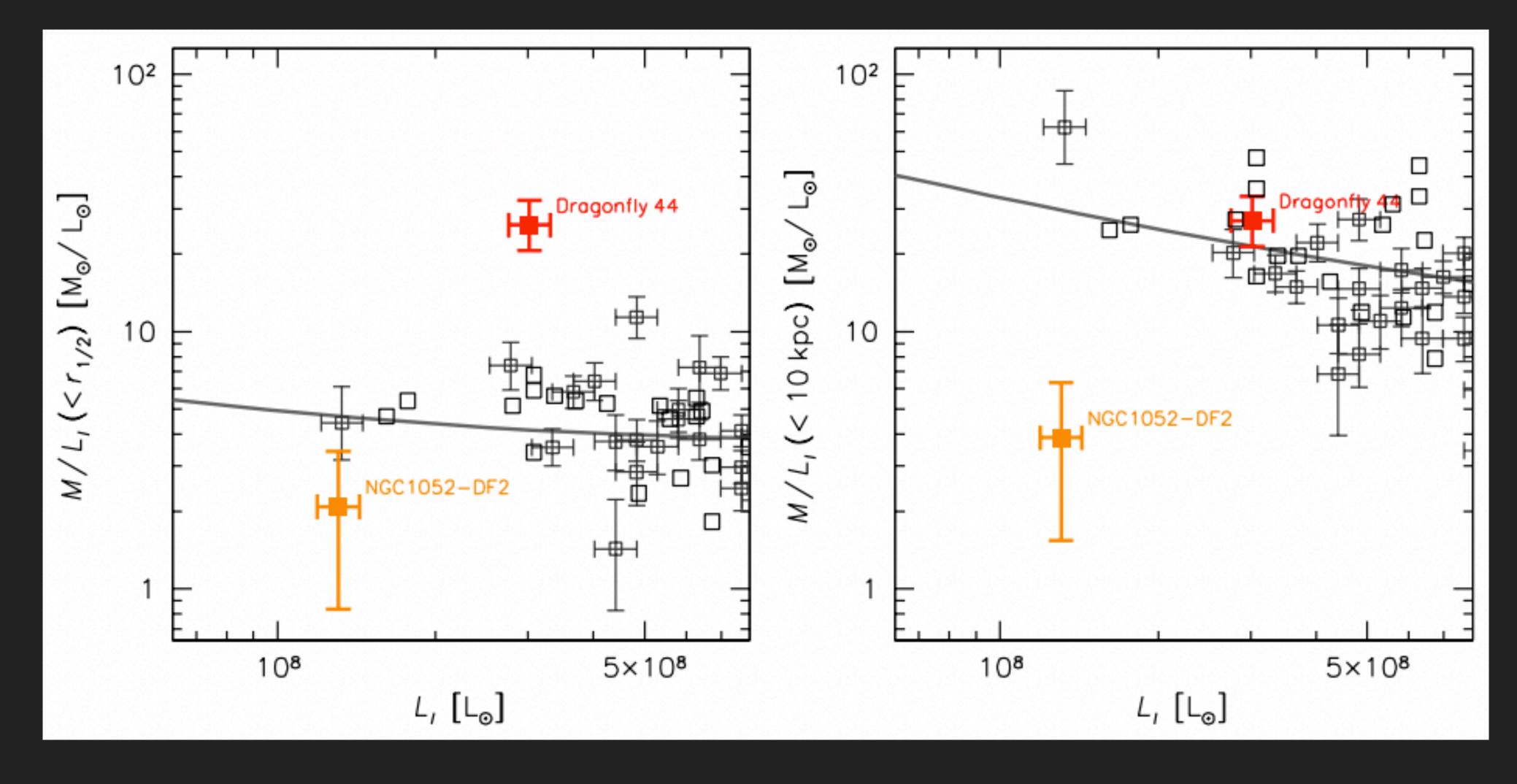


Some have crazy dynamic mass to light ratio

Turned out that some are very dark matter dominated while others are "dark matter free"?

Using the kinematics of both resolved stellar population and the globular cluster systems

van Dokkum et al. 2019

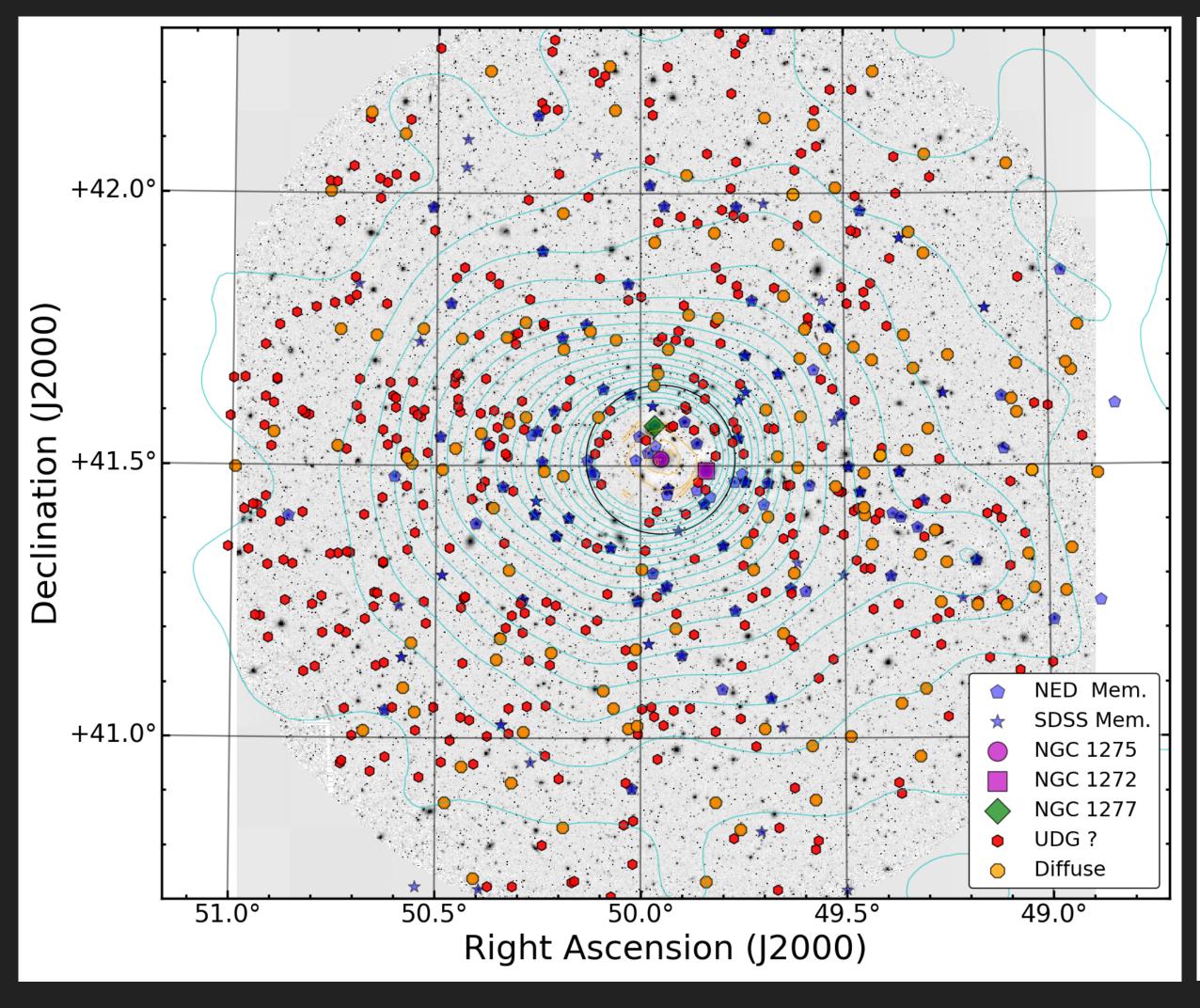


QUESTION

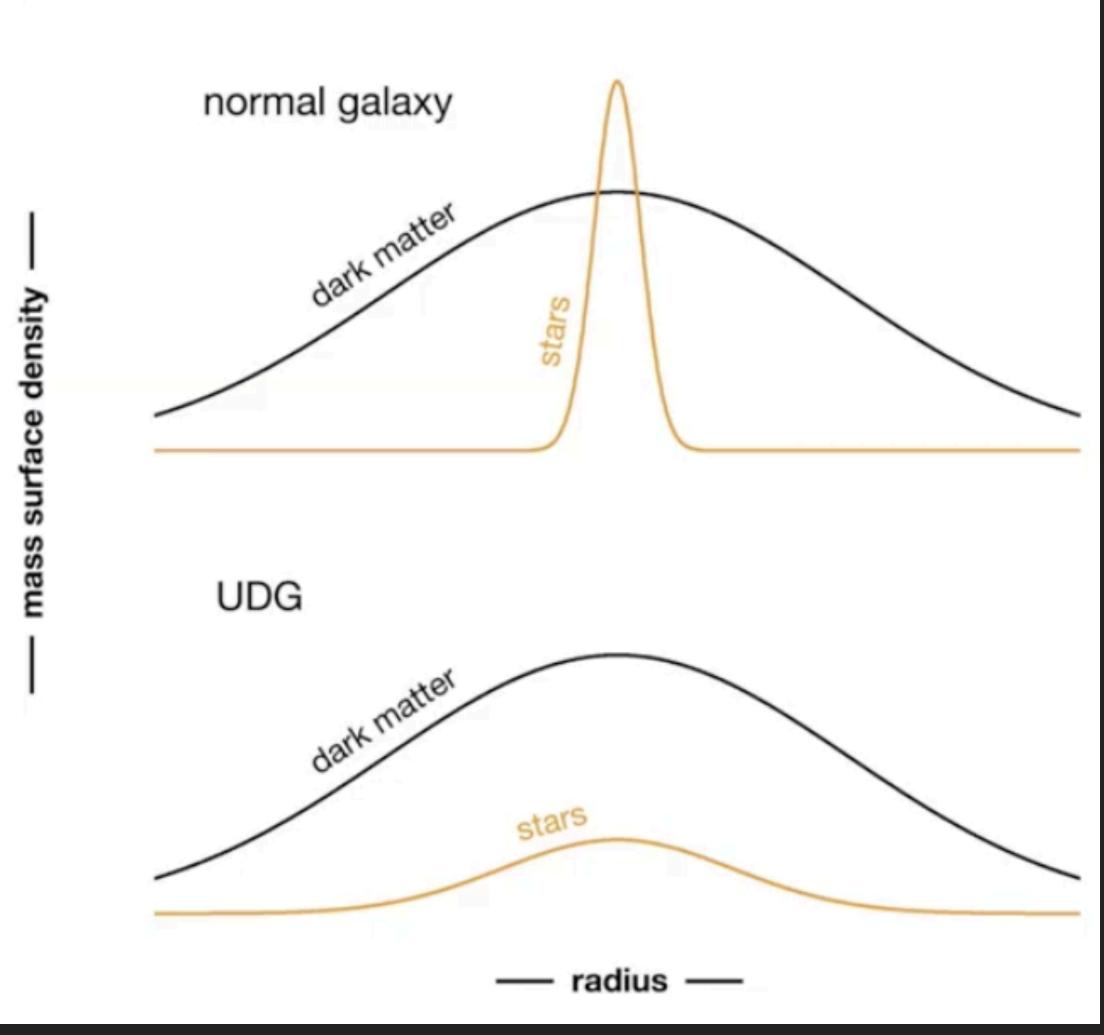
Both DM dominated and DM free? Seriously?

Why do we think the UDGs should be "DM dominated" in the first place?

UDG candidates in the Perseus cluster

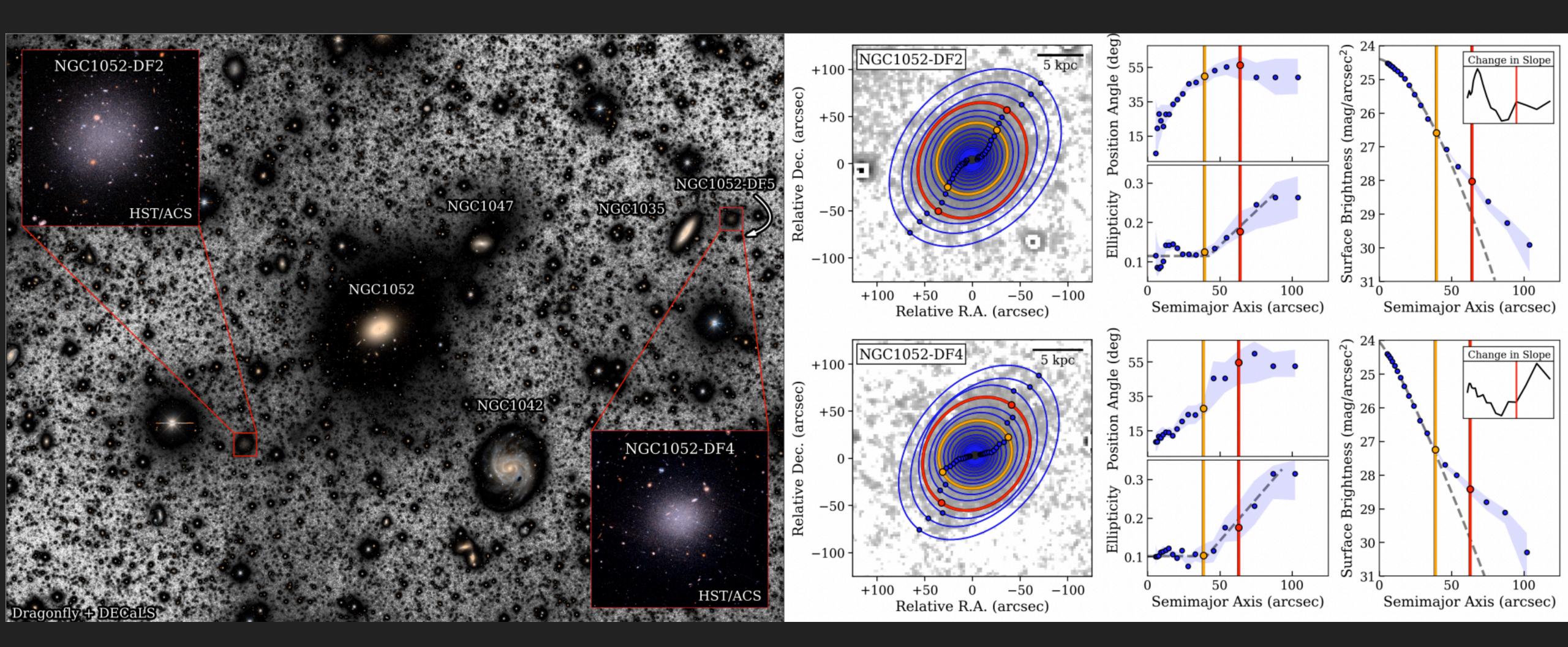


Credit: Pieter van Dokkum



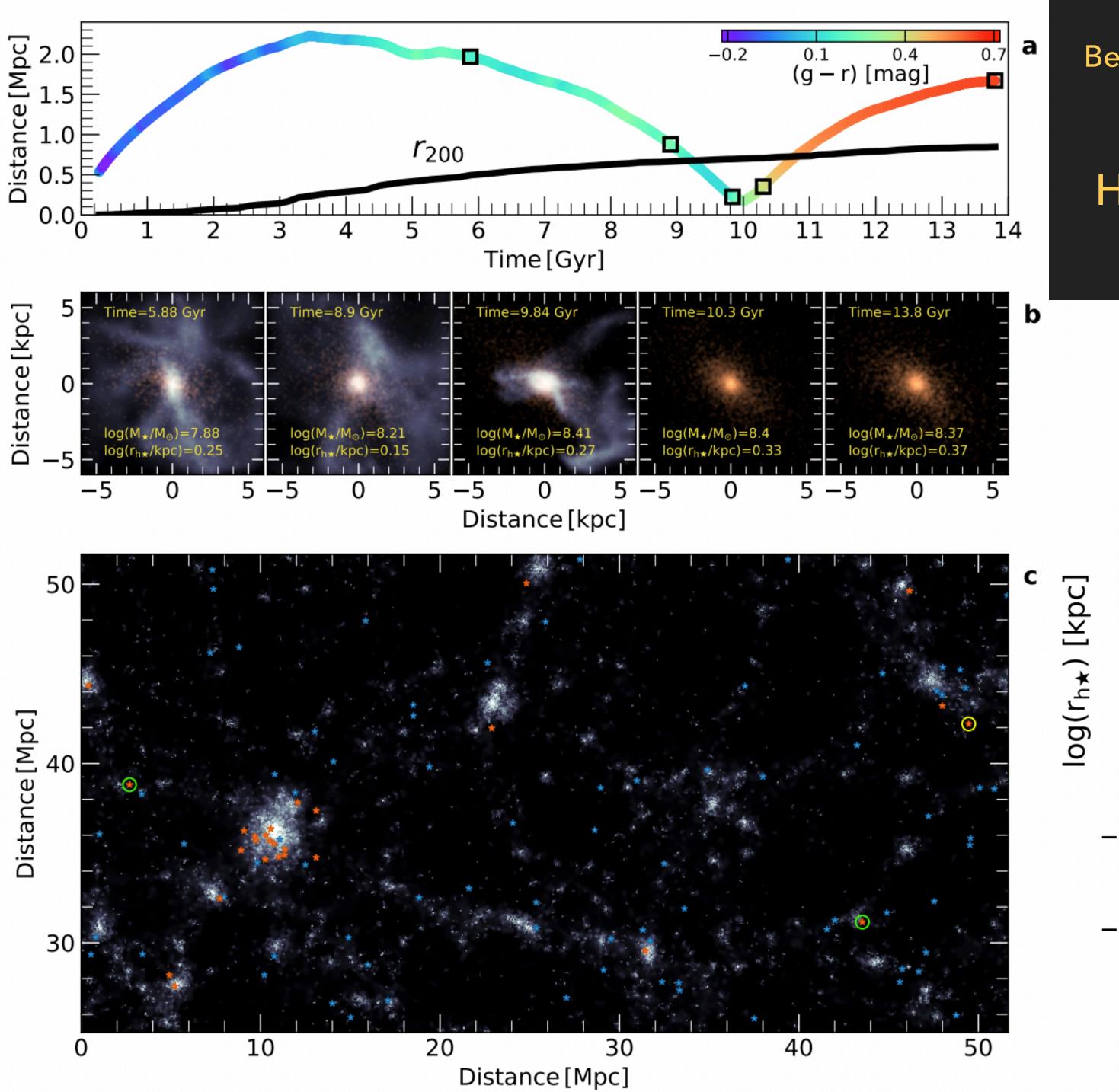
So...they are getting tidal distorted/disrupted anyway...

Keim et al. 2021



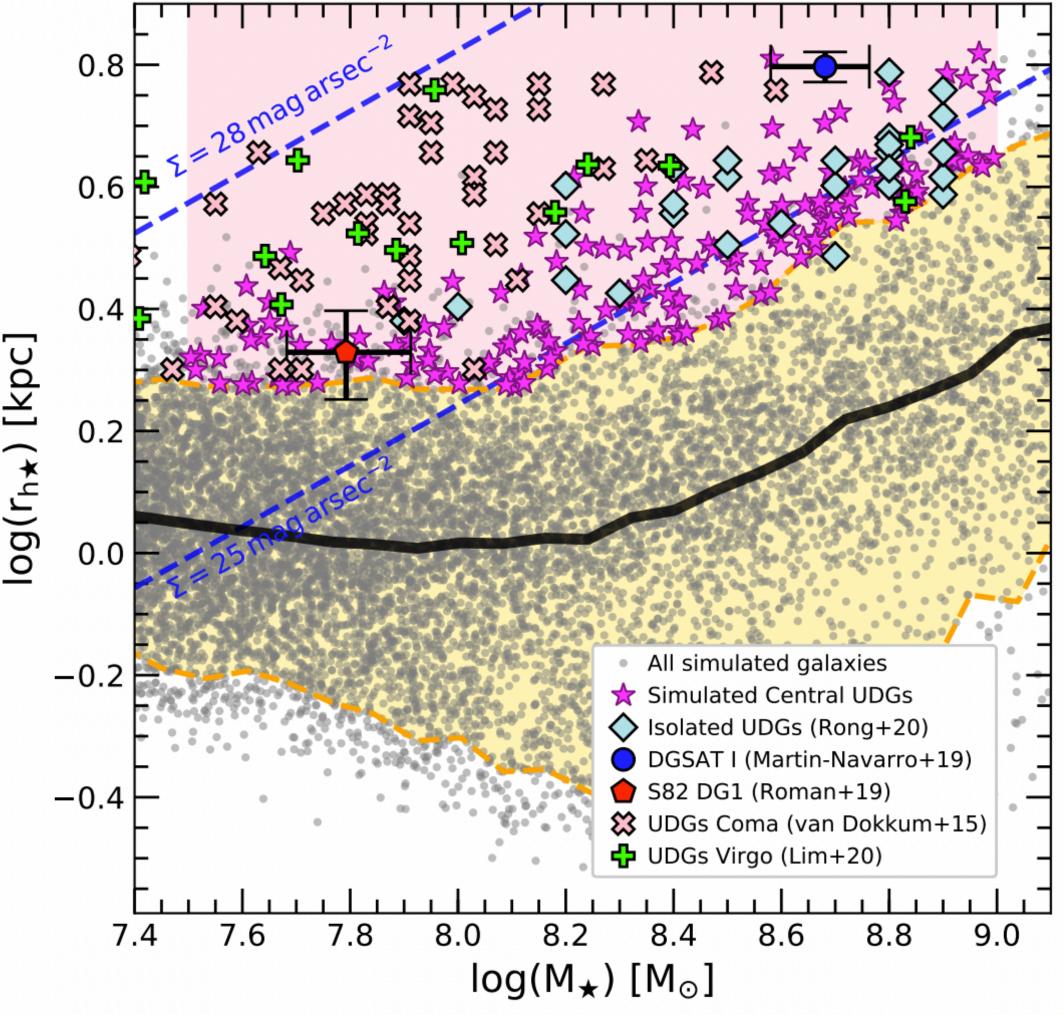
QUESTION

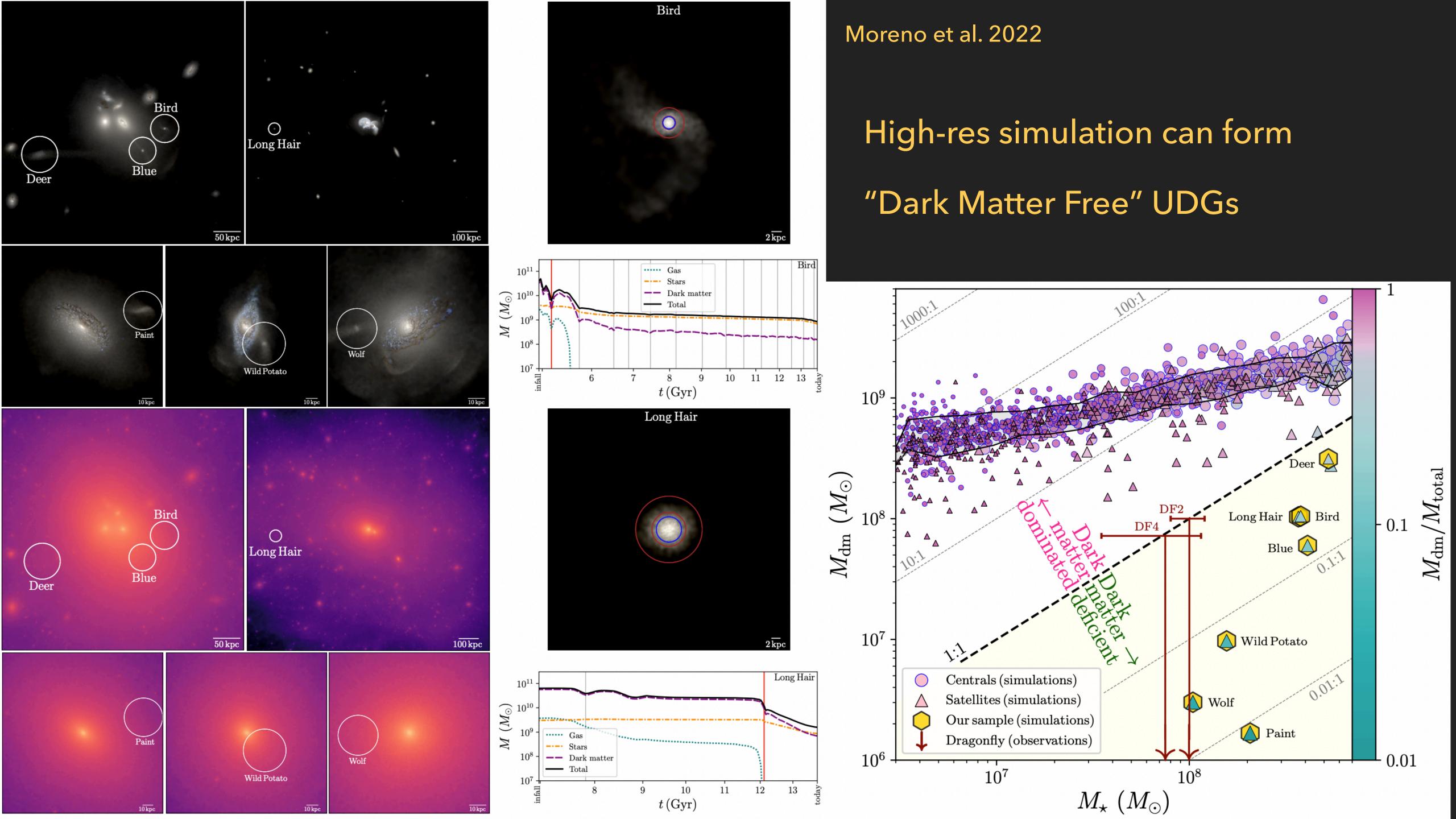
Can they challenge galaxy formation? Dark Matter? Everything we know about the Universe?



Benavides et al. 2021

High-res simulation can form UDGs

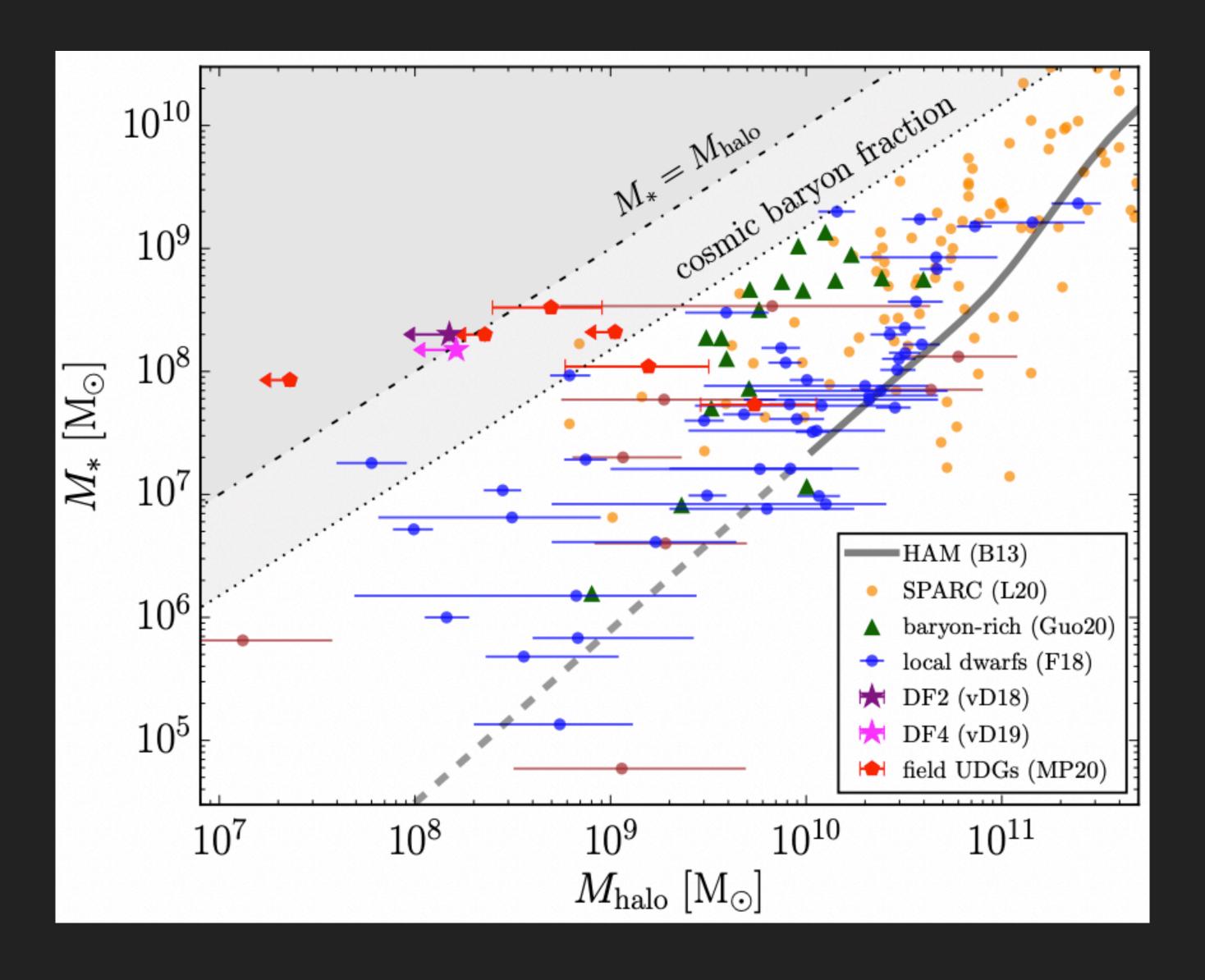




Are we good?

POSITIVE

Force us to think about galaxy formation and halo connection in a new angle

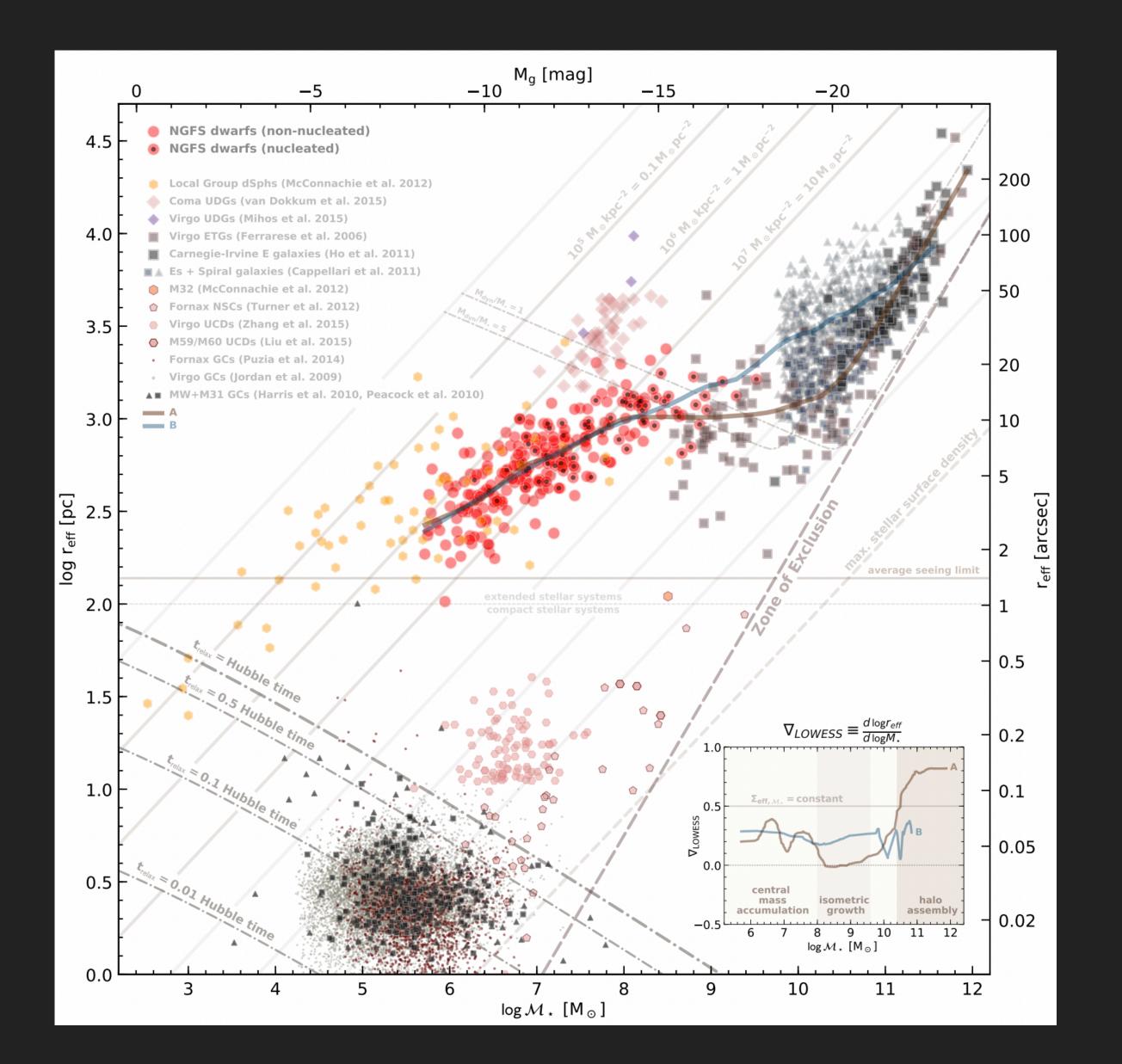


POSITIVE

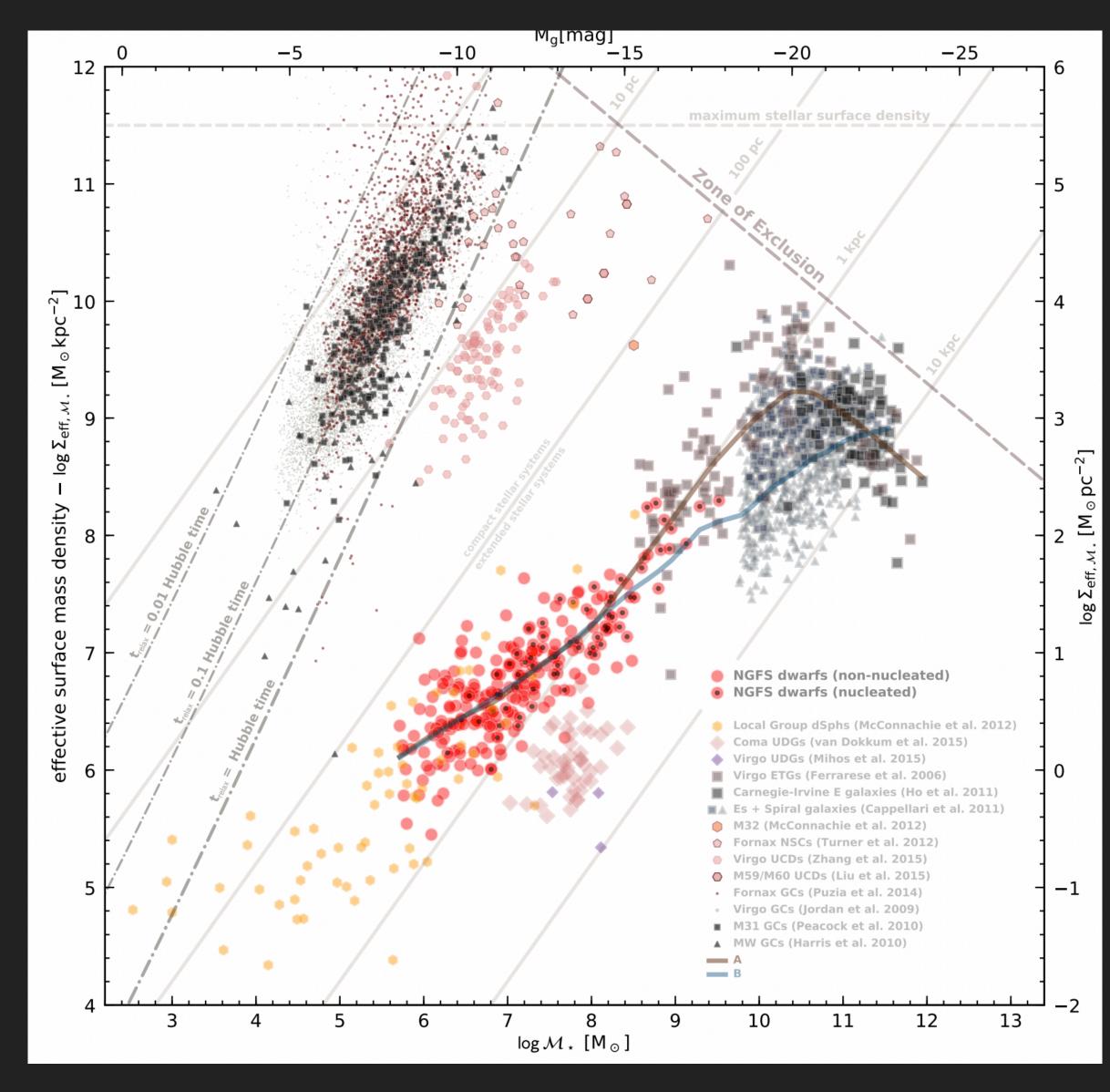
LSB science is back under the spotlight

NEGATIVE

Fail to see the wood for the trees

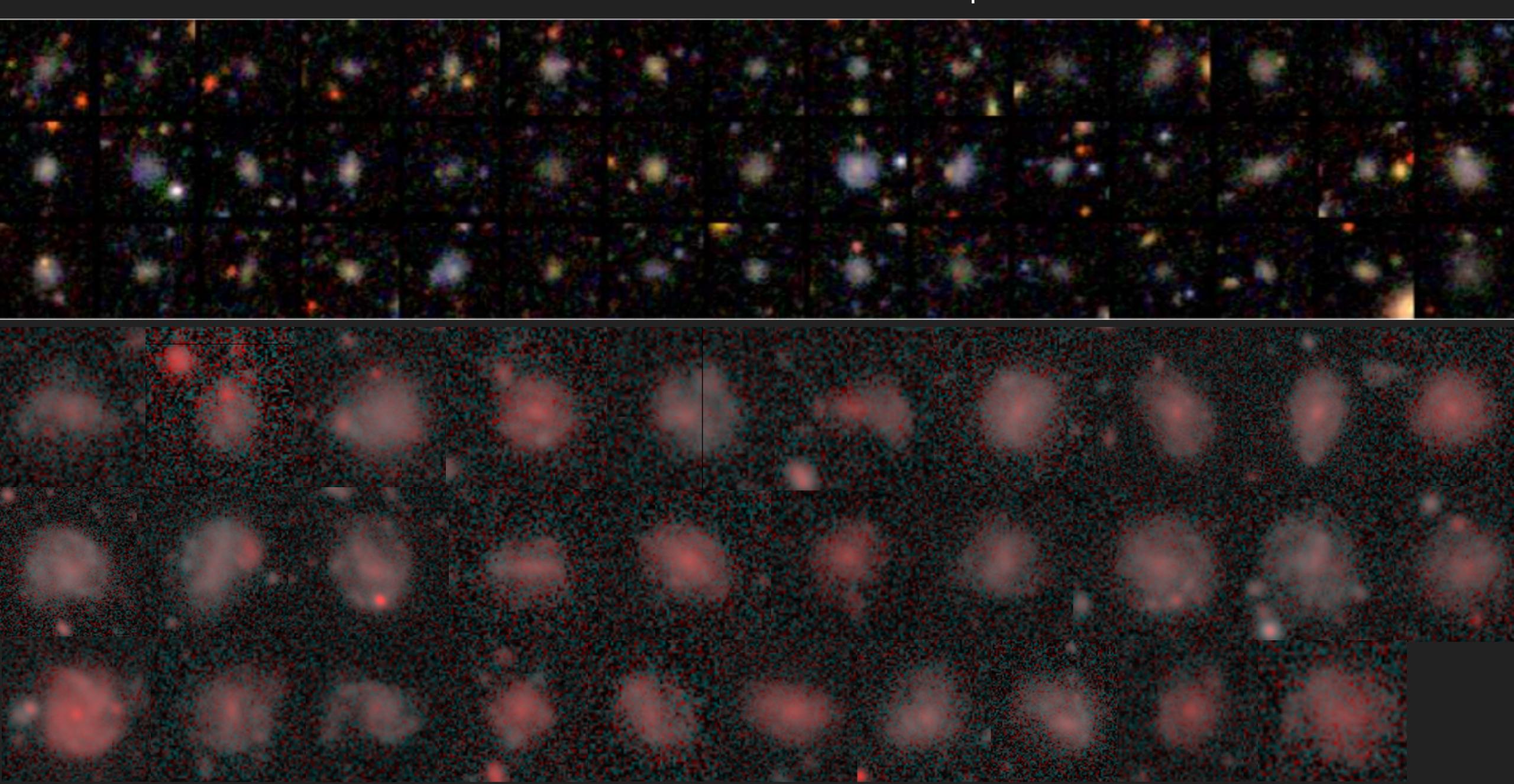


Eigenthaler et al. 2018



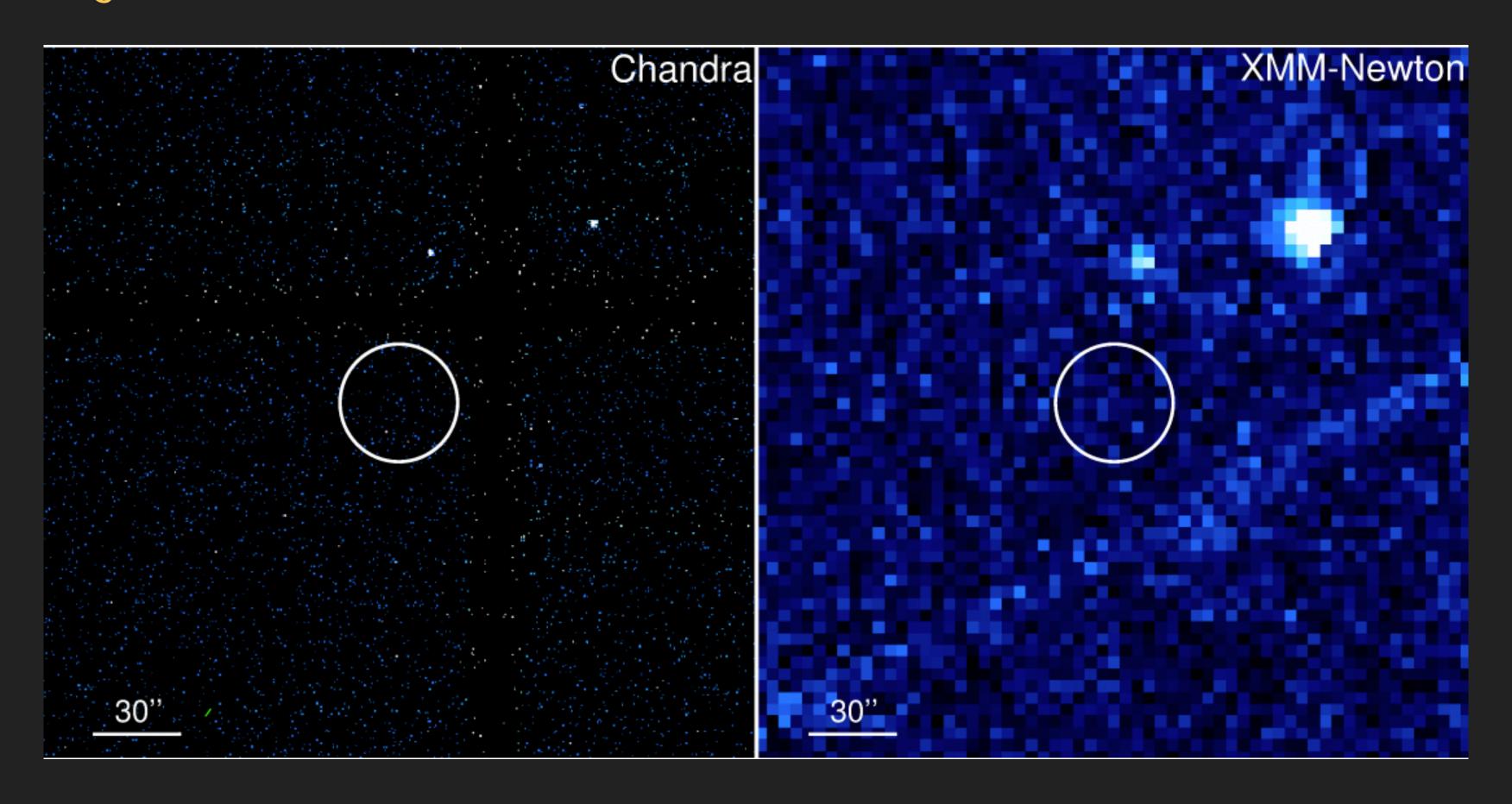
NEGATIVE

There are noises...



Bogdan 2020

DF44 in Coma Cluster



SUMMARY



Hunting Diffuse Dwarfs in the Perseus Clusters

....and several more clusters using Hyper Suprime-Cam

In collaboration with

Nobuhiro Okabe (Hiroshima)

Aaron Romanowsky (SJSU & UCSC)

and, maybe You?

