LCO:Las Cumbres Observatory



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Outline

1. Telescopes and sites of LCO

2. Power of LCO global network

3. Science and LCO Key projects

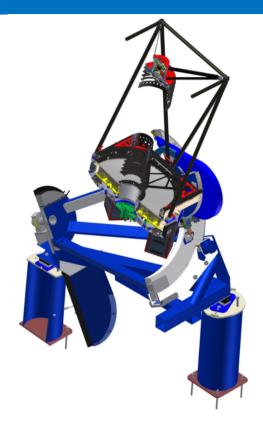
4. How to apply LCO time

Telescopes and sites

- twenty-two telescopes at eight (seven) sites around the world.
- 10 1-meter telescopes
- 2 2-meter telescopes
- 10 0.4-meter telescopes

1-meter telescope

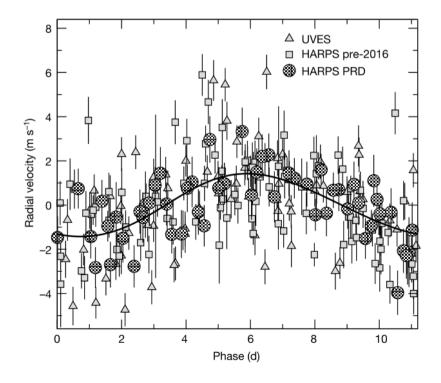
Optical: FOV: 26x26 arcmin Pixel scale: 0.39 arcsec CCD: 4Kx4K pixel Cycle time (overhead+readout): 42s Filter wheel options: 21s Airmass: <3.8 (15degree horizon limit) **Precision:** ~0.3% for V=11mag (single point) **Filter: Johnson/Cousins UBVRI Sloan primed ugri** PanSTARRS z,y

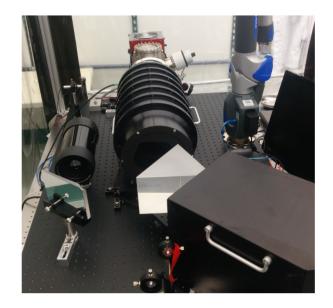


1-meter telescope

Robotic Echelle Spectrographs (NRES) primary motivation: measure RV of planet system. Number: six telescopes at different sites. high-resolution: R~53,000

Precision: 3 m/s for V=12mag



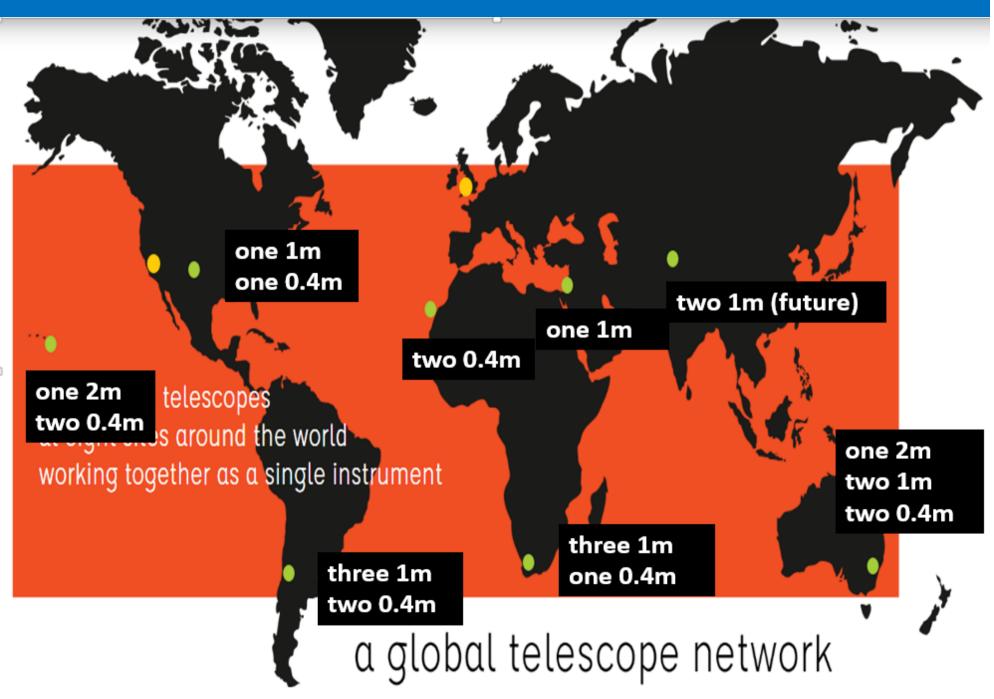


0.4-meter telescope

Optical: FOV: 29x29 arcmin **Pixel scale: 1.14 arcsec** CCD: 2Kx3K pixel Cycle time (overhead+readout): 32s Filter wheel options: 9s Filter: Johnson/Cousins BV **Sloan primed ugriz PanSTARRS** w

Purpose: science observations and education

Sites



Motivations

- Theory appears to predict fewer freefloating planets than observations (Ma, Mao et al. 2016)
- needs to measure the population empirically (e.g. with K2C9)
 - Applied for CFHT time for (Penny, Zhu, Fouque, Dong + SM)

•Gradually building up the microlensing team

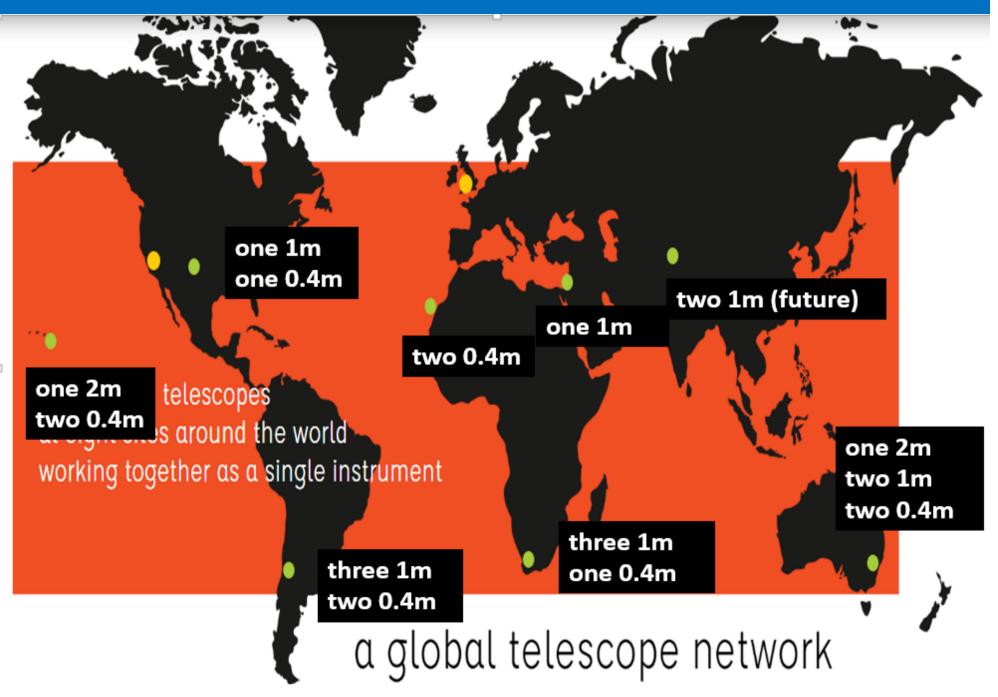
- Joined RoboNET through LCO
- two 1m telescopes to be built in Tibet: \$4m
- Needs to build up observational and theoretical expertise



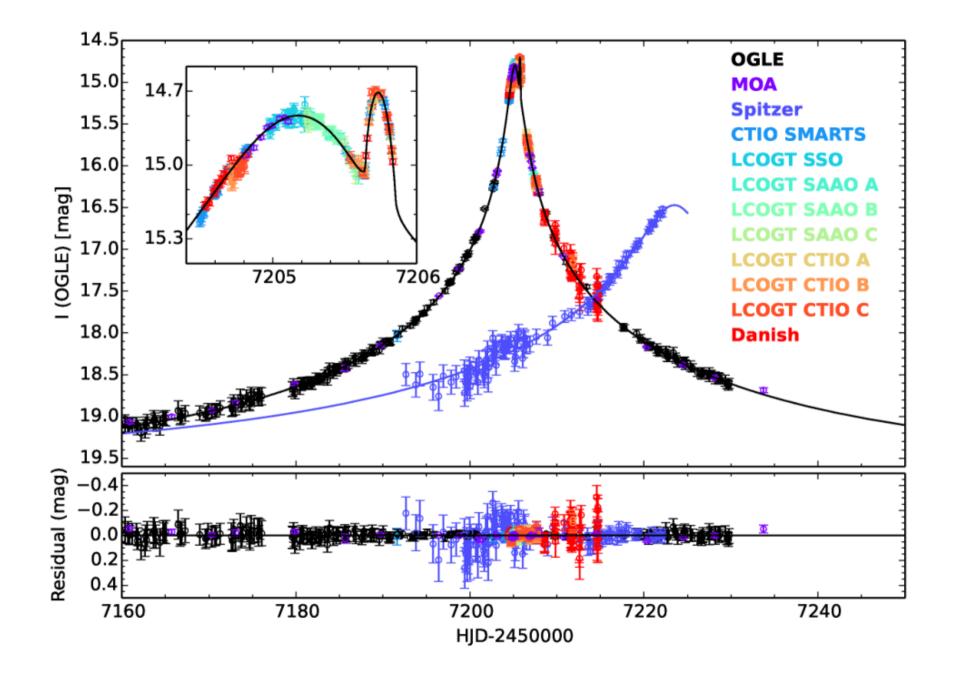


2019: ~2500 hours

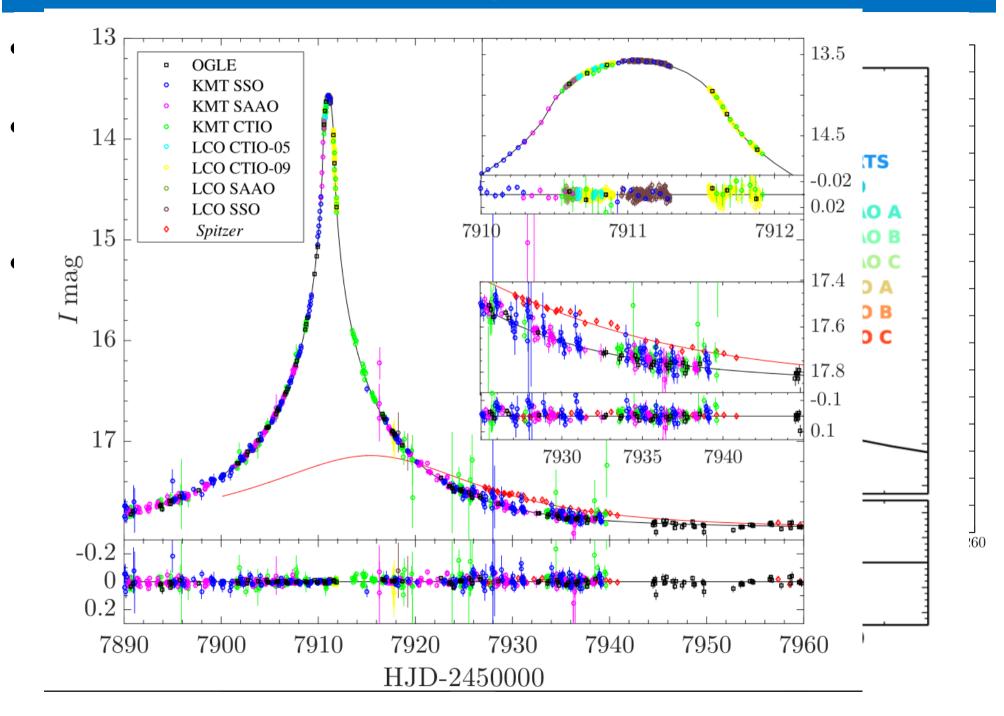
Power of LCO



Global Network



Flexible (high) Cadence



Flexible exposure time

Observatory Tools

Saturated at I ~ 13mag



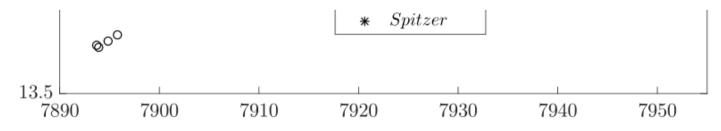
need on our different classes of telescope and using different filter sets.

Exposure Time Calculator

Target Visibility Calculator

When planning your observing request it is import to make sure your targets will be visible, from which sites and during which time periods. The Visibility Calculator will show you all of this.

We have provided a tool to help you decide what exposure time your target will



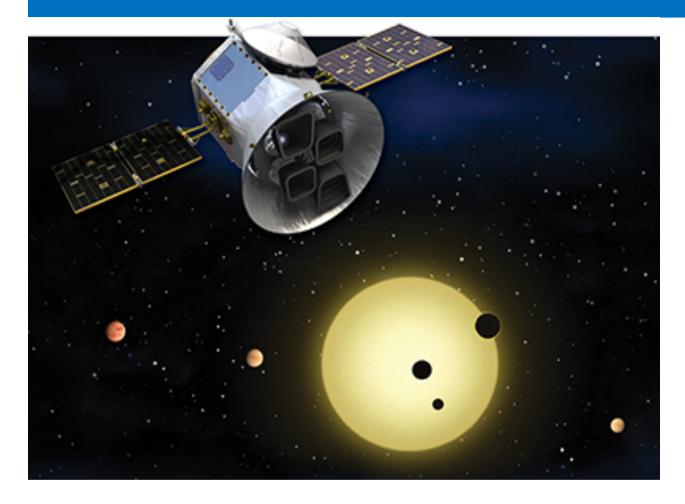
Science of LCO

- •The areas of astronomy most interesting to us are those which take full advantage of this sort of network.
- •Targets which appear suddenly and without warning like supernovae and asteroids.
- •Objects which need to be observed for long periods in darkness like exoplanets and binary star systems.

Science Collaboration Key Projects

PI name	PI Institution	Contributing Institutions	Title	Semesters	2018A hours				
					FLOYDS (2m)	Spectral (2m)	NRES (1m)	Sinistro (1m)	SBIG (0.4m)
Brown, T., Mazeh, T.	LCO, U. Tel Aviv	LCO, SUPA	Using NRES to Validate and Characterize Exoplanets Found by TESS and Other Surveys	2017AB- 2020A			2150		
Horne, K.	St. Andrews U.	SUPA, LCO, ANU, SAAO	Echo Mapping of AGN Accretion Flows	2014A- 2018A		140		975	
Howell, A.	LCO	LCO, NAOC, NOAO, ANU, TAU, UTexas	The Global Supernova Project	2017AB- 2020A	274	250		1109	
Robertson, P.	Penn State Univ.	LCO	High-Cadence Monitoring of the Sun's Coolest Neighbors	2017AB- 2020A				100	1400
Shporer, A.	Caltech	LCO, ANU, SUPA, SAAO, IAC, UTexas, UHawaii	Transiting Exoplanet Science with LCO - The Network Awakens	2017AB- 2019A		65	301	300	339
Tsapras, Y.	U. Heidelberg	SUPA, LCO, NAOC	ROME/REA - A three-color window to planets beyond the snow-line	2017AB- 2020A				740	

TESS: Transit Exoplanet Survey Satellite



- TESS was launched yesterday
- Will discover more than 5000 nearby exoplanets

- Two problems:
- 1. Pixel Scale of TESS is 0.7 arcmin/pixel. Highly blended
- 2. Measure the Radius of the components, not the mass

LCO-TESS Key Project

Key Project 1: Transiting Exoplanet Science with LCO - The Network Awakens

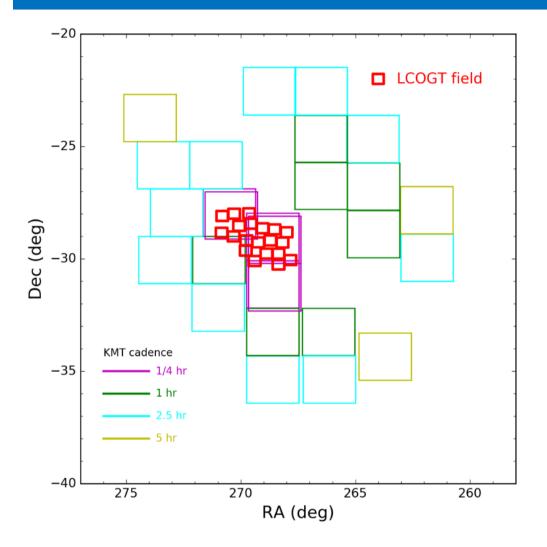
Find the host star of the planet system.

Key Project 2: Using NRES to Validate and Characterize Exoplanets Found by TESS

Measure Radius Velocity of Host star, in order to answer

- 1. The exoplanet mass/period distribution.
- 2. The inclination

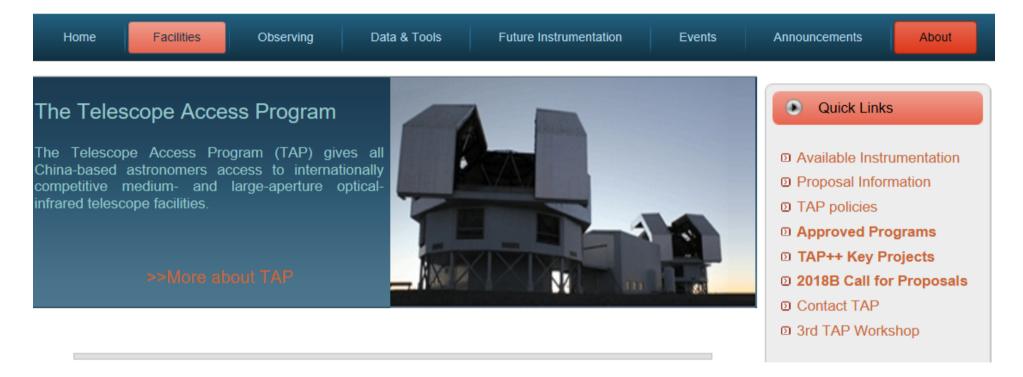
LCO Microlensing Key Project



- FOV of LCO: 26x26 arcmin
- FOV of KMTNet: 2x2 degree
- ROME: 3 exposures per 8hrs
- **REA:** < 30min cadence
- LCO Network is suitable for Follow up, but not for wide-field survey!

How to apply LCO time

Telescope Access Program



- Each semester China has 200 hours 1-m network
- 2017B: 3 proposals; 2018A: 4 proposals

Trade among aperture sizes

- Todd's official rate: 4 (0.4m) = 2 (1m) = 1 (2m).
- Shude's rate: 1 (1m) = 6.25 (0.4m) by aperture size
- Todd's personal rate: The 0.4m's are usually **undersubscribed**, so, while I give the official rate above, let me know if you need more than that.

Spitzer Microlensing project:
2017: 200 hours 1-m
2018: 120hr (0.4m) from China 100hr (0.4m) from IPAC 90hr (1m) from Wise.

Purchase network observing time

- •1-meter: 3000 hours
- •2-meter: 400 hours
- •0.4-meter: enough
- \$350/hr + \$3,000 (admin fee) for 20-99 hours per year (adr advance)
- \$300/hr for 100-499 hours per year (admin fee waived)
- \$250/hr for 500 or more hours per year (admin fee waived)



• LCO is a Global Telescope Network and can conduct 24hours observations.

•LCO Global Network has flexible cadence and exposure time.

• 1-meter telescope is the main scientific telescopes of LCO and has two mode: Optical and NRES.