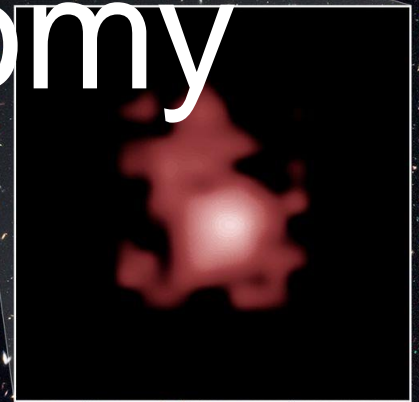


# Detecting First Stars and First Galaxies with IR astronomy

Meng Zhou

2022.4.1

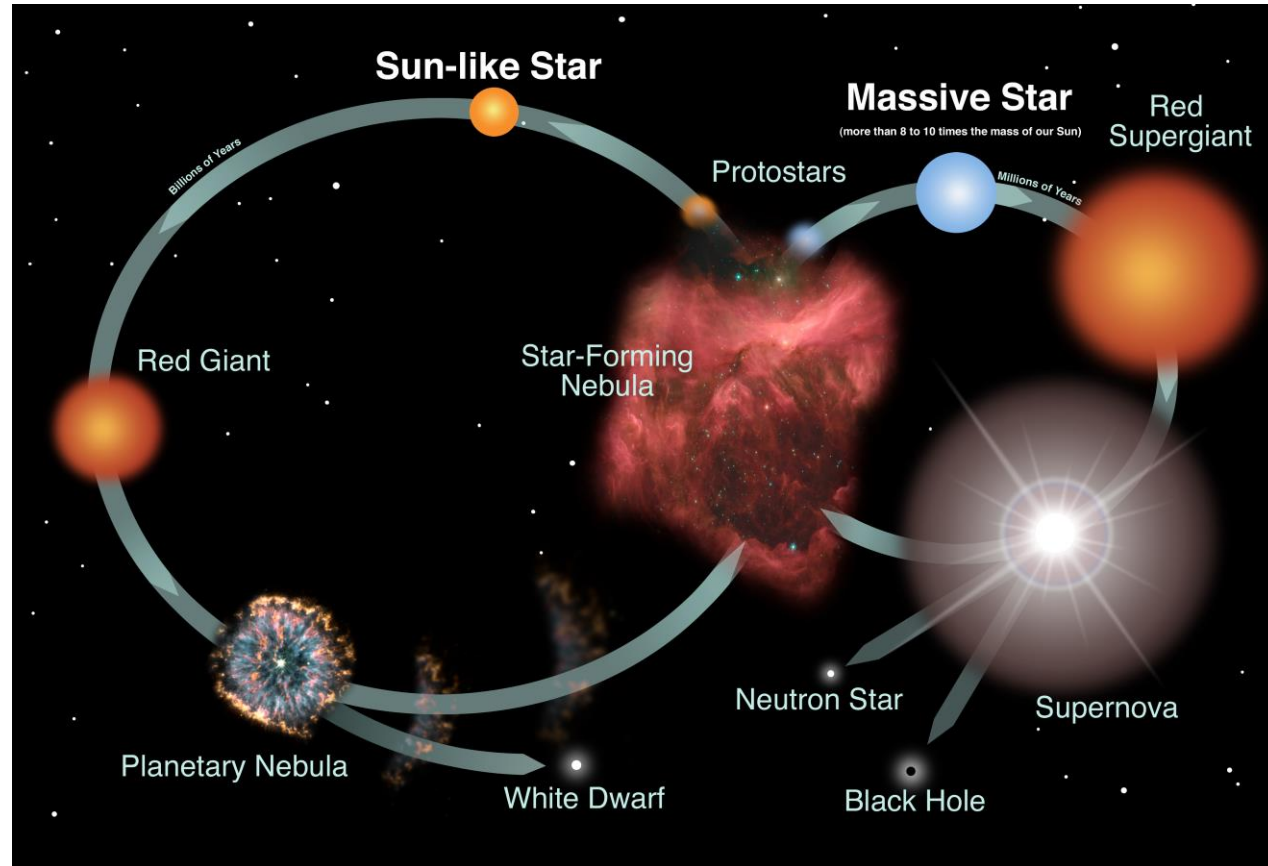
@Student Seminar





# How stars form and cycle?

*Beyond cycle:*  
SN feedback  
Metal enrichment



Credit: NASA and the Night Sky Network

# How galaxies form and evolve?

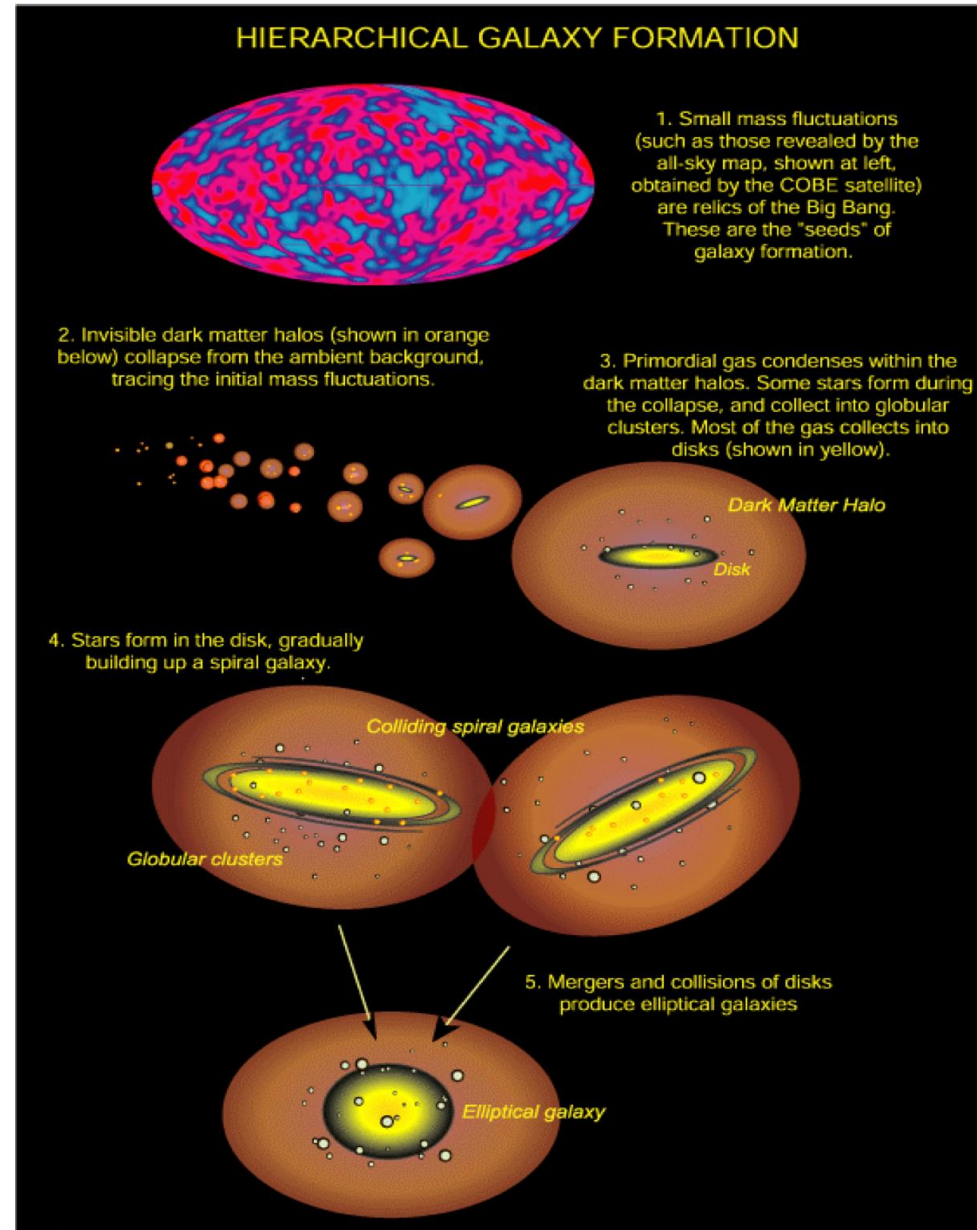
Seeds

**DM halos**

Gas clouds collapse

Stars and Galaxies

Merge and Collision



# Open questions

- When were first stars/galaxies form?
- Stars first, galaxies later? Or coincide?
- Mass function of host DM halos? DM models?
- Environment for first galaxies? Gas rich/poor? Metal rich/poor? SN? Dust? AGNs?

Only Observations can answer!

Lyman  $\alpha$  emitter

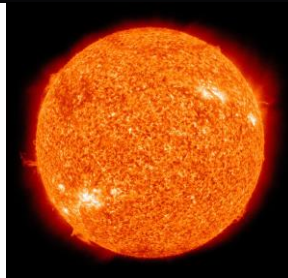
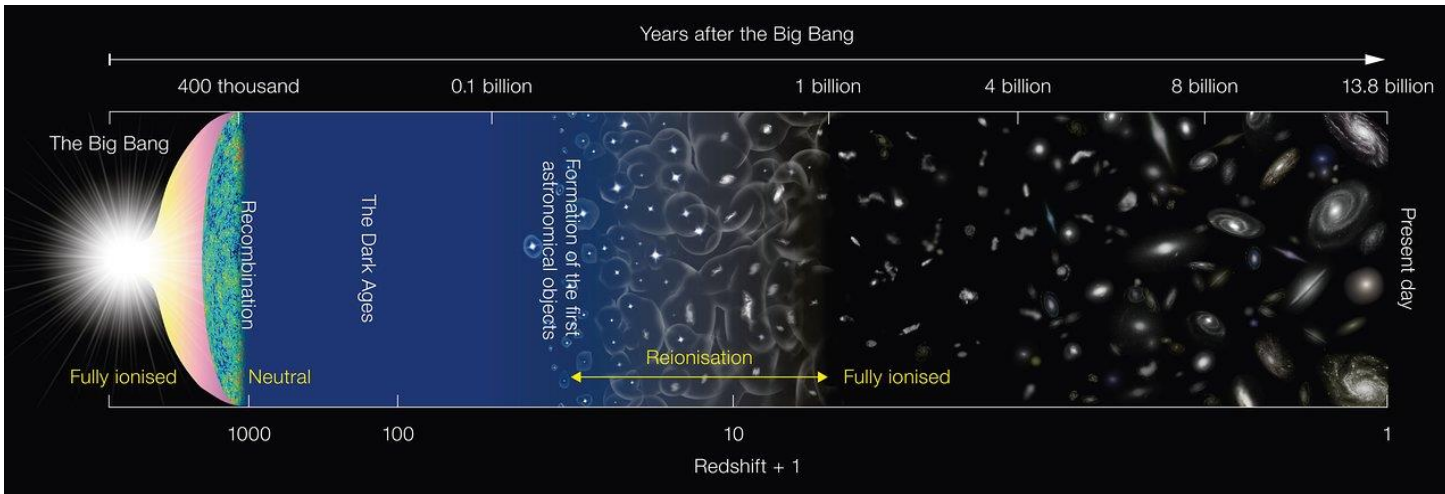
Lyman break

GRB

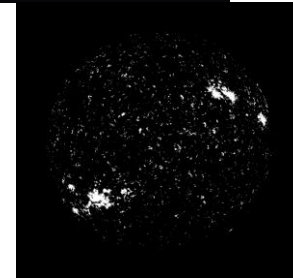
H $\alpha$

...

# Let's do some calculation first.



"Suns"  $z=20$



"Suns"  $z=0$

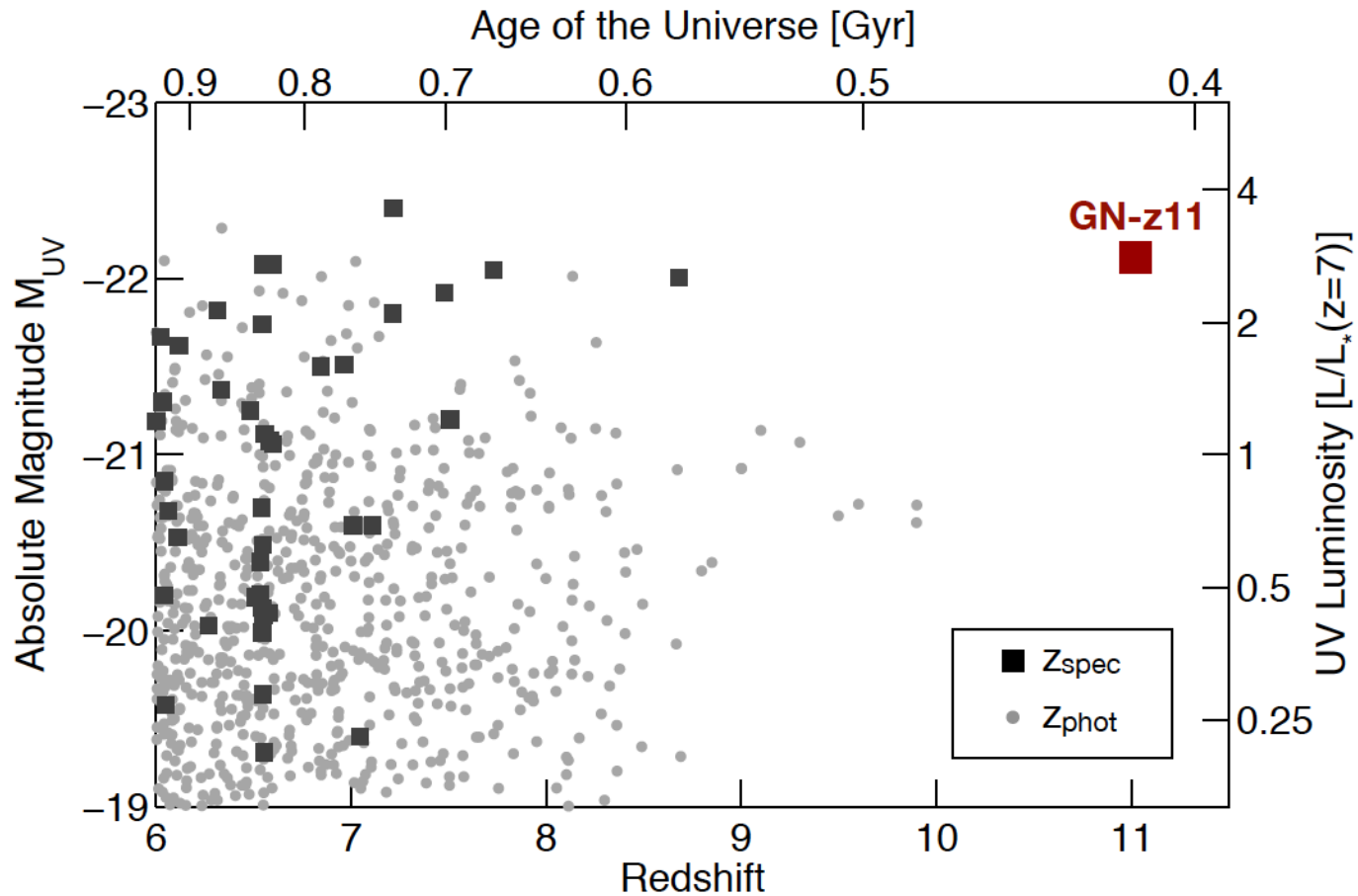
$$\lambda_{obs} = \lambda_{rest} \frac{1 + z_{emit}}{1 + z_{obs}}$$

$$\lambda_{rest} = 580\text{nm}$$

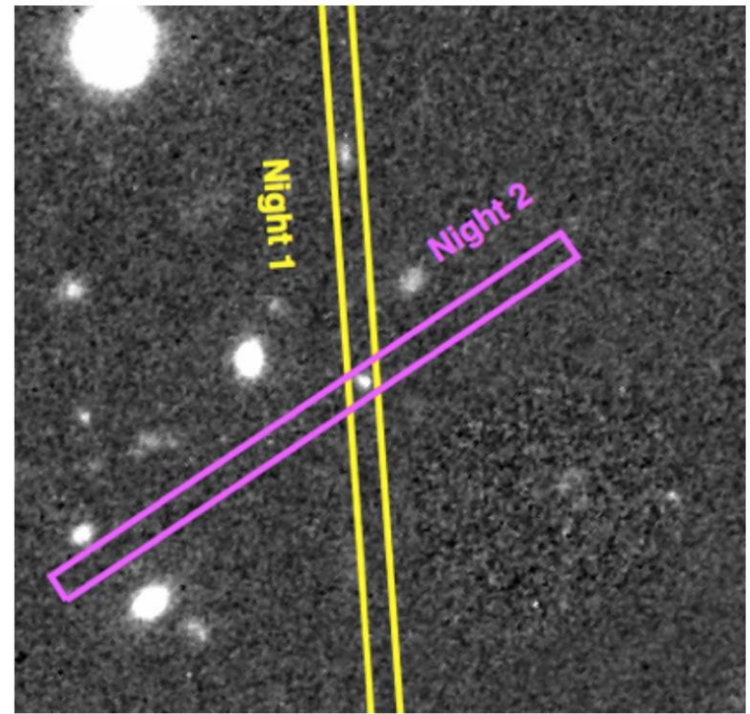
$$\lambda_{obs} = 12.18\mu\text{m}$$

**We need Infrared observations!**

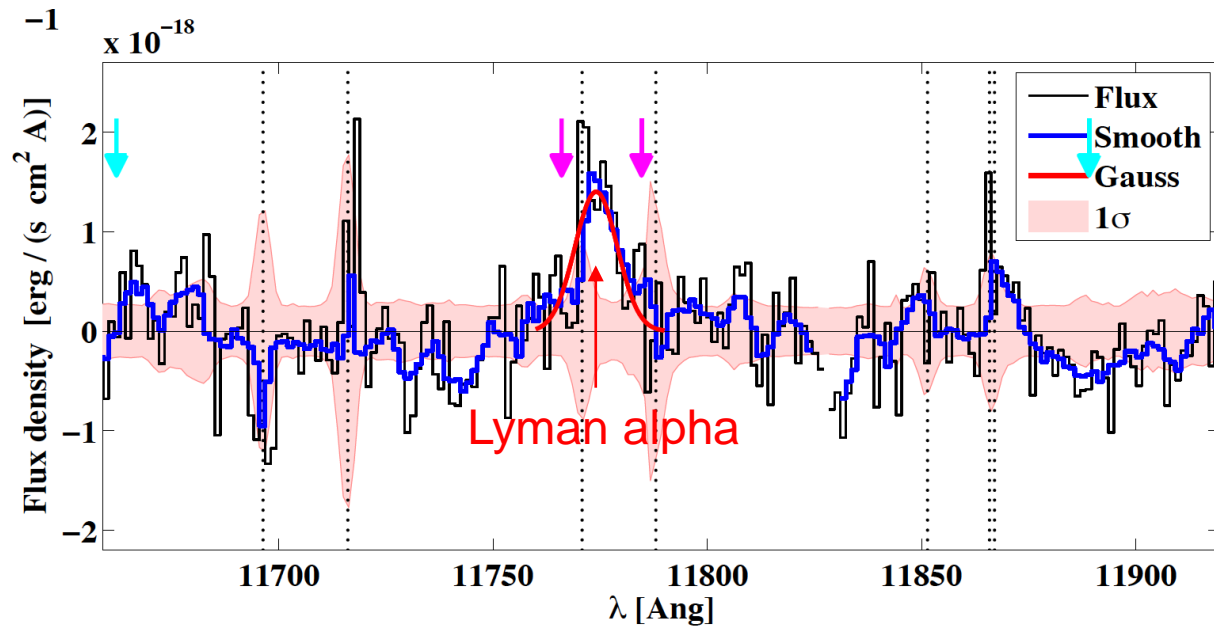
# Current observations cannot provide enough samples!



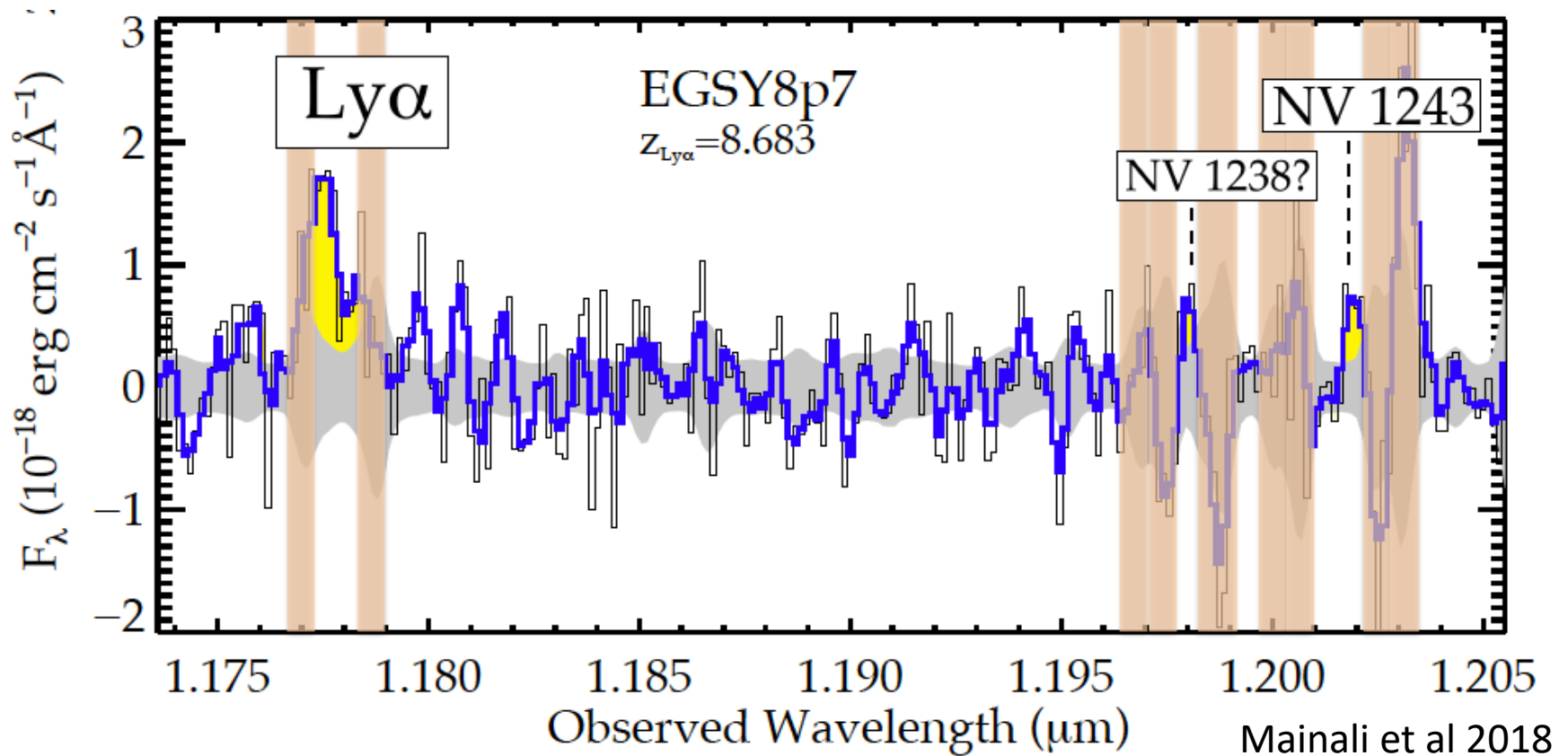
# EGSY8p7



- Keck/MOSFIRE, July 2015
- $z=8.68$
- $m_{AB} = 25.26$ ,  $M \sim 10^{10} M_{\odot}$



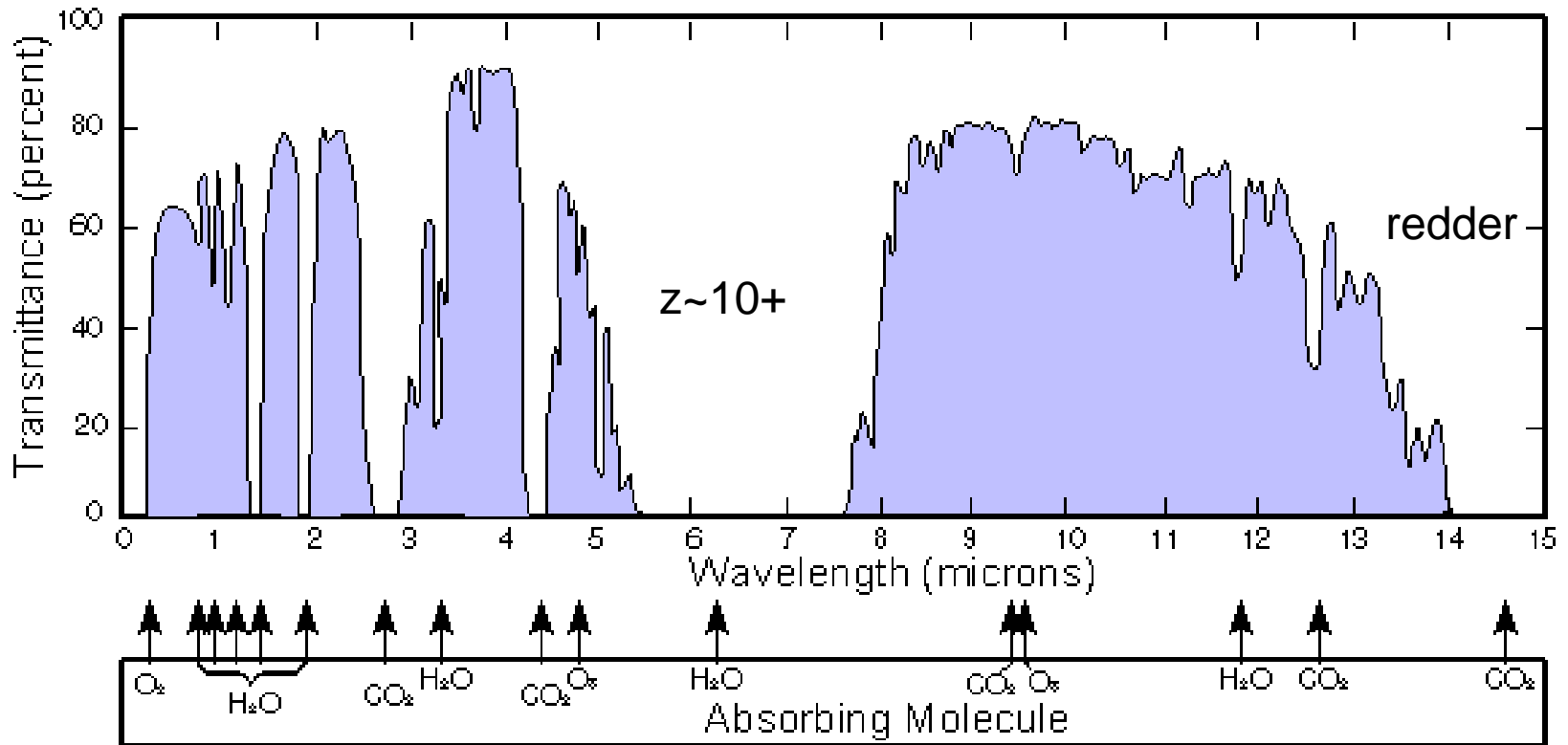
# EGSY8p7



The detection of NV emission indicated AGN activities or fast radiative shocks.

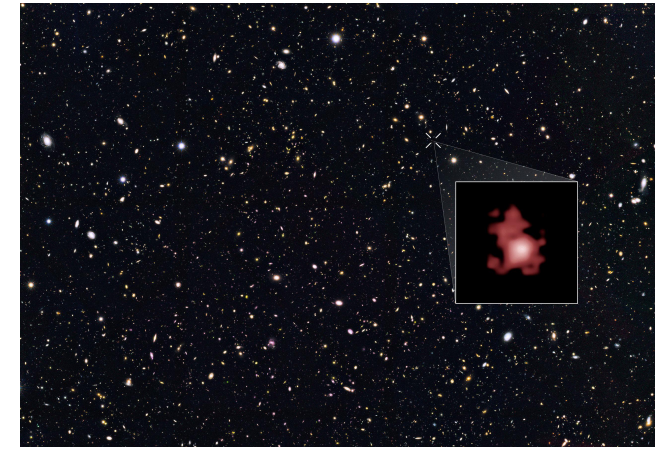
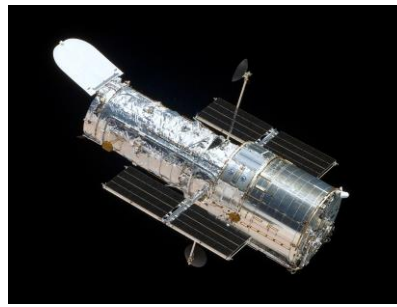


# IR window

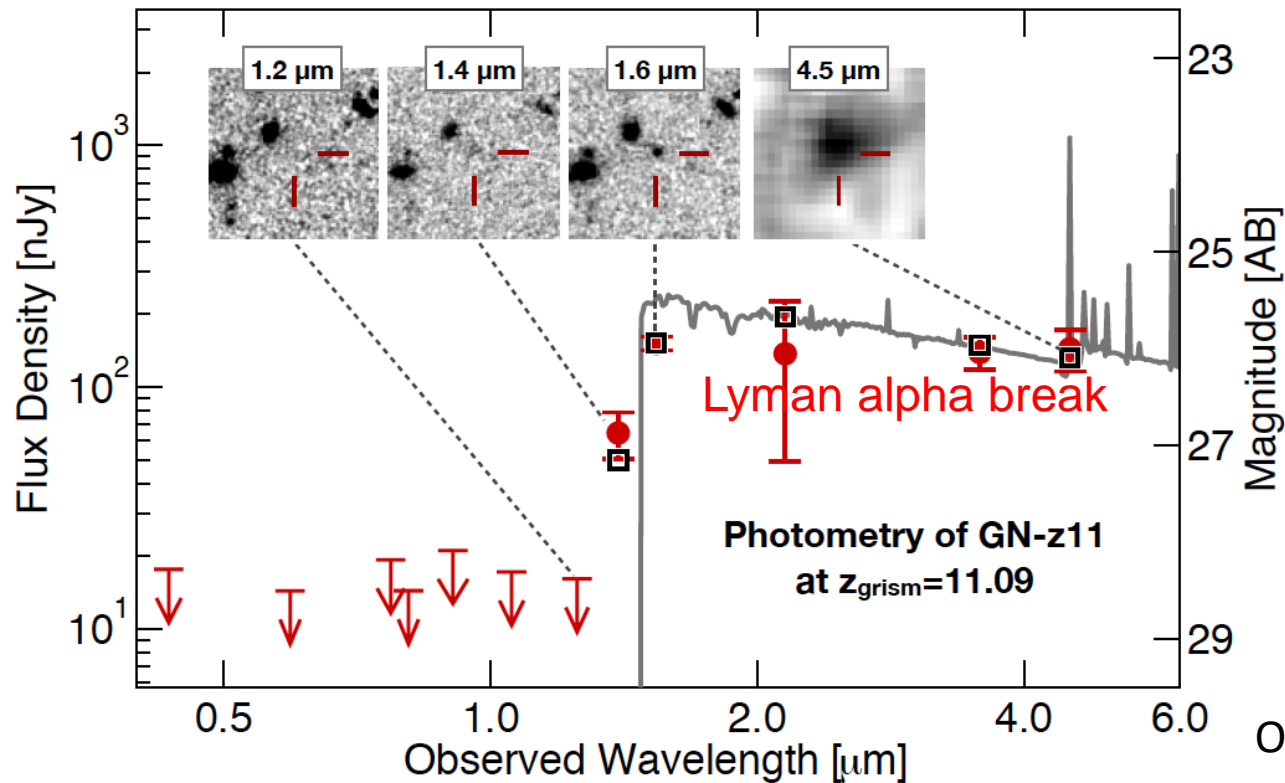


If we want to see some “old” first galaxies, we need space telescopes!

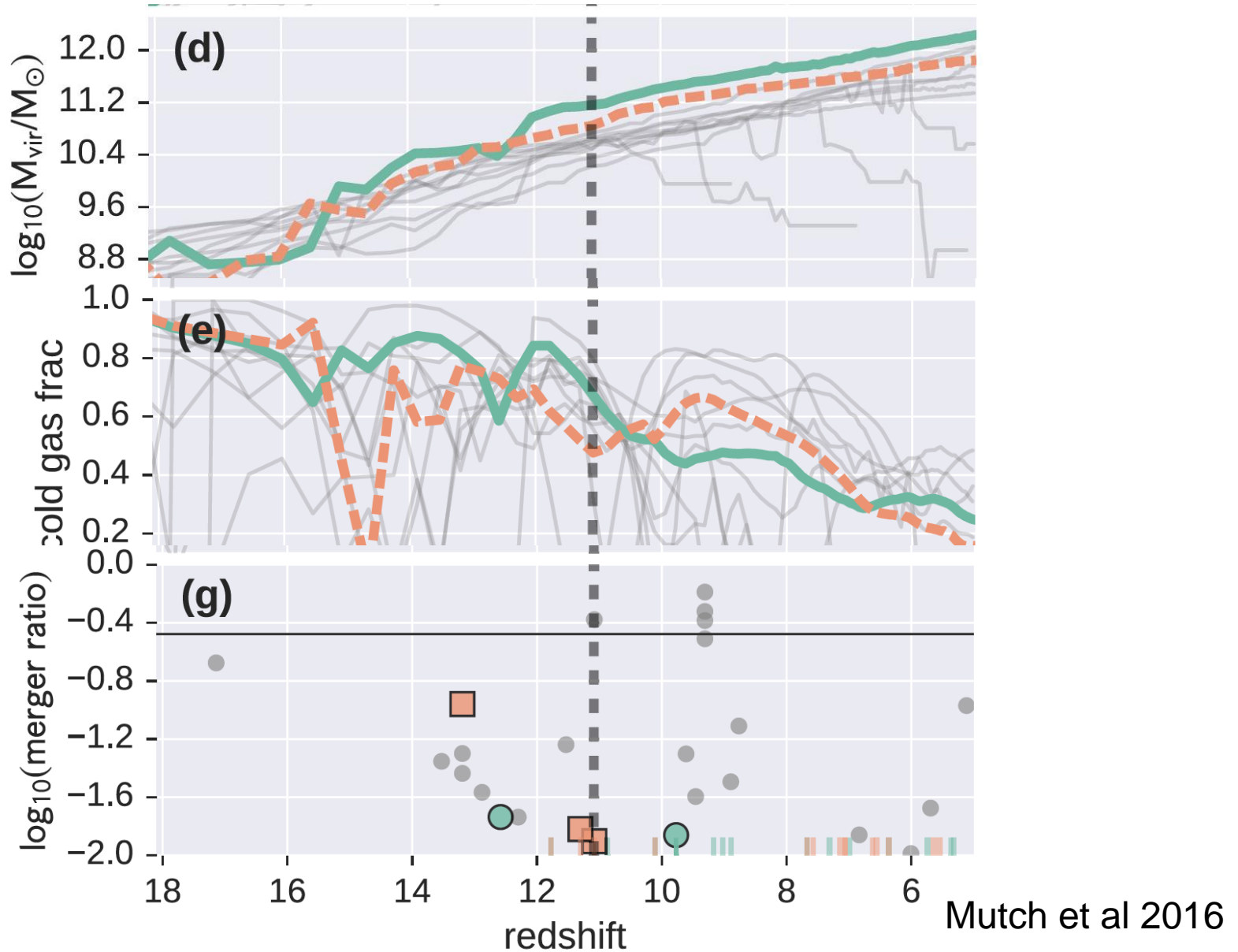
# GN-z11



- Hubble WFC3/IR, March 2016
- $z = 11.1$
- $m_{AB} = 26.0$ ,  $M \sim 10^9 M_{\odot}$










GN-z11 might be formed in a growing, gas-rich and isolated DM halo.



# Debate about GN-z11 flash

Original paper suggests a UV flash associated with GRB

## A possible bright ultraviolet flash from a galaxy at redshift $z \approx 11$

Linhua Jiang <sup>1,2</sup> , Shu Wang<sup>1,2</sup>, Bing Zhang <sup>3</sup> , Nobunari Kashikawa <sup>4,5</sup>, Luis C. Ho<sup>1,2</sup>, Zheng Cai<sup>6</sup>, Eiichi Egami<sup>7</sup>, Gregory Walth<sup>8</sup>, Yi-Si Yang <sup>9,10</sup>, Bin-Bin Zhang <sup>9,10</sup> and Hai-Bin Zhao<sup>11,12</sup>

Supporters

GN-z11-flash in the context of Gamma-Ray Burst Afterglows

D. A. Kann,<sup>1</sup> M. Blazek,<sup>1</sup> A. de Ugarte Postigo,<sup>1,2</sup> and C. C. Thöne<sup>1</sup>






A solar system object?

## A more probable explanation for a continuum flash towards a redshift $\approx 11$ galaxy

Charles Louis Steinhardt <sup>1,2</sup> , Michael I. Andersen <sup>1,2</sup>, Gabriel B. Brammer <sup>1,2</sup>, Lise Christensen <sup>1,2</sup>, Johan P. U. Fynbo <sup>1,2</sup>, Peter Laursen <sup>1,2</sup>, Bo Milvang-Jensen <sup>1,2</sup>, Pascal A. Oesch <sup>1,2,3</sup> and Sune Toft <sup>1,2</sup>

a Russian Proton rocket?

## GN-z11-flash from a man-made satellite not a gamma-ray burst at redshift 11

Michał Jerzy Michałowski <sup>1</sup> , Krzysztof Kamiński <sup>1</sup> , Monika Katarzyna Kamińska  and Edwin Wnuk

A Satellite Glint?

The GN-z11-Flash Event can be a Satellite Glint

Guy Nir<sup>1</sup> , Eran O. Ofek<sup>1</sup> , and Avishay Gal-Yam<sup>1</sup> 

SNe?

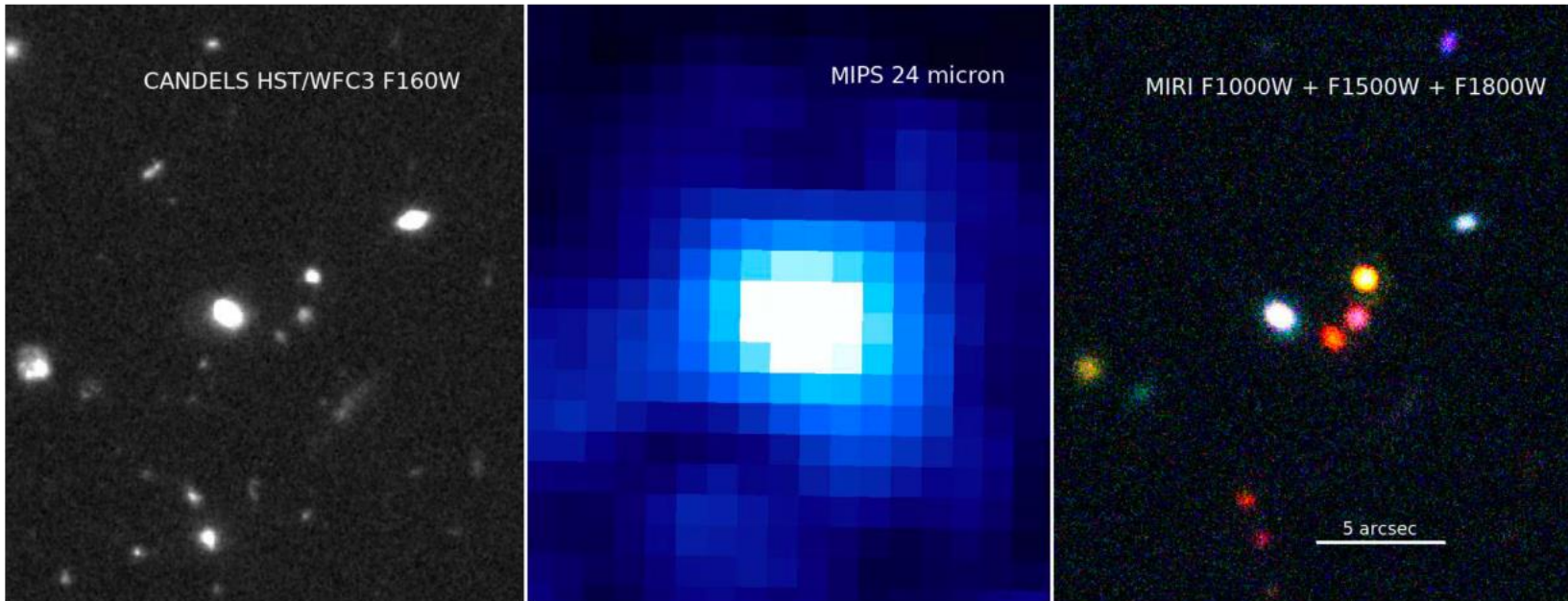
## Signatures of population III supernovae at Cosmic Dawn: the case of GN-z11-flash

Hamsa Padmanabhan<sup>1</sup>  · Abraham Loeb<sup>2</sup>



# Future: JWST

@Chen & Zhang's talk



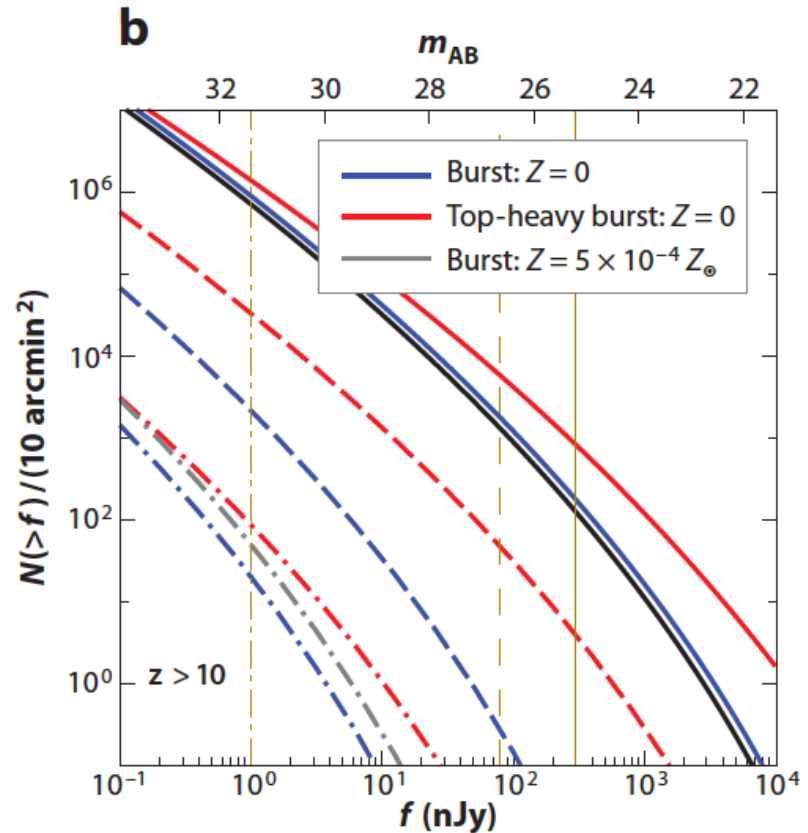
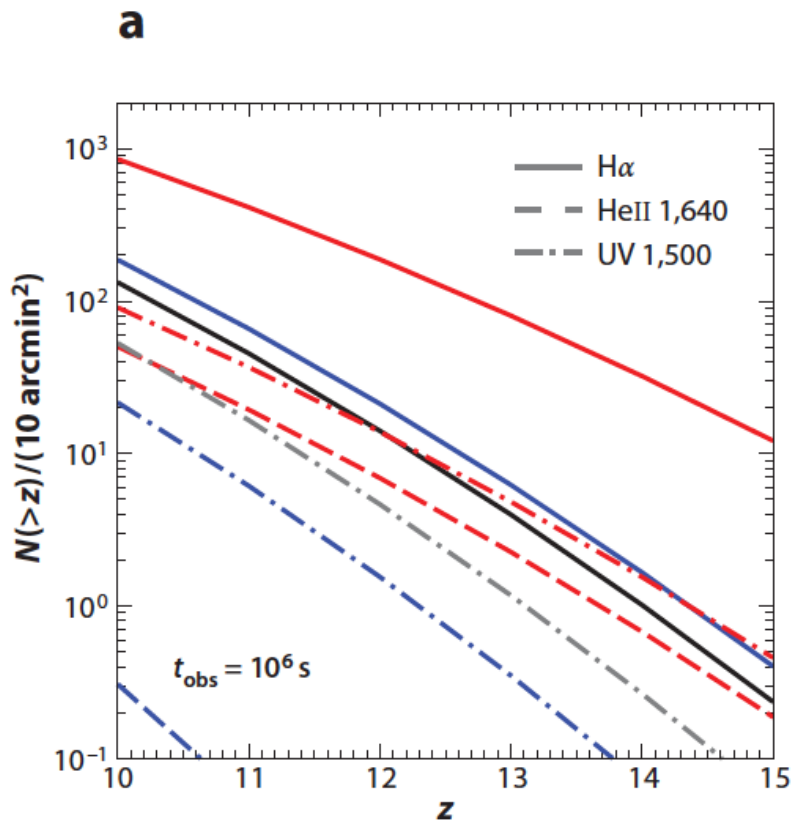
HST/WFC3

Spitzer/MIPS

JWST/MIRI (simulated)

Jason 2018

# JWST can detect 10~1000 star-forming galaxies with $z > 10$ .



Pawlik et al 2011

We need other detections as IR cannot provide complete samples by itself.

- 21cm intensity mapping
- Thomson scattering optical depth/Global 21cm signal
- Early BHs
- Local ultrafaint dwarf galaxies
- ...

# Take home messages

- First stars and galaxies are missing from current formation history.
- IR astronomy has made numerous efforts to study their formations and environments, and will make more in the future.
- We need other detections.

