

The background of the slide is a light gray gradient with several realistic water droplets of various sizes scattered across it. The droplets have highlights and shadows, giving them a three-dimensional appearance. The largest droplet is in the bottom right corner, while others are smaller and more numerous in the top left and bottom center areas.

# LABORATORY EXPERIMENTS ON SHOCKS

XINYAN HUA

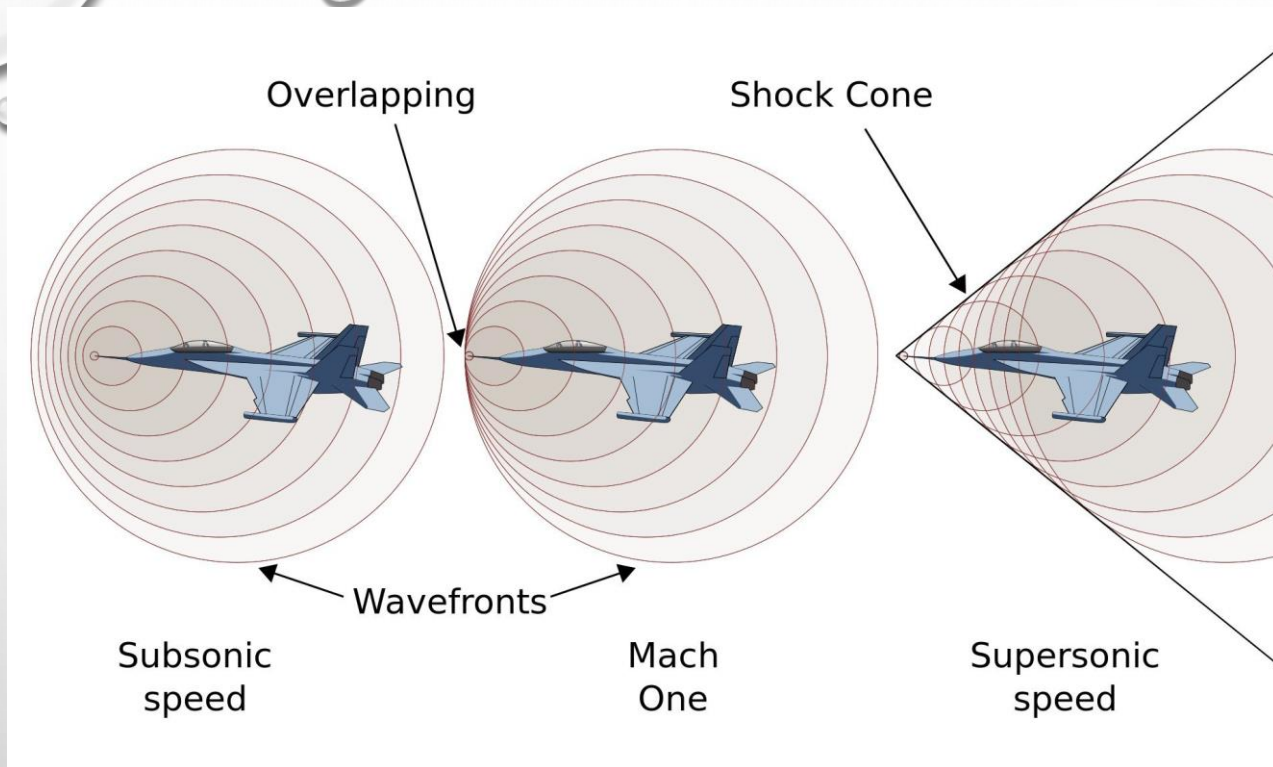
SUPERVISOR: H. FENG; W. CUI

# Outline

- Introduction of shocks
- Why need to study shocks
- Shock experiments in laboratory
- Summary

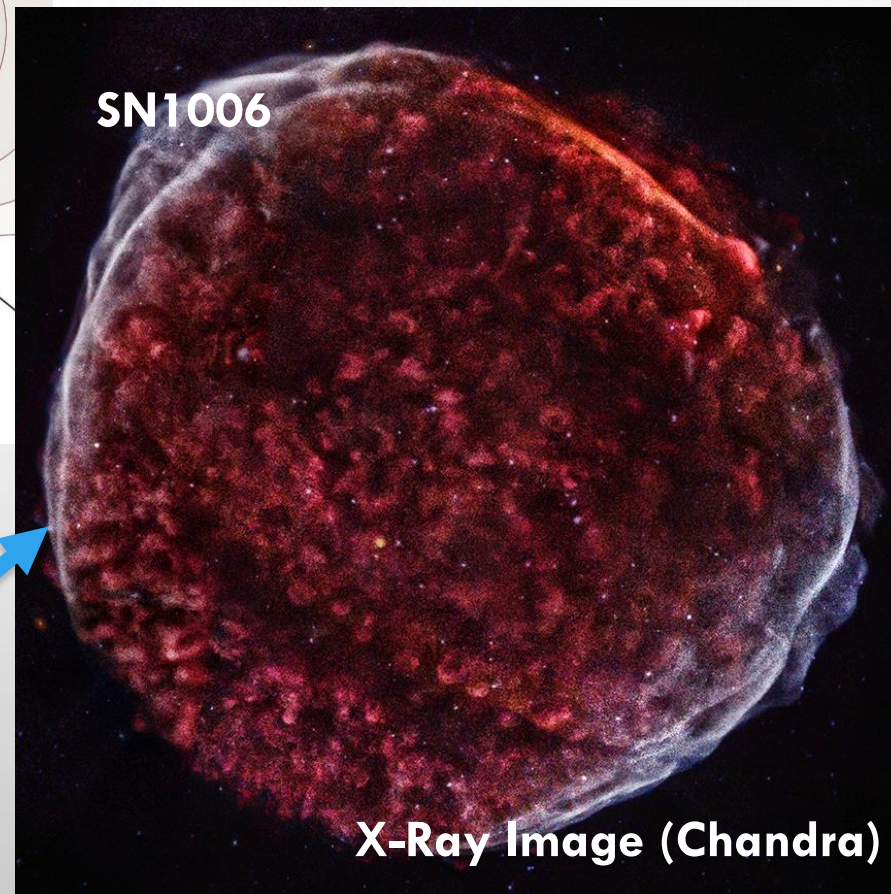
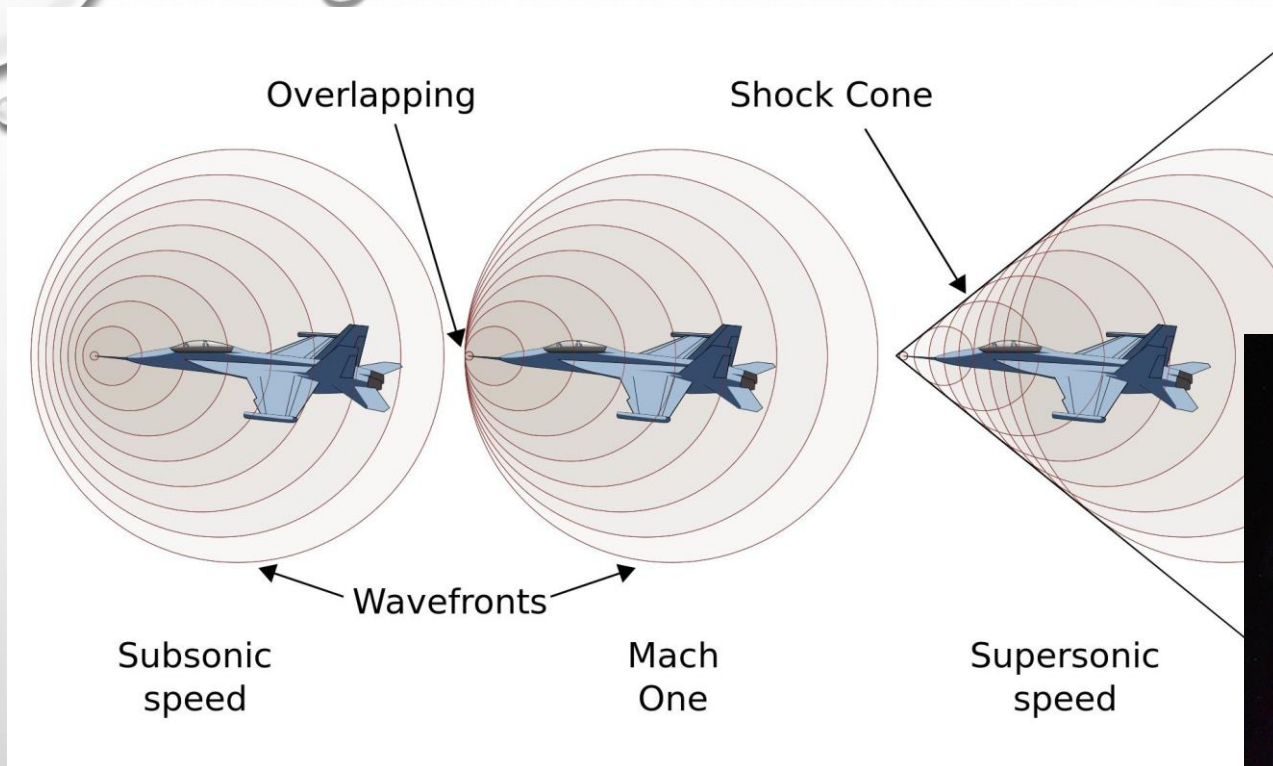
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**Mach number = object velocity / speed of sound**





**Shock front**

$$l_{Coul} = \frac{m_e^2 v^4}{8\pi n Z^2 e^4 \ln \Lambda} \approx 1.4 \times 10^4 \left(\frac{T}{K}\right)^2 \left(\frac{n}{cm^{-3}}\right)^{-1} cm$$

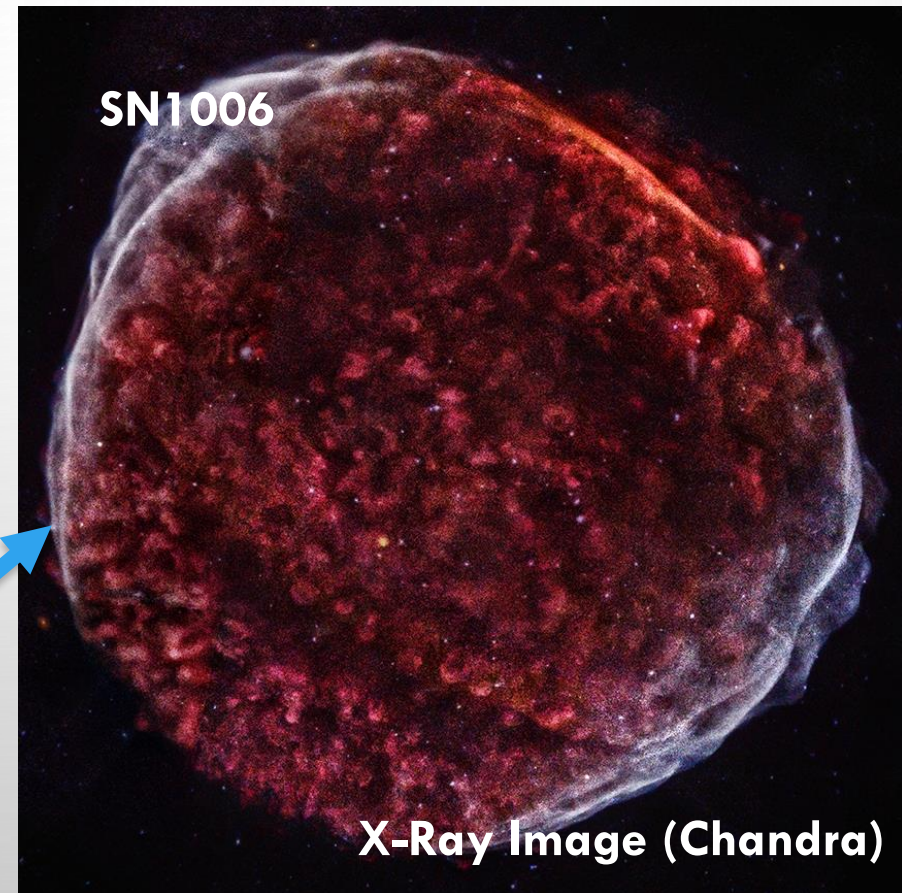
**Coulomb mean free path:  $\sim 10^{20}$  cm**

**Diameter of SN1006:  $\sim 10$  pc =  $3 \times 10^{19}$  cm**

**Shock front width:  $\sim 0.04$ pc =  $1.2 \times 10^{17}$  cm**

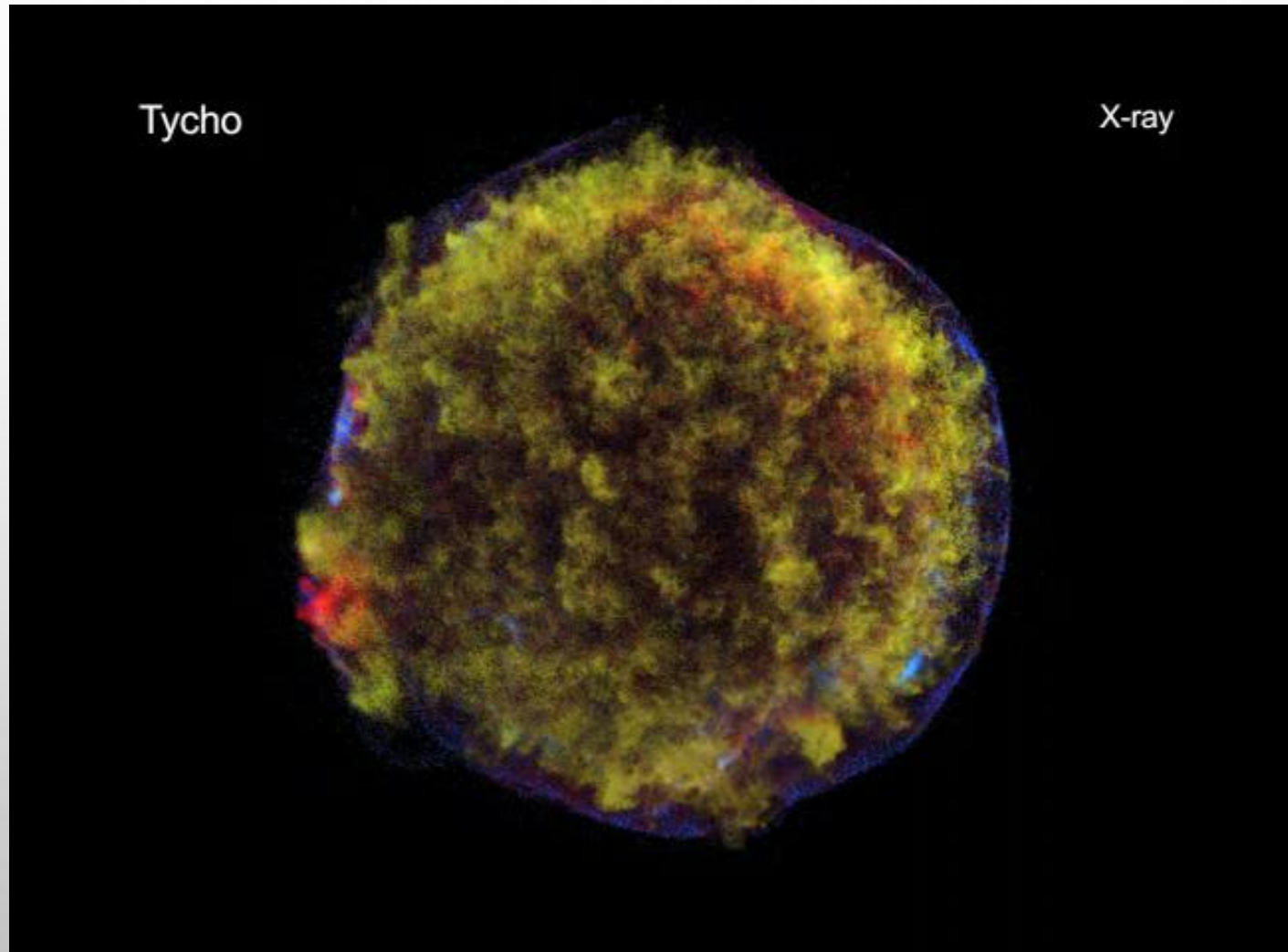
## Collisionless Shocks

Shock front





# Tycho's supernova remnant X-Ray Image (Chandra)



# Tycho

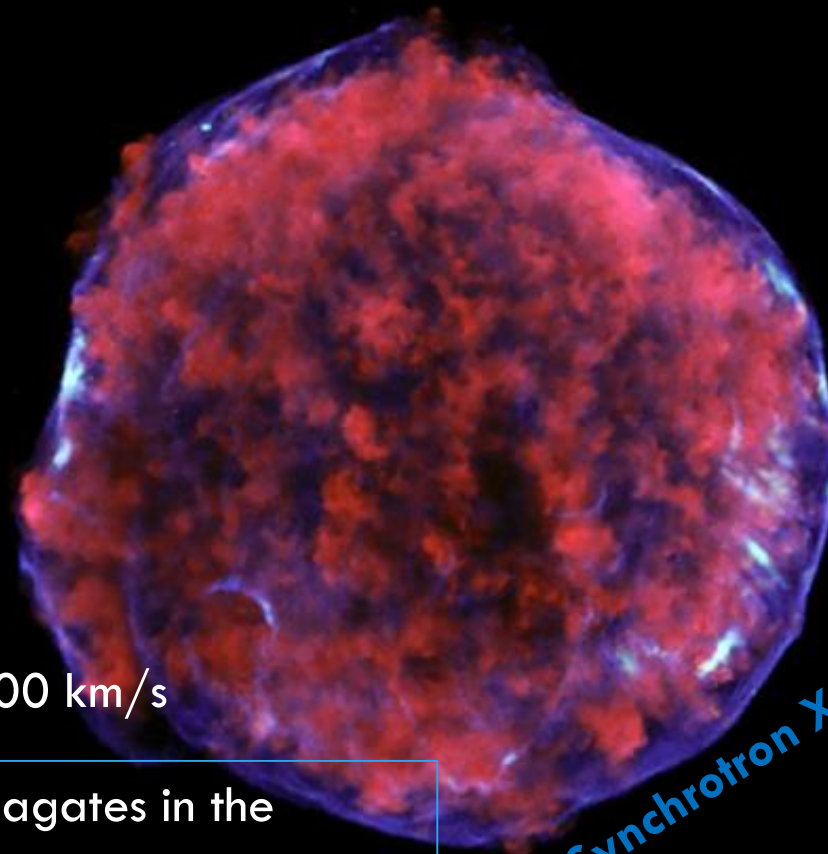
Red: low energy X-Ray  
Blue: high energy X-Ray

Mechanism driving shocks:  
Plasma instabilities

Fermi  
Acceleration

Shock velocity: 4000 km/s

Shock wave propagates in the  
interstellar medium  
TeV electrons producing **synchrotron  
X-Rays**



Synchrotron X-rays



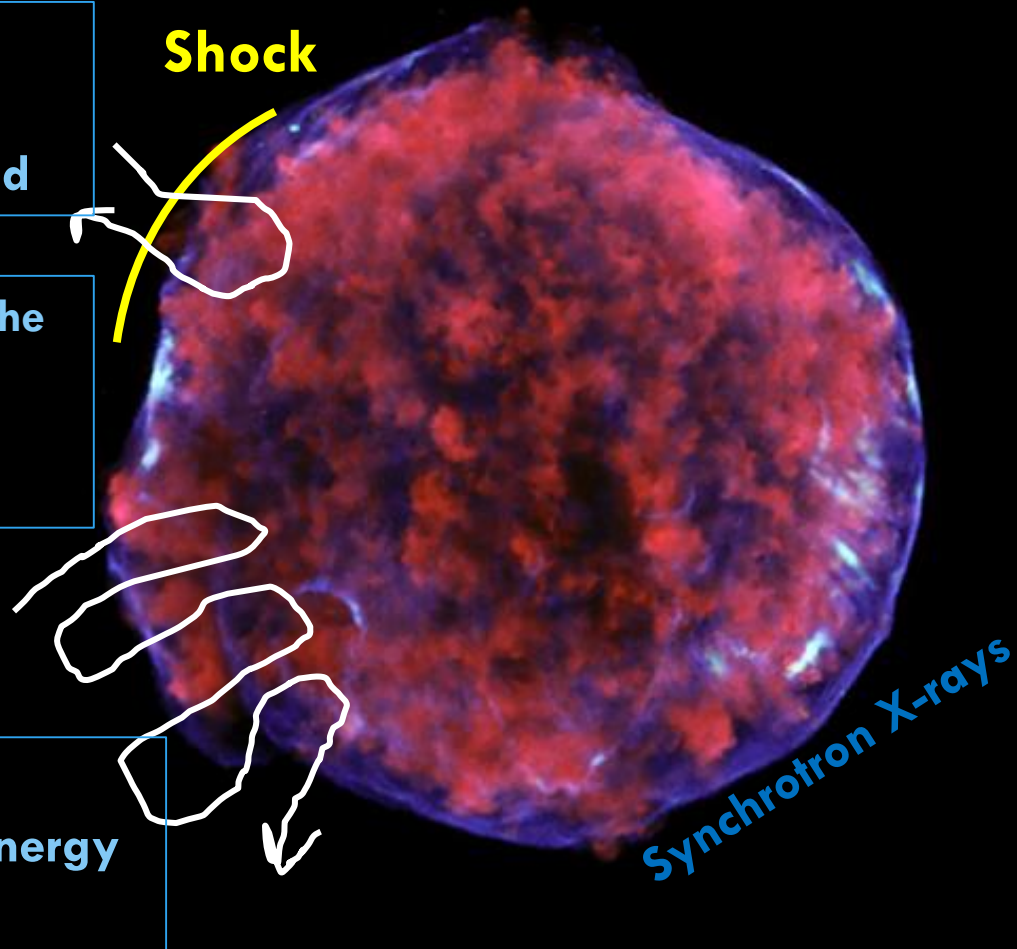
# Tycho

Red: low energy X-Ray  
Blue: high energy X-Ray

**“Random Walk”** of  
charged particles in a  
presence magnetic field

**One round trip** crossing the  
shock wave:  
Gain **1%** of the original  
energy

**1000** round trips:  
**20000** times bigger energy  
eg. **1 GeV** → **20 TeV**



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# Why we need to study (collisionless) shocks ?

- Collisionless shocks are associated with extremely high energy particles.
- Shock waves act as cosmic ray accelerator.
- Shock waves and magnetic field ?
- Condition to generate shocks ?



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# Shock experiments in laboratory

## Laser-produced collisionless shocks

- USA: Livermore (LLNL), Los Alamos (LANL), Rochester (LLE), Berkeley, UT, Michigan
- UK: Didcot (Rutherford lab), Aldermaston (AWE)
- France: CEA (CESTA, Saclay), Ecole Polytechnique (LULI, LOA), CELIA
- Germany: Garching, Jena, Munich, Dresden, Dusseldorf
- Czech Republic: PALS
- Japon: Osaka (ILE), Kyoto (JAERI)
- South Korea: GIST
- Russia: Moscou (Lebedev, MEPHI), Arsamas
- China: Pekin, Shanghai



## Shock experiments in laboratory

### Plasma diagnostics in high power laser experiments

**Velocities:** particle diagnostics (spectrometers, Thomson parabola), Doppler

**Magnetic fields:** proton radiography, polarimetry, faraday rotation, B dots

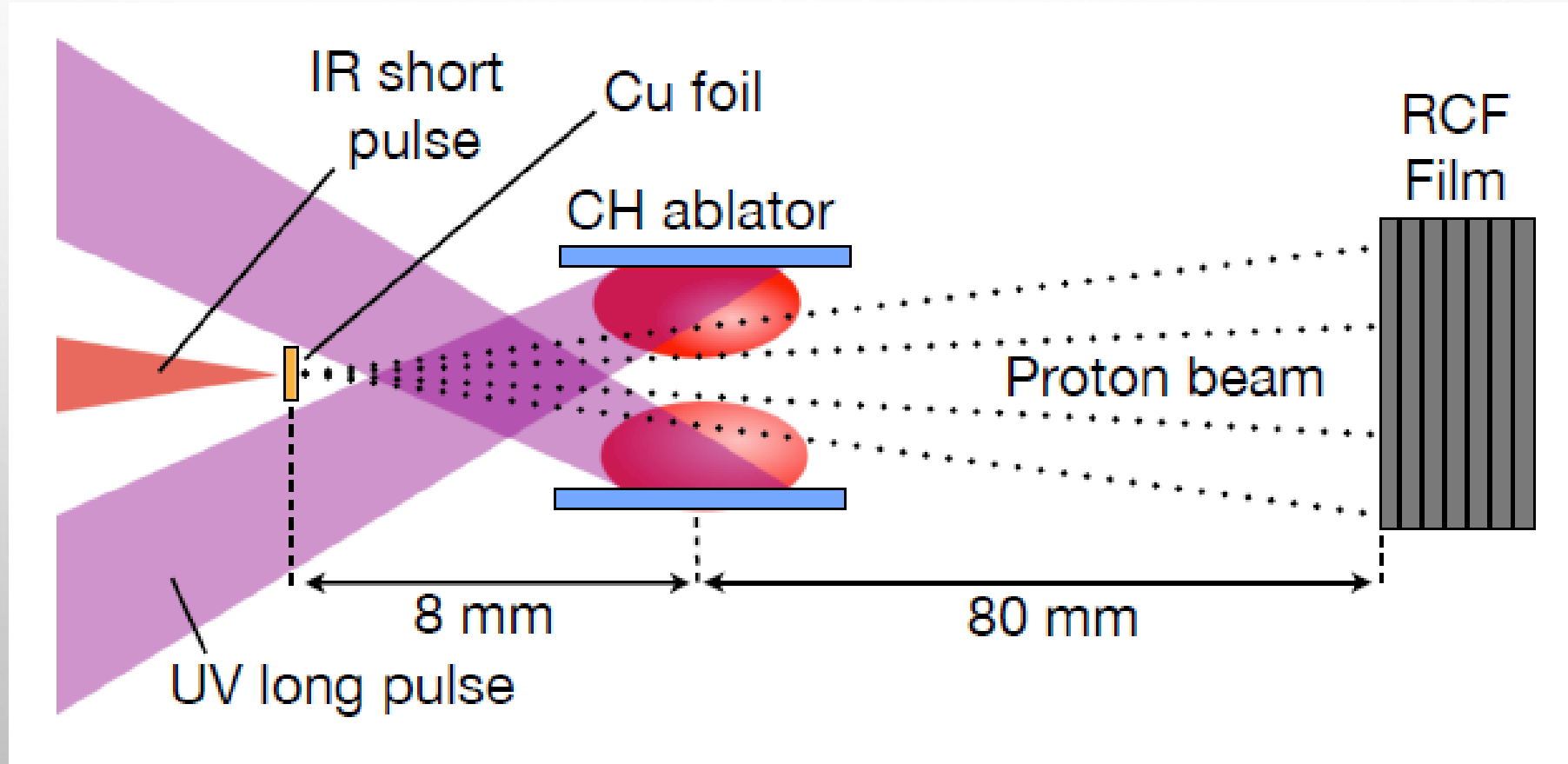
**Densities:** interferometry, X-ray radiography, shadowgraphy

**Temperatures:** spectroscopy, Thomson scattering



# Shock experiments in laboratory

## Laser-produced collisionless shocks

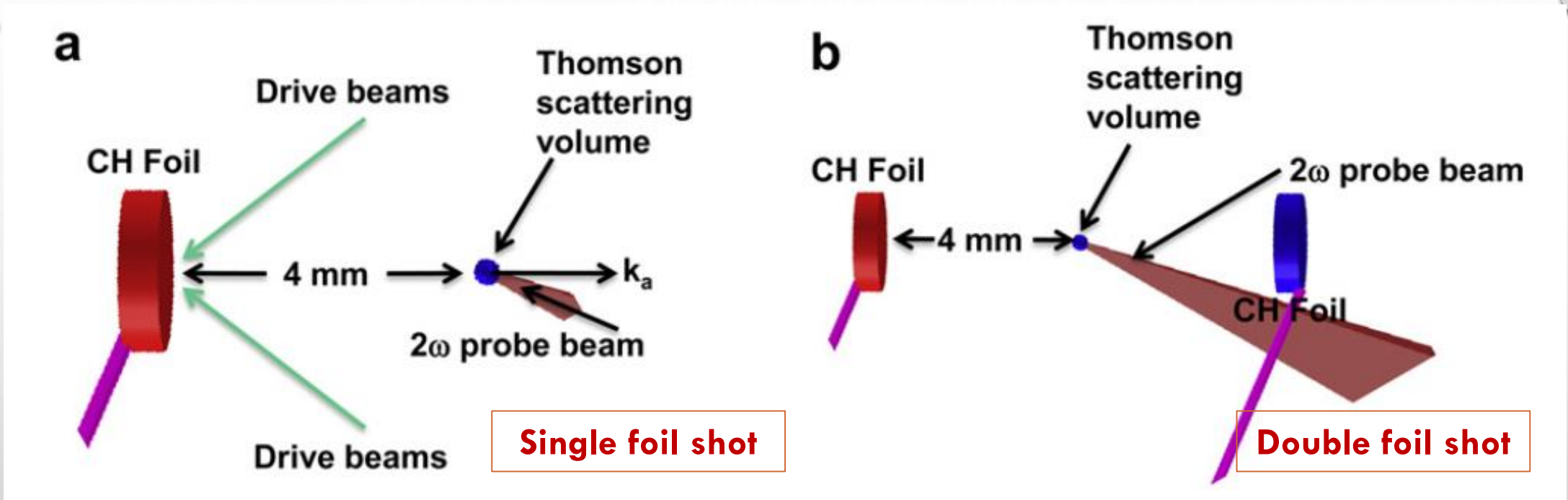


## Shock experiments in laboratory

- Collision mean free path  $\lambda$
- Spatial scale of the interacting flows  $l_{int}$
- Scale needed to generate plasma instabilities  $l^*$

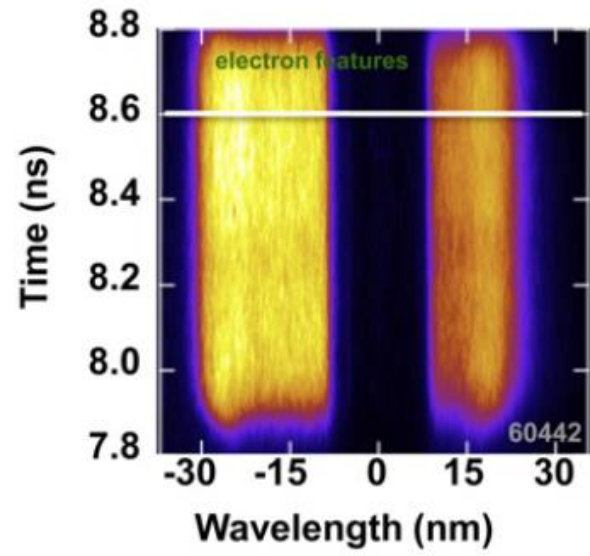
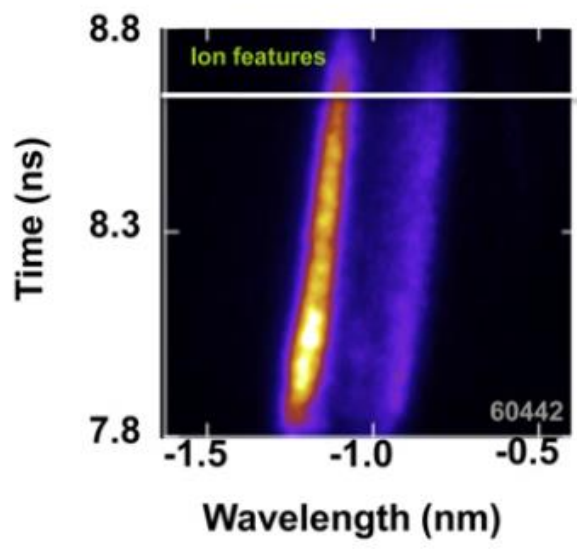
Conditions:  $\lambda \gg l_{int} \gg l^*$

# Shock experiments in laboratory

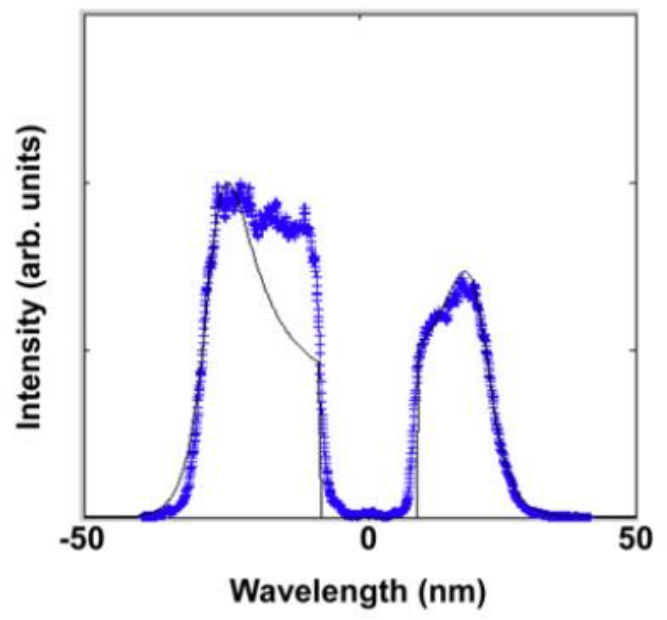
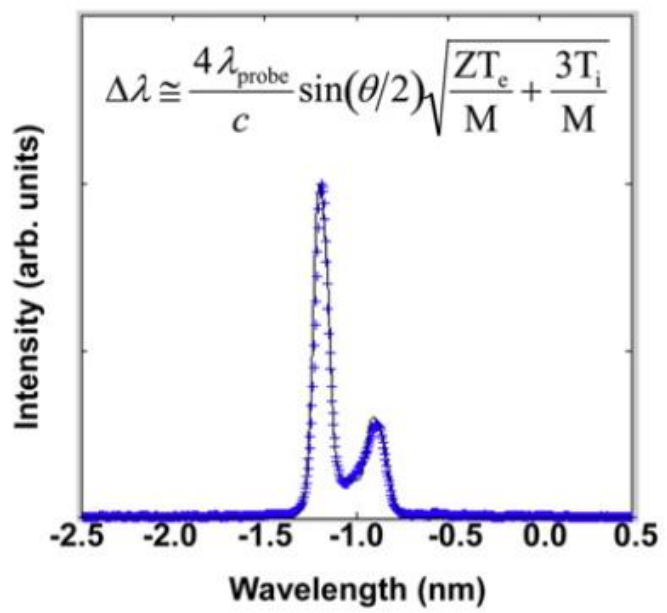


- Heater-drive beams: 5 kJ in a 1 ns square pulse shape
- Probe beam: 0.53 micrometer wavelength
- Thomson scattering is measured from the electron feature and the ion feature

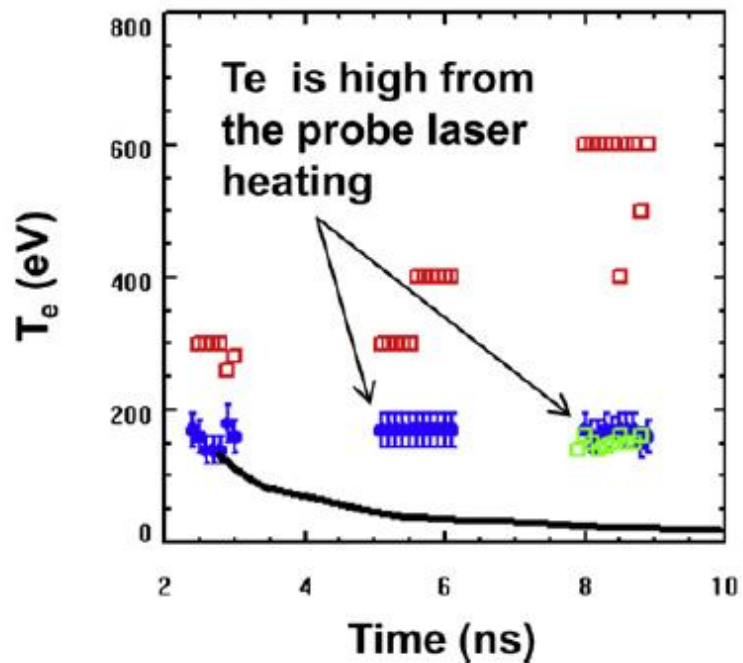
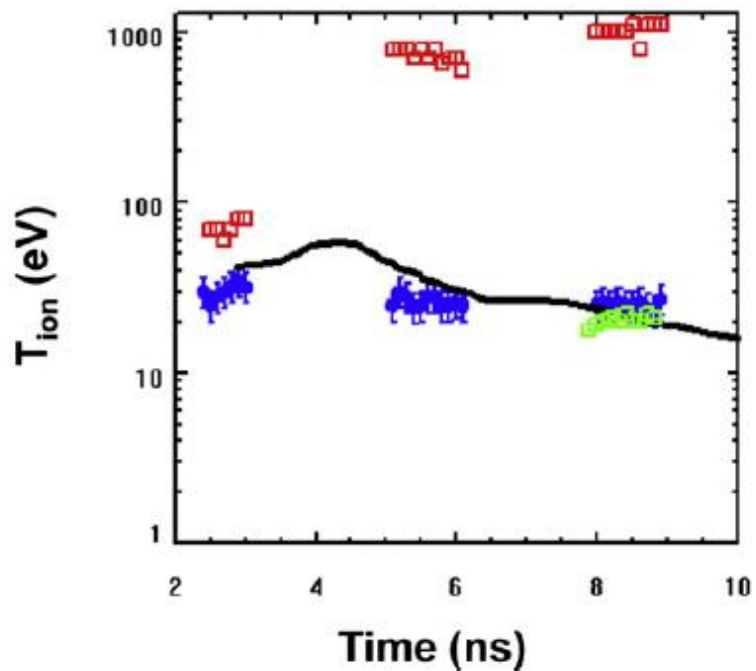
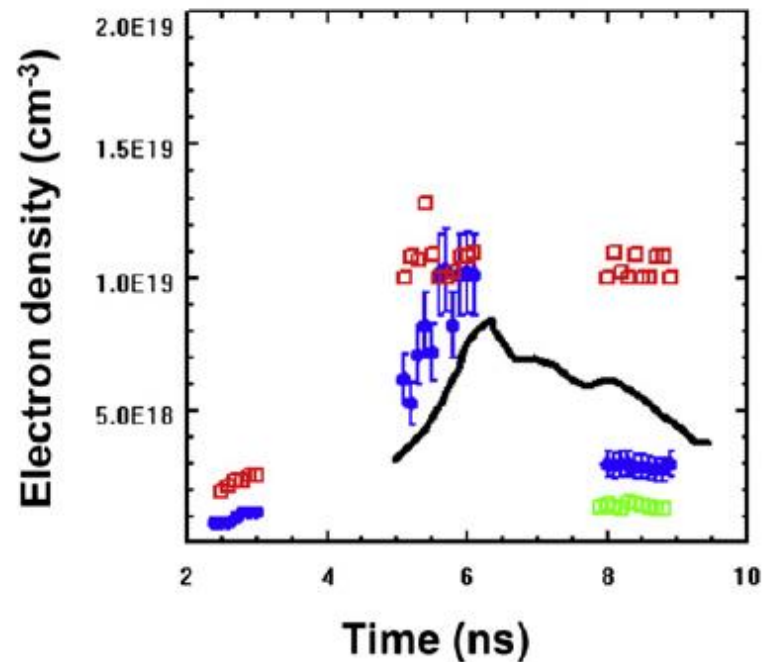
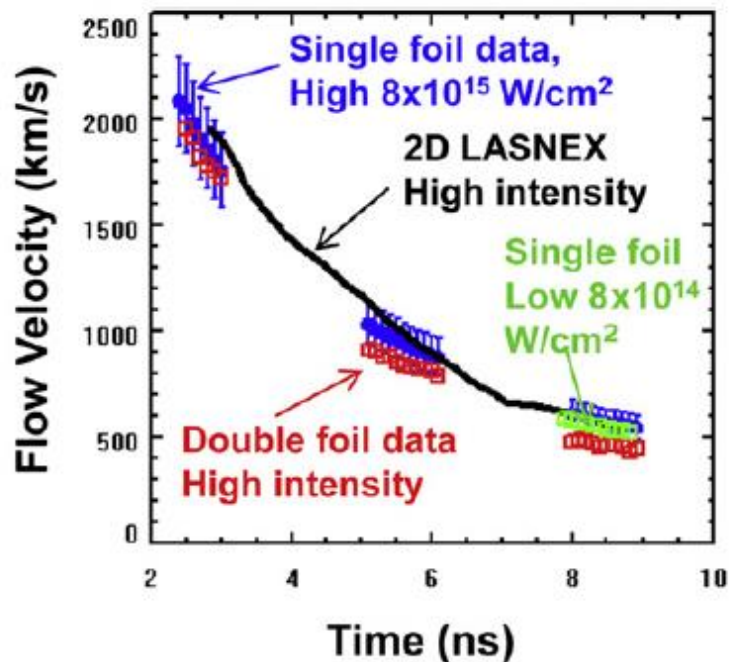
# Single foil shot



$T_e = 170 \text{ eV}$   
 $T_i = 26 \text{ eV}$   
 $n_e = 3.1 \times 10^{18} \text{ cm}^{-3}$   
 $v_{\text{flow}} = 5.7 \times 10^7 \text{ cm/s}$   
 $u_{\text{drift}} = 1.2 \times 10^8 \text{ cm/s}$







**Blue: single foil shots**

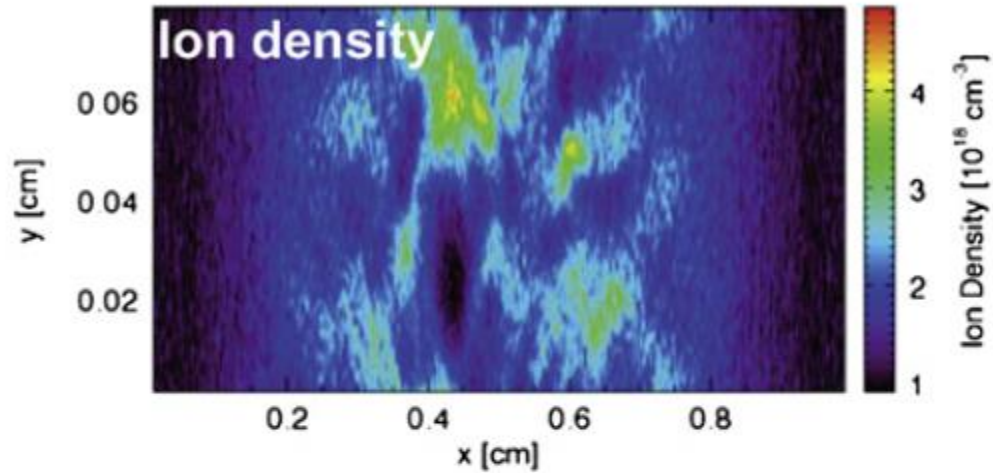
**Red: double foil shots**

**Green: single foil with 10 times lower driven laser intensity**

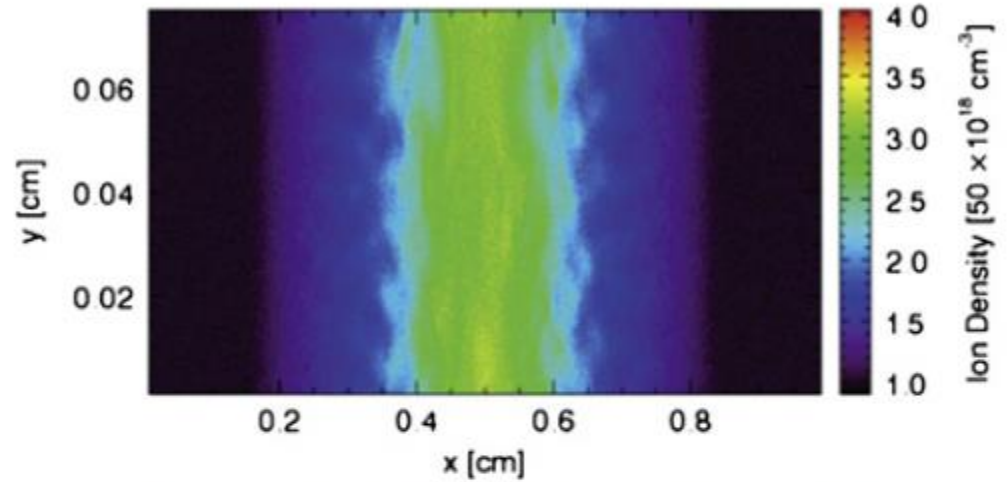
**Solid line: 2D simulations**

# Shock experiments in laboratory

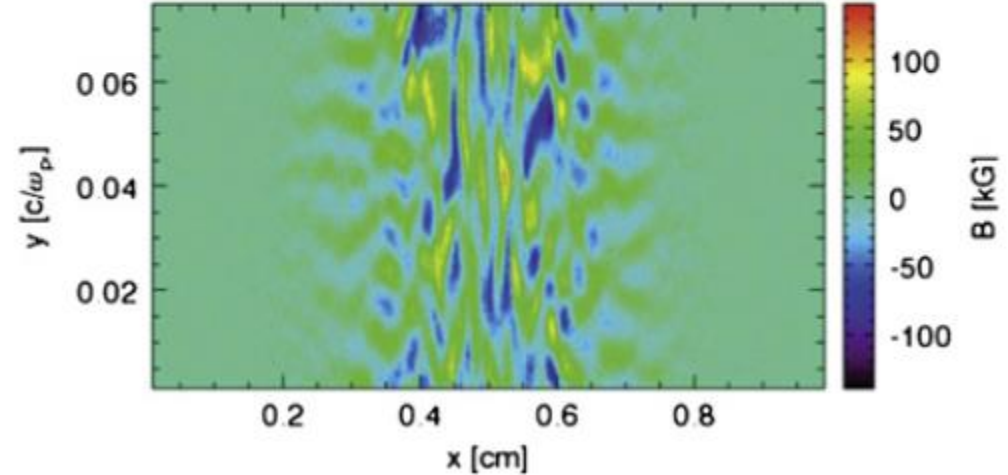
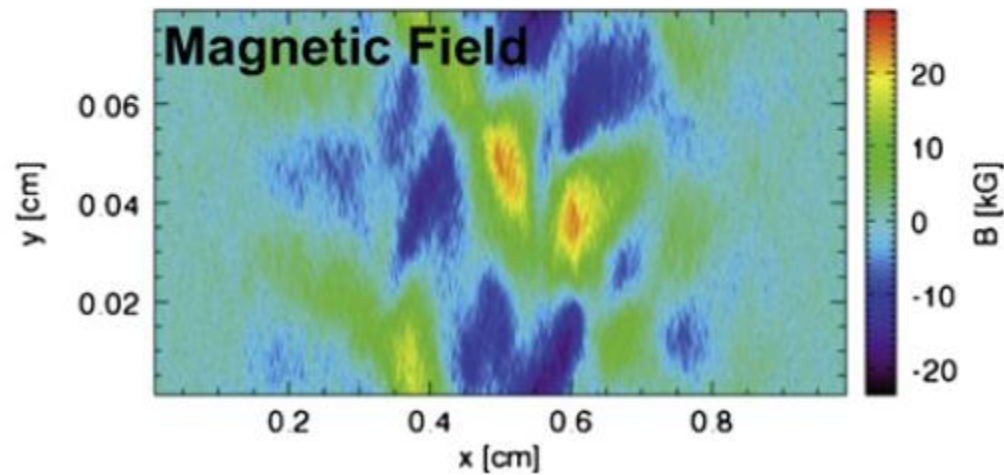
Dec. experimental condition



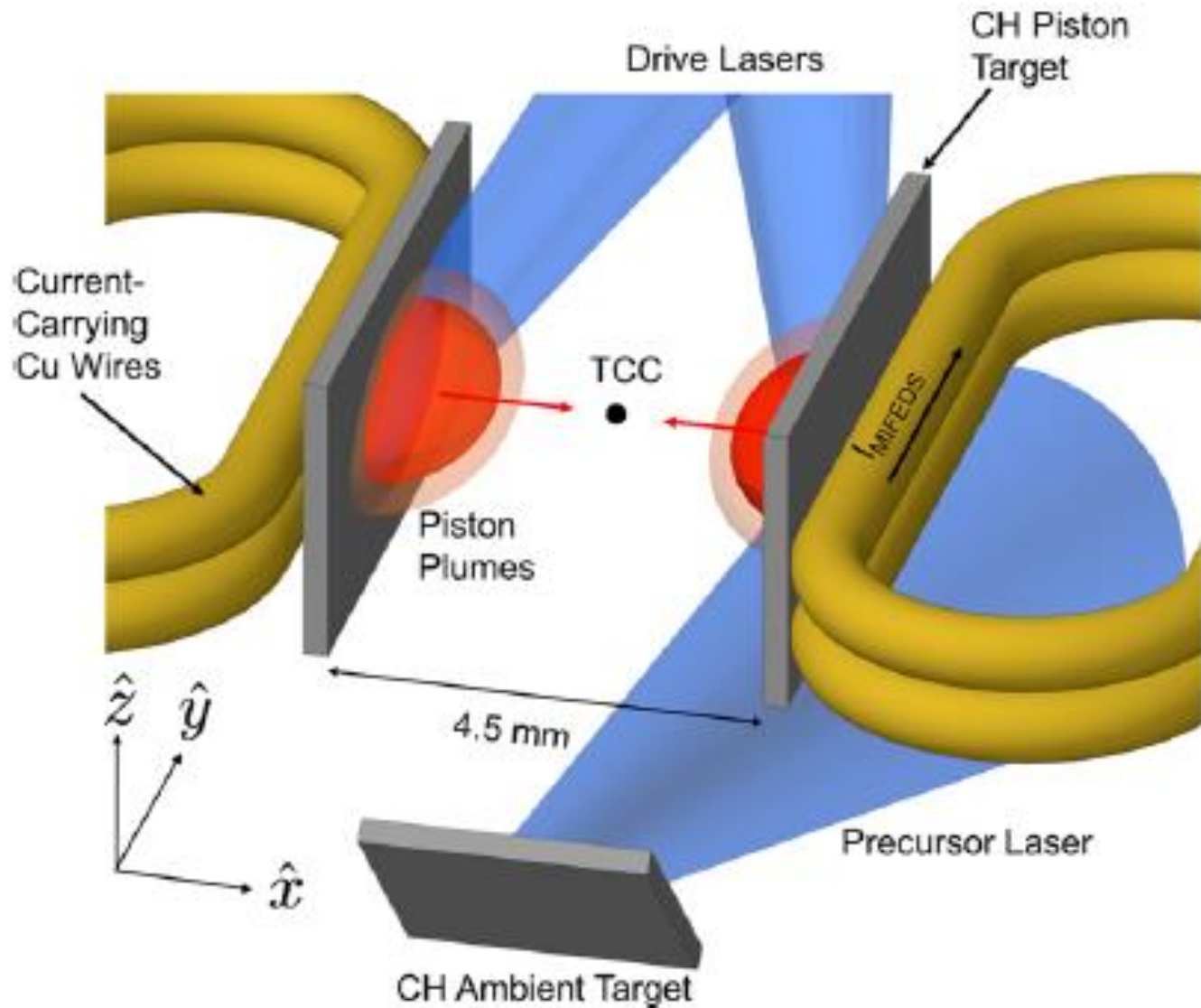
Fully formed shock



**Magnetic Field**



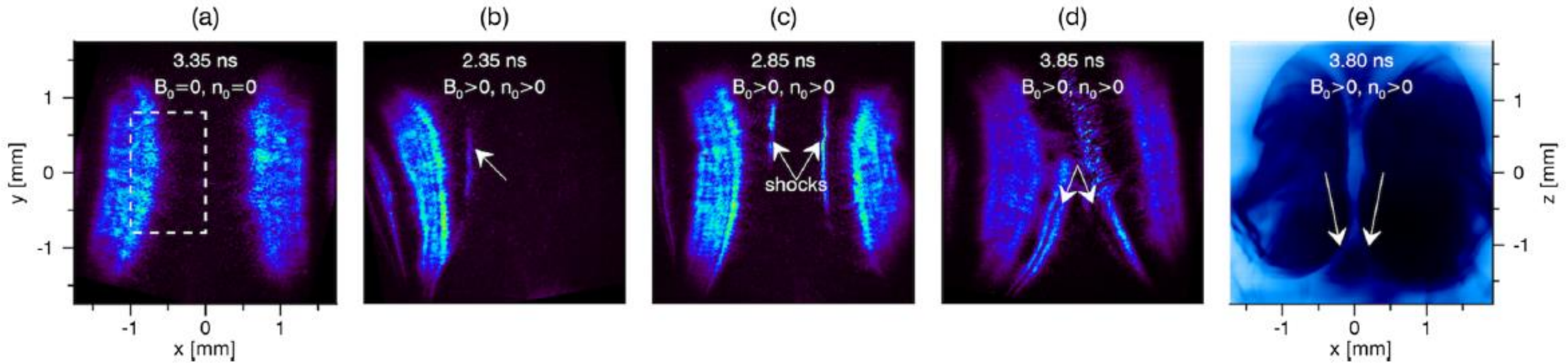
## Shock experiments in laboratory



- Two opposing plastic piston targets irradiated by a  $5 \times 10^{13} \text{ W/cm}^2$  drive beam (351 nm, 1.5 kJ, 2 ns)
- Ambient CH target irradiated by a  $9 \times 10^{12} \text{ W/cm}^2$  beam (351 nm, 100 J, 1 ns)
- An externally imposed background magnetic field  $B_{peak} \sim 8 \text{ T}$



# Shock experiments in laboratory



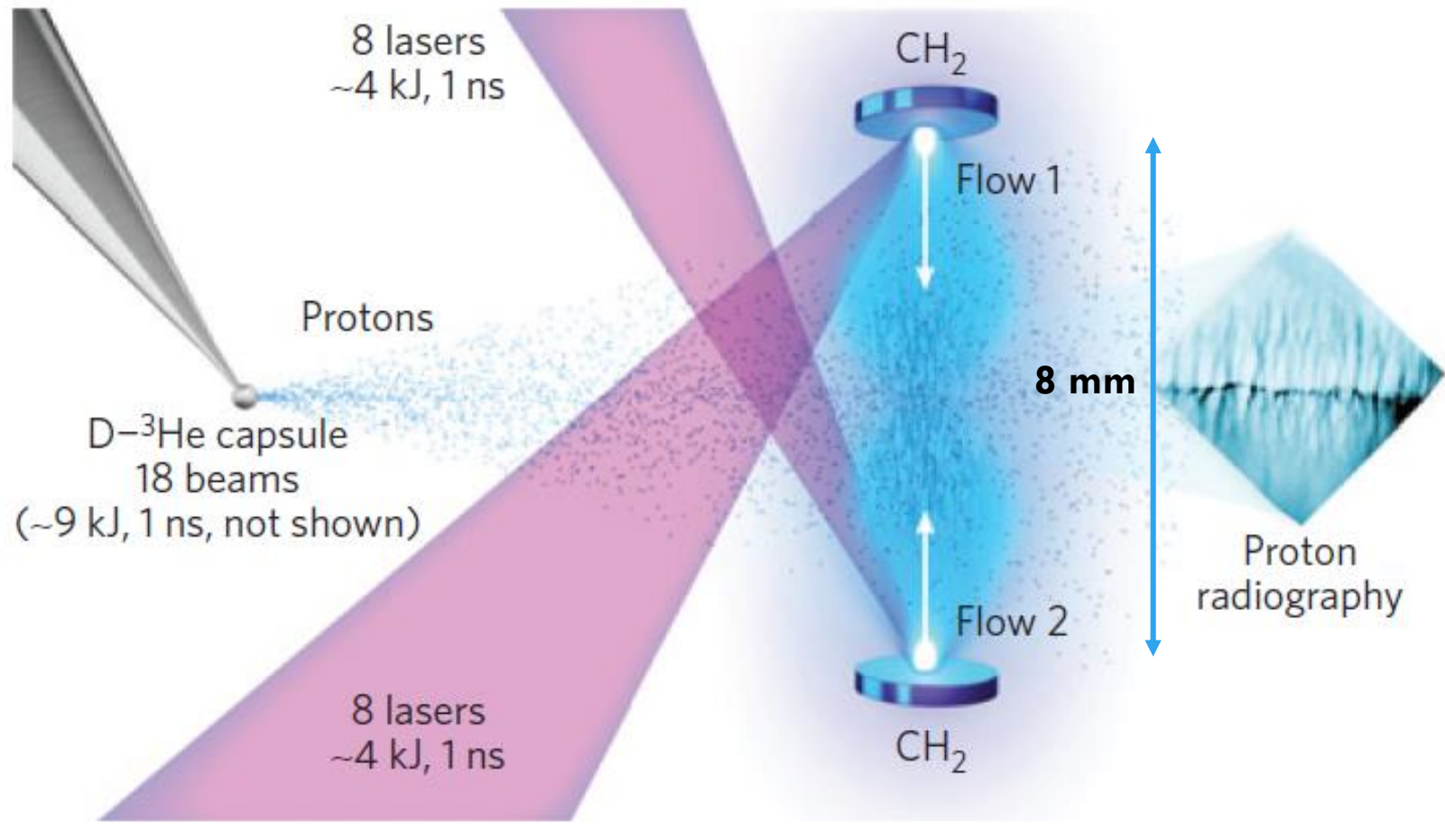
$$n_0 \sim 6 \times 10^{18} \text{ cm}^{-3}$$

Wide bands: piston plasma plums

Narrow bands: shocks (strong density gradients, piston and ambient plasma interact)

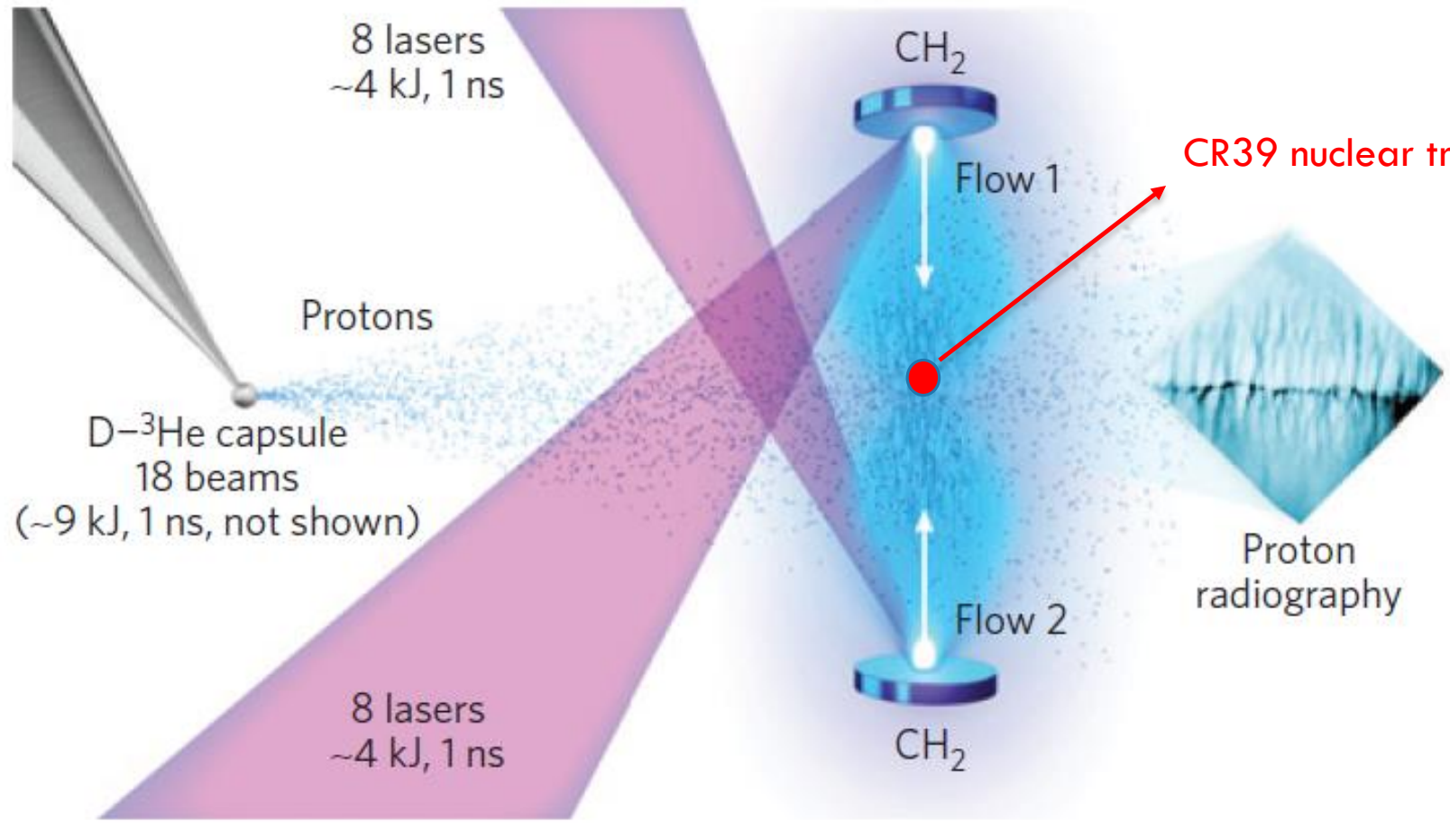


# Shock experiments in laboratory



- 2 foils irradiated simultaneously
- Plasma flows interact near the midplane
- Plasma flow velocity 1000-2000 km/s
- Electron density  $n_e \sim 5 \times 10^{18}$  per cube centimeter
- Electron and ion temperature < 200 eV
- Proton imaging to detect magnetic fields
- Protons at 3MeV (from D-D reactions) and at 14.7MeV (from D-3He reactions) are produced

# Shock experiments in laboratory

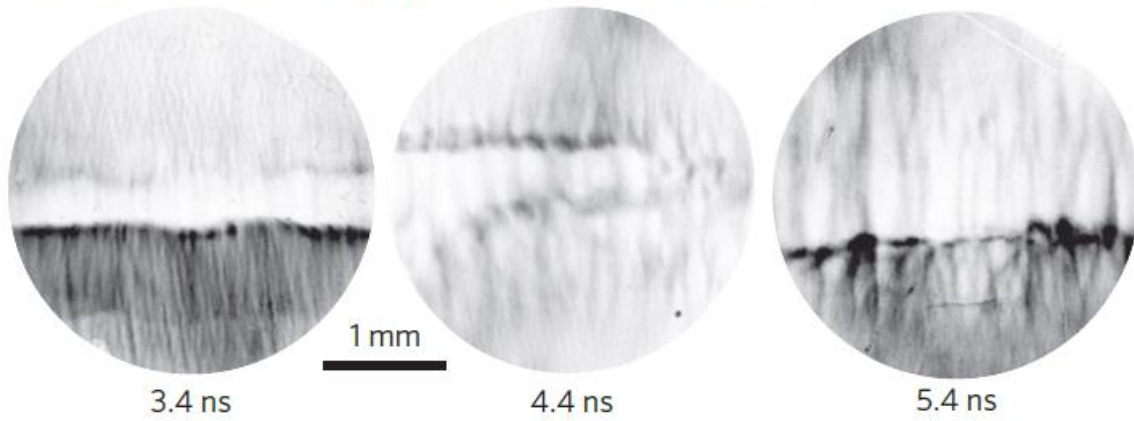


$$v_{flow} \sim 1000 - 2000 \text{ km/s}$$

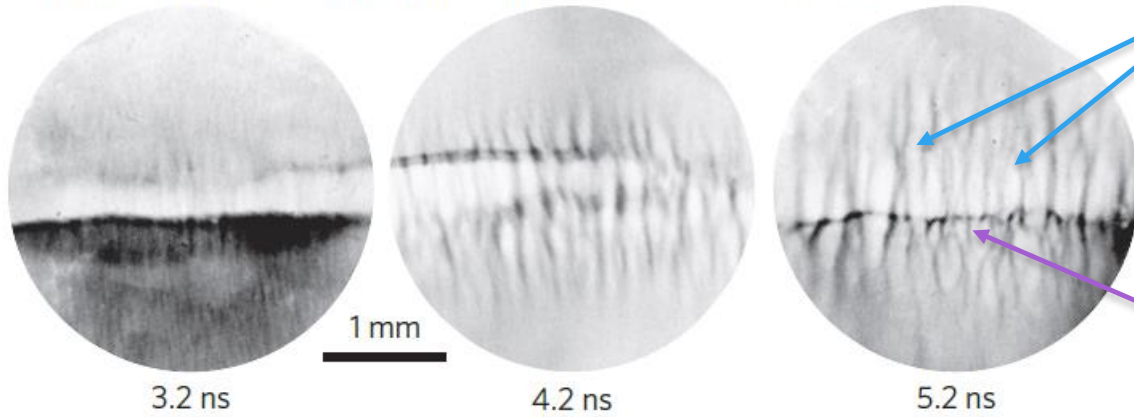
$$n_e \sim 5 \times 10^{18} \text{ cm}^{-3}$$

$$T_e, T_i < 200 \text{ eV}$$

Experimental proton radiographs from 3.0 MeV (D-D) protons



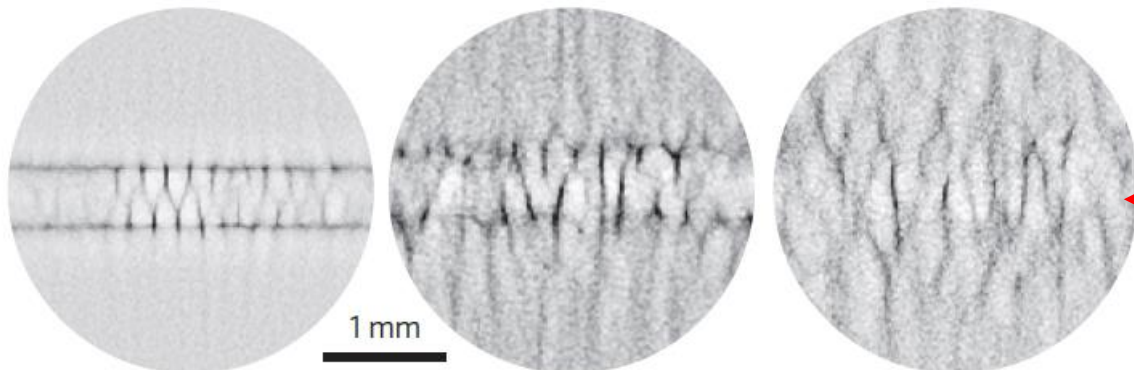
Experimental proton radiographs from 14.7 MeV (D-<sup>3</sup>He) protons



**Magnetic Weibel filaments**

**Extended magnetic "plates"**

Synthetic proton radiographs from 14.7 MeV protons



**Particle-in-cell (PIC) simulations**

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# Summary

- In astronomy shocks are collisionless
- Shock waves act as cosmic ray accelerator
- In laboratory, high power laser can produced collisionless shocks
- Shocks can generate magnetic fields